## AY2 Homework for Quiz 2: Fall 2019

Erg=CGS unit of energy=gram·cm $^2$ /second $^2$ ; E=mc $^2$  in units of ergs if "m" is in grams and "c"=3x10 $^{10}$  cm/sec;

- 1. Star A has twice the trigonometric parallax and twice the luminosity of Star B (assume no dust toward either star).
  - a) What is the relative distance of the two stars?

Star A is ½ the distance of Star B

b) What is the relative brightness of the two stars?

If the same luminosity, Star A would be 4x brighter than B based on the distance ratio. At twice the luminosity of B, Star A will be 4x 2= 8 times brighter than B

- 2. The Sun will eventually go through which of the following phases?
- \_x\_ planetary nebula
- $_{\bf x}$ \_ red-giant branch
- \_\_\_ SNII
- x white dwarf
- 3. In the fusion of four protons into helium,  $4.7 \times 10^{-26}$  grams of material is converted into energy. How much energy does this amount of matter produce?

 $E=mc^2=(4.7\times10^{-26} \text{ gr}) \times (3\times10^{10})^2 = (4.7\times9) \times 10^{-6} \text{ ergs}$ 

- 4. Which of the following are True (T), which False (F)?
- **\_T**\_ The fraction of the Sun composed of helium is larger than it was 1 billion years ago.
- \_T\_ The Sun is losing mass every day
- \_F\_ The fraction of the Sun composed of Fe is larger than it was 1 billion years ago
- \_F\_\_ The luminosity of the Sun decreases a small amount every day as it uses up its hydrogen fuel

- 5. Given that hydrogen fusion to helium produces 10<sup>18</sup> ergs of energy per gram of hydrogen:
  - A. How much energy can the Sun produce via this mechanism with the  $2 \times 10^{32}$  grams of hydrogen in the core where the temperature is hot enough for fusion to occur?

 $E=10^{18} \text{ ergs/gr} \times 2 \times 10^{32} \text{ gr} = 2 \times 10^{50} \text{ ergs}$ 

B. How long could the Sun produce energy via this hydrogen fusion at its luminosity of  $4 \times 10^{33}$  ergs/second?

Time=  $(2 \times 10^{50} \text{ ergs})/(4 \times 10^{33} \text{ ergs/second})=0.5 \times 10^{17} \text{ seconds}$ 

- 6. In a SN I outburst, the initial burst of light is due to the energy released in fusion reactions. What keeps the SN glowing after the first 15 days? (select one)
  - \_\_ A. Neutrino heating
  - \_\_\_ B. Photo-disintegration of Iron nuclei
  - \_\_\_ C. Radioactive decay or Nickel and Cobalt formed during the explosion
  - \_\_\_ D. Radioactive decay of Hydrogen and Helium
- 7. Which of the following are used in measuring stellar masses (check any that are true)?
  - \_ A. Proper motion measurements of nearby stars
  - \_\_ B. Radial velocity measurements of stars in binary systems
  - \_\_\_ C. Red Giants that are within 100pc of the Sun
  - \_\_ D. Newton's Laws of gravity
- 8. The Solar luminosity at the Earth is  $3.9 \times 10^{33}$  ergs/sec. What is it at the distance of Jupiter (5 AU or 5 times the distance from the Sun compared to Earth)?
  - \_\_\_ A. 3.9 x 10<sup>33</sup> ergs/sec
  - \_\_\_ B.  $(3.9 \times 10^{33})/5 \text{ ergs/sec}$
  - \_\_\_ C.  $(3.9 \times 10^{33})/5^2 \text{ ergs/sec}$
  - D.  $(3.9 \times 10^{33}) \times 5^2 \text{ ergs/sec}$

9.	"Hydrostatic" models for the Sun or other stars are based on (check any that are correct):		
	-	Gas pressure compressing stars to the point just before they become liquid	
	A. B.	Balancing the force of gravity and gas (thermal) pressure at every radius	
	C.	The laws of physics governing the fusion of the elements	
	D.	Static electricity providing support against gravitational collapse	
10	_	ciple behind determining stellar radius or surface area is best d by (select one):	
	A.	Using stellar luminosity combined with Wien's Law for the peak radiation	
		inction of temperature	
	В.	•	
		Using stellar luminosity, stellar surface temperature and Stephan's Law that	
		d temperature and radiated energy for a surface	
	D.	Using binary stars and radial velocity measurements	
11		f the following statements are True (T), which (F) regarding the Main e in the Hertzsprung-Russell Diagram:	
	A.		
	tempe	erature/low-luminosity corner and the higher-mass stars at the high-T/high-ner TRUE	
	B.	It is the sequence of stars that are in equilibrium fusing hydrogen to helium ir cores TRUE	
	C.	Stars start their lives in the cool-T/low-L corner and evolve along the Main ence as they age and get hotter FALSE	
	D.	Once a star is on the Main Sequence, it stays there for at least 10 billion FALSE	
12	.High tem	perature is required for hydrogen fusion because (check one):	
	A.	Only at high temperatures do the protons approach close enough for the	
	<mark>nuclea</mark>	ar force to overcome the electrical force	
	B.	Only at high temperatures can gravity be balanced	
	C.	Hydrogen can not be fused unless oxygen is present and that only occurs at	
	"burn	ing" temperatures	
13	.Why doe	s hydrogen fusion only occur at the center of the Sun?	
Th	e required	temperature of at least 10 million K only occurs in the central region	

14. How much energy is released by the reactions in the core of the Sun each second?  $3.9 \times 10^{33}\,erg$ 

15. For two stars on the Main Sequence of the H-R Diagram, compare the luminosity of Star F and Star G if Star G has a higher surface temperature and is at twice the distance of Star F.

Distance is not relevant. On the main-sequence, the hotter star is more luminous

16. Four star	rs occupy the four corners of the H-R diagram (UL, LL, UR, LR):
A.	In which corner(s) is (are) the largest star(s)? UR
	In which corner(s) is (are) the most luminous star(s) (UL, UR)
	In which corner(s) is (are) the hottest star(s) UL, LL
D.	In which corner(s) is (are) the lowest-mass star(s) LR
	the following are true (T), which false (F) about the white dwarf the eventually become?
ligł	It will have become slightly more massive than the Sun is now because htweight hydrogen has been converted into heavier Helium FALSE It will be enriched in helium compared to the Sun TRUE
C.	It will initially have a much higher surface temperature than the Sun TRUE It will be supported against gravity by electron degeneracy forces TRUE
18.Which of	the following support the theory of SN II: core-collapse supernovae?
A.	SN II are always seen near regions of star formation
	The supernova remnants in the Galaxy show evidence of heavy element
	cements There are pulsars (rotating neutron stars) at the centers of some SN II
remna	<u> </u>
	They have luminosities similar to red giant stars
19.Which of by a SNII?	the following are predicted properties of the neutron star left behind
	They will initially have very high surface temperatures They will be spinning very rapidly
C.	They will have much higher densities that main sequence stars or white rarfs
	They will be supported by gravity by neutron degeneracy

20. Which of the following are part of the story that explains a SN I explosion?			
<ul> <li>A. Core collapse of the iron core of a massive star</li> <li>B. Mass transfer in a close binary onto a white dwarf</li> <li> C. Violent stellar collisions in star clusters that force objects over the Chandrasekar Limit</li> <li> D. Explosive nucleosynthesis that produces radioactive nickel and cobalt</li> </ul>			
21. Why do we expect a freshly-formed neutron star to be rapidly rotating.			
Conservation of angular momentum and the large ratio of the radius of the iron core that collapsed to the radius of the neutron star			
22. The lowest-luminosity white dwarfs in the Galaxy have about $10^{-5}$ the luminosity of the Sun. Why are there no whites dwarfs at lower luminosities than this?			
The Galaxy has a finite age			
23. How long will a star with 10 times the mass of the Sun and 10,000 times the luminosity of the Sun spend on the Main Sequence of the H-R Diagram?			
Relative age is proportional to the mass and inversely proportional to the luminosity so it will last $10/10000 = 1/1000$ the main-sequence lifetime of the Sun. This is			
10 <sup>-3</sup> x 10 <sup>10</sup> =10 <sup>7</sup> year			
<ul> <li>24.Why don't White Dwarfs collapse to smaller radius due to gravity? (check any that are correct).  A. They are supported by hydrogen fusion B. All white dwarfs are in the process of slowly collapsing till they become neutron stars C. They are supported by electron degeneracy D. They are already a super density fluid and fluids can not be compressed</li> </ul>			