

AY1 Homework for Quiz 2: Spring 2017

Erg=CGS unit of energy=gram·cm²/second²; E=mc² in units of ergs if “m” is in grams and “c”=3x10¹⁰ cm/sec;

1. Star A has twice the trigonometric parallax and twice the luminosity of Star B (assume no dust toward either star).

a) What is the relative distance of the two stars?

Star A is $\frac{1}{2}$ the distance of Star B

b) What is the relative brightness of the two stars?

If the same luminosity, Star A would be 4x brighter than B based on the distance ratio.
At twice the luminosity of B, Star A will be 4 x 2= 8 times brighter than B

2. The Sun will eventually go through which of the following phases?

planetary nebula

red-giant branch

SNII

white dwarf

3. In the fusion of four protons into helium, 4.7×10^{-26} grams of material is converted into energy. How much energy does this amount of matter produce?

$$E=mc^2=(4.7 \times 10^{-26} \text{ gr}) \times (3 \times 10^{10})^2 = (4.7 \times 9) \times 10^{-6} \text{ ergs}$$

4. Which of the following are True (T), which False (F)?

T The fraction of the Sun composed of helium is larger than it was 1 billion years ago.

T The Sun is losing mass every day

F The fraction of the Sun composed of Fe is larger than it was 1 billion years ago

F The luminosity of the Sun decreases a small amount every day as it uses up its hydrogen fuel

5. Given that hydrogen fusion to helium produces 10^{18} ergs of energy per gram of hydrogen:

- A. How much energy can the Sun produce via this mechanism with the 2×10^{32} grams of hydrogen in the core where the temperature is hot enough for fusion to occur?

$$E = 10^{18} \text{ ergs/gr} \times 2 \times 10^{32} \text{ gr} = 2 \times 10^{50} \text{ ergs}$$

- B. How long could the Sun produce energy via this hydrogen fusion at its luminosity of 4×10^{33} ergs/second?

$$\text{Time} = (2 \times 10^{50} \text{ ergs}) / (4 \times 10^{33} \text{ ergs/second}) = 0.5 \times 10^{17} \text{ seconds}$$

6. In a SN I outburst, the initial burst of light is due to the energy released in fusion reactions. What keeps the SN glowing after the first 15 days? (select one)

- A. Neutrino heating
 B. Photo-disintegration of Iron nuclei
 C. Radioactive decay of Nickel and Cobalt formed during the explosion
 D. Radioactive decay of Hydrogen and Helium

7. Which of the following are used in measuring stellar masses (check any that are true)?

- A. Proper motion measurements of nearby stars
 B. Radial velocity measurements of stars in binary systems
 C. Red Giants that are within 100pc of the Sun
 D. Newton's Laws of gravity

8. The Solar luminosity at the Earth is 3.9×10^{33} ergs/sec. What is it at the distance of Jupiter (5 AU or 5 times the distance from the Sun compared to Earth)?

- A. 3.9×10^{33} ergs/sec
 B. $(3.9 \times 10^{33})/5$ ergs/sec
 C. $(3.9 \times 10^{33})/5^2$ ergs/sec
 D. $(3.9 \times 10^{33}) \times 5^2$ ergs/sec

9. "Hydrostatic" models for the Sun or other stars are based on (check any that are correct):

- A. Gas pressure compressing stars to the point just before they become liquid
- B. Balancing the force of gravity and gas (thermal) pressure at every radius
- C. The laws of physics governing the fusion of the elements
- D. Static electricity providing support against gravitational collapse

10. The principle behind determining stellar radius or surface area is best described by (select one):

- A. Using stellar luminosity combined with Wien's Law for the peak radiation as a function of temperature
- B. Using the distance determined from parallax and the apparent size of a star
- C. Using stellar luminosity, stellar surface temperature and Stephan's Law that related temperature and radiated energy for a surface
- D. Using binary stars and radial velocity measurements

11. Which of the following statements are True (T), which (F) regarding the Main Sequence in the Hertzsprung-Russell Diagram:

- A. It is a mass sequence with the lower-mass stars at the low-temperature/low-luminosity corner and the higher-mass stars at the high-T/high-L corner **TRUE**
- B. It is the sequence of stars that are in equilibrium fusing hydrogen to helium in their cores **TRUE**
- C. Stars start their lives in the cool-T/low-L corner and evolve along the Main Sequence as they age and get hotter **FALSE**
- D. Once a star is on the Main Sequence, it stays there for at least 10 billion years **FALSE**

12. High temperature is required for hydrogen fusion because (check one):

- A. Only at high temperatures do the protons approach close enough for the nuclear force to overcome the electrical force
- B. Only at high temperatures can gravity be balanced
- C. Hydrogen can not be fused unless oxygen is present and that only occurs at "burning" temperatures

13. Why does hydrogen fusion only occur at the center of the Sun?

The required temperature of at least 10 million K only occurs in the central region

14. How much energy is released by the reactions in the core of the Sun each second? 3.9×10^{33} erg

15. For two stars on the Main Sequence of the H-R Diagram, compare the luminosity of Star F and Star G if Star G has a higher surface temperature and is at twice the distance of Star F.

Distance is not relevant. On the main-sequence, the hotter star is more luminous

16. Four stars occupy the four corners of the H-R diagram (UL, LL, UR, LR):

- _____ A. In which corner(s) is (are) the largest star(s)? UR
- _____ B. In which corner(s) is (are) the most luminous star(s) (UL, UR)
- _____ C. In which corner(s) is (are) the hottest star(s) UL, LL
- _____ D. In which corner(s) is (are) the lowest-mass star(s) LR

17. Which of the following are true (T), which false (F) about the white dwarf the Sun will eventually become?

- ___ A. It will have become slightly more massive than the Sun is now because lightweight hydrogen has been converted into heavier Helium FALSE
- ___ B. It will be enriched in helium compared to the Sun TRUE
- ___ C. It will initially have a much higher surface temperature than the Sun TRUE
- ___ D. It will be supported against gravity by electron degeneracy forces TRUE

18. Which of the following support the theory of SN II: core-collapse supernovae?

- ___ A. SN II are always seen near regions of star formation
- ___ B. The supernova remnants in the Galaxy show evidence of heavy element enhancements
- ___ C. There are pulsars (rotating neutron stars) at the centers of some SN II remnants
- ___ D. They have luminosities similar to red giant stars

19. Which of the following are predicted properties of the neutron star left behind by a SNII?

- ___ A. They will initially have very high surface temperatures
- ___ B. They will be spinning very rapidly
- ___ C. They will have much higher densities than main sequence stars or white dwarfs
- ___ D. They will be supported by gravity by neutron degeneracy

20. Which of the following are part of the story that explains a SN I explosion?

- A. Core collapse of the iron core of a massive star
- B. Mass transfer in a close binary onto a white dwarf
- C. Violent stellar collisions in star clusters that force objects over the Chandrasekar Limit
- D. Explosive nucleosynthesis that produces radioactive nickel and cobalt

21. Why do we expect a freshly-formed neutron star to be rapidly rotating.

Conservation of angular momentum and the large ratio of the radius of the iron core that collapsed to the radius of the neutron star

22. The lowest-luminosity white dwarfs in the Galaxy have about 10^{-5} the luminosity of the Sun. Why are there no whites dwarfs at lower luminosities than this?

The Galaxy has a finite age

23. How long will a star with 10 times the mass of the Sun and 10,000 times the luminosity of the Sun spend on the Main Sequence of the H-R Diagram?

Relative age is proportional to the mass and inversely proportional to the luminosity so it will last $10/10000 = 1/1000$ the main-sequence lifetime of the Sun. This is

$$10^{-3} \times 10^{10} = 10^7 \text{ year}$$

24. Why don't White Dwarfs collapse to smaller radius due to gravity? (check any that are correct).

- A. They are supported by hydrogen fusion
- B. All white dwarfs are in the process of slowly collapsing till they become neutron stars
- C. They are supported by electron degeneracy
- D. They are already a super density fluid and fluids can not be compressed