

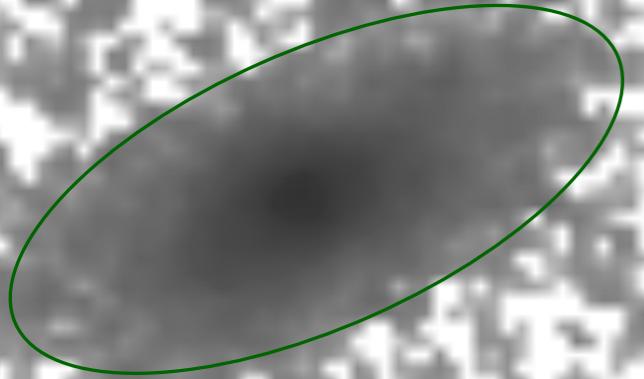
# **CANDELS-VAPC: HST Multi-band Multi-Aperture Photometry Catalogs in the Five CANDELS Fields**

Fengshan Liu, Dongfei Jiang ,  
*(Shenyang Normal University)*

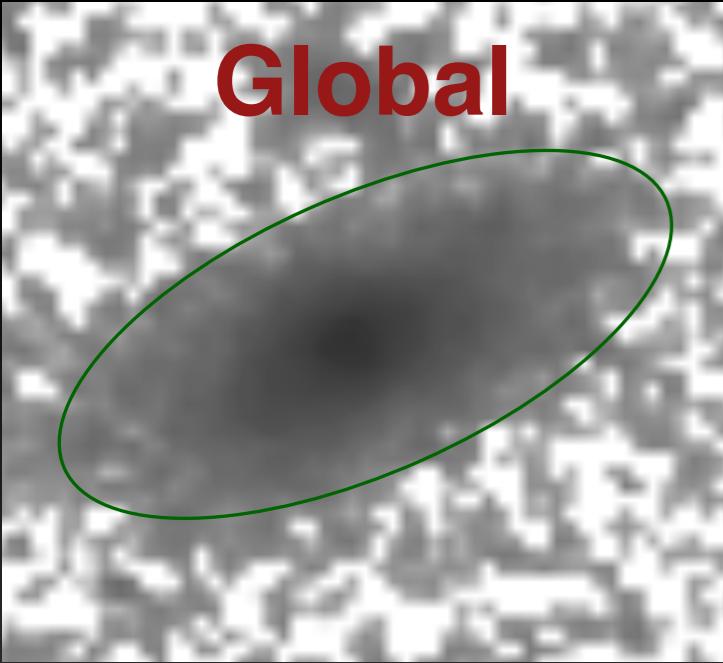
David Koo, Sandy Faber ,Yicheng Guo  
*(University of Californian, Santa Cruz)*  
and CANDELS Team

2017 CANDELS Meeting  
5 August, 2017 @ Santa Cruz

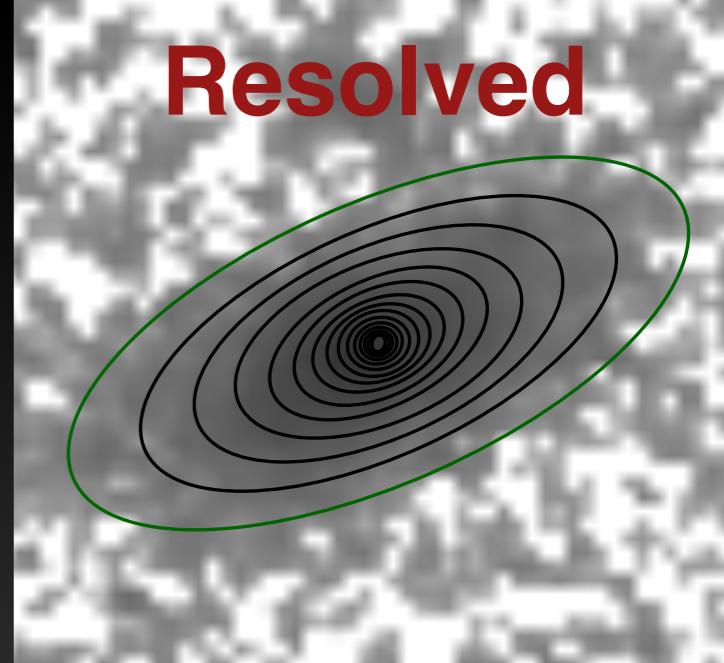
**Global**



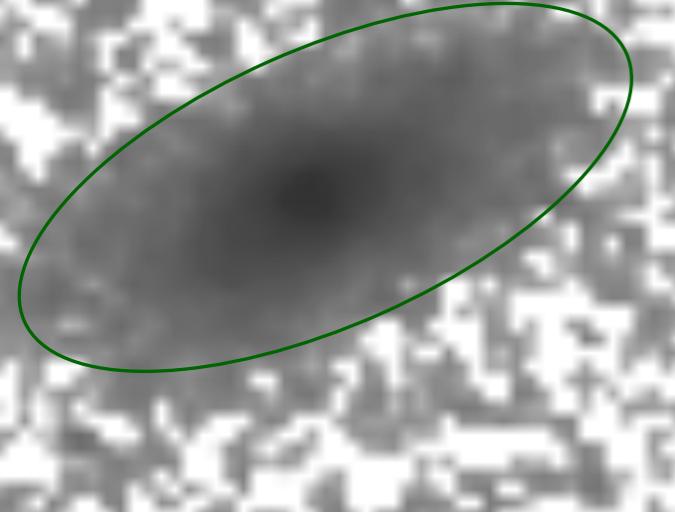
**Global**



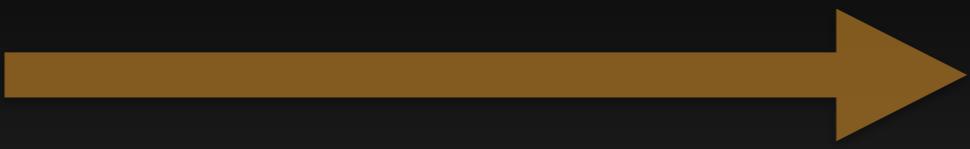
**Resolved**



**Global**

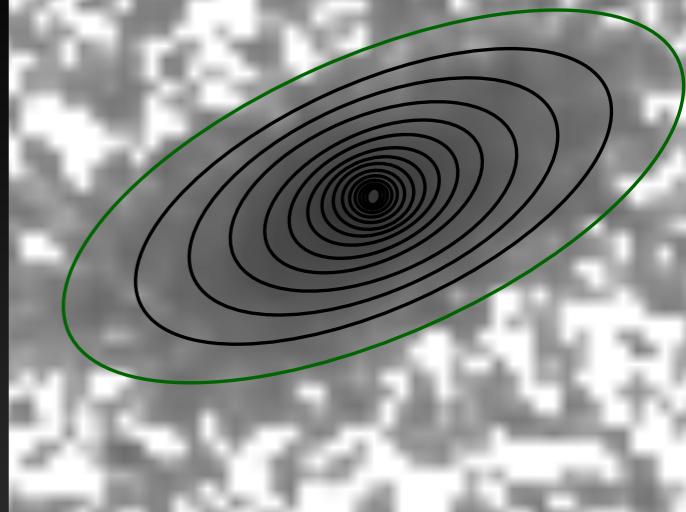


**CANDELS-VAPC: HST Multi-Aperture  
Photometry Catalogs**

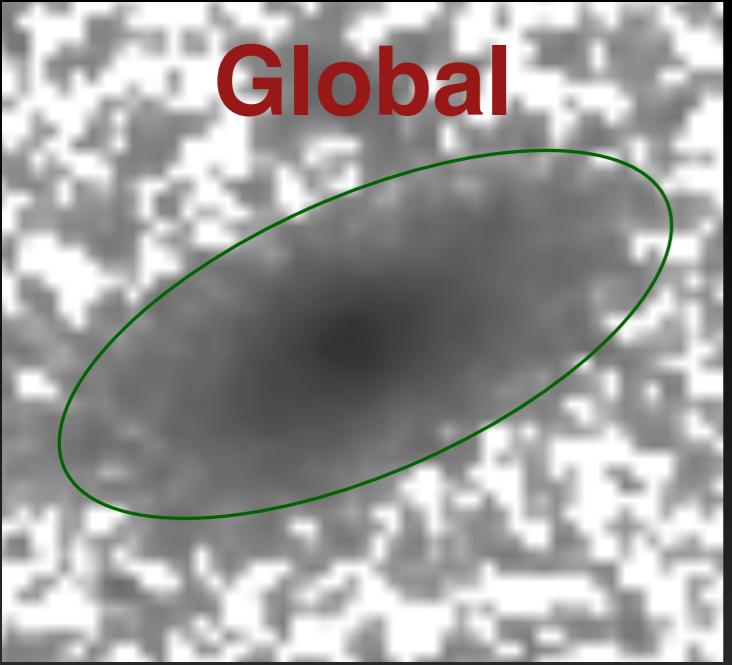


*(Fengshan Liu, Dongfei Jiang,  
David C. Koo, S. M. Faber,  
Yicheng Guo, and CANDELS Team)*

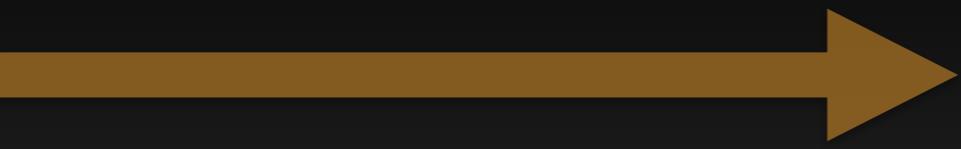
**Resolved**



**Global**

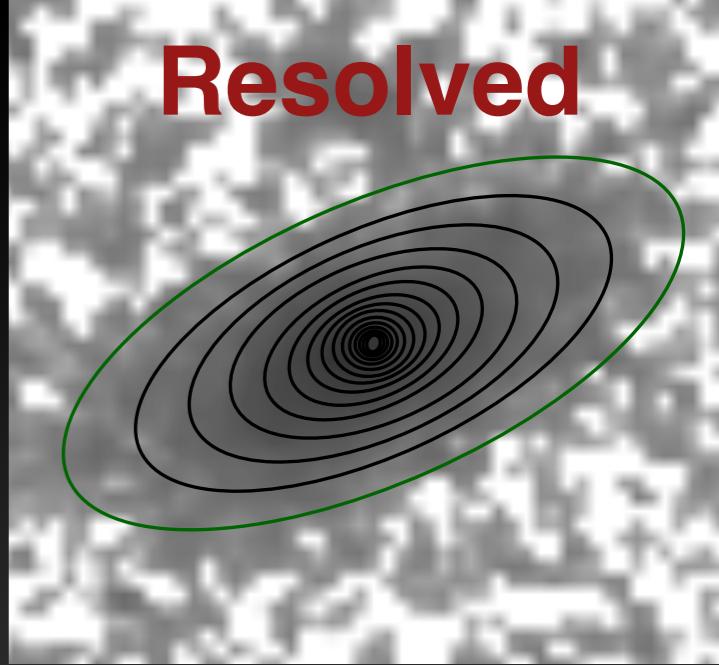


**CANDELS-VAPC: HST Multi-Aperture  
Photometry Catalogs**



*(Fengshan Liu, Dongfei Jiang,  
David C. Koo, S. M. Faber,  
Yicheng Guo, and CANDELS Team)*

**Resolved**

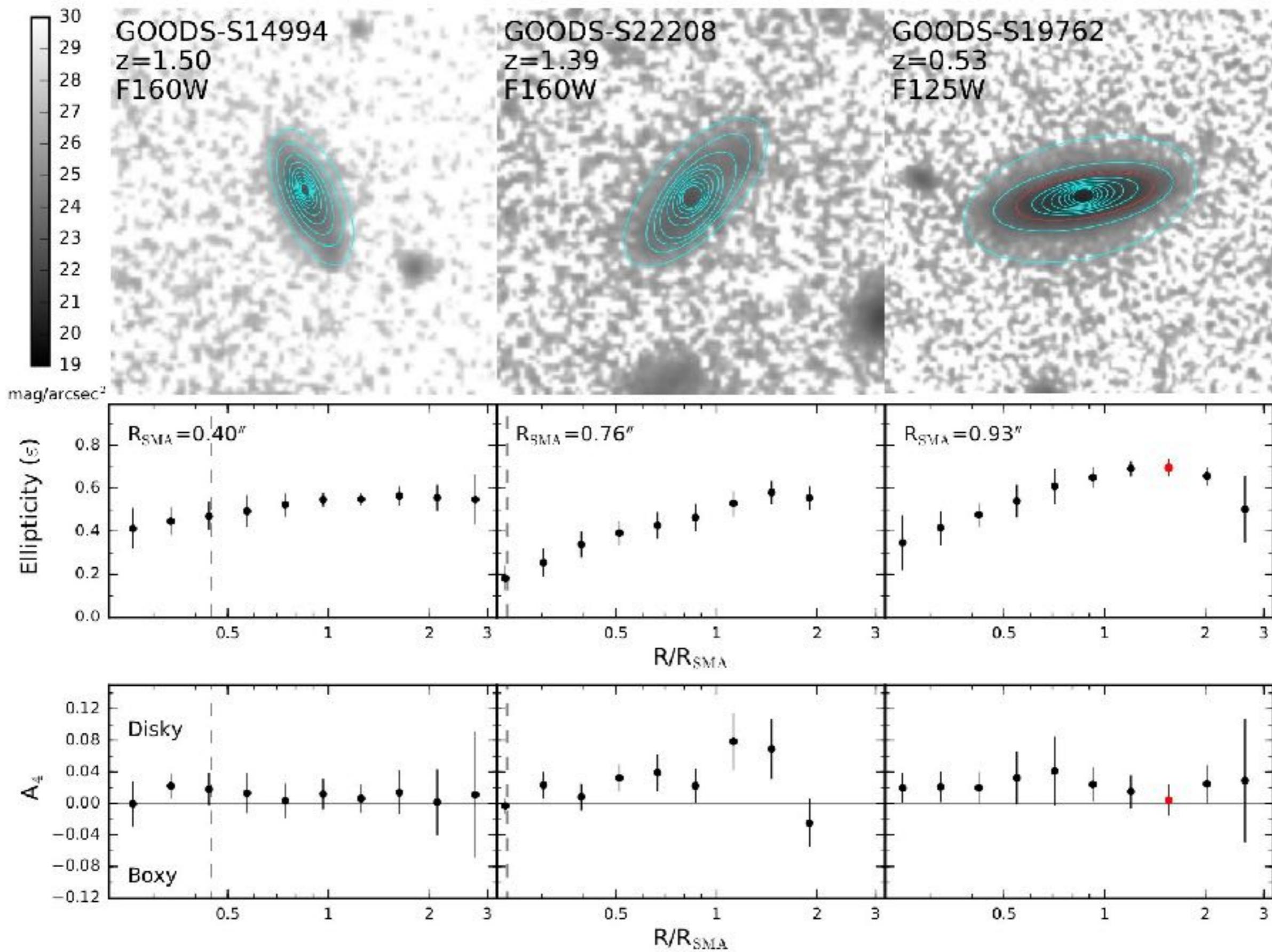


**Part I : Isophotal Aperture Photometry (Fourier Expansion)**

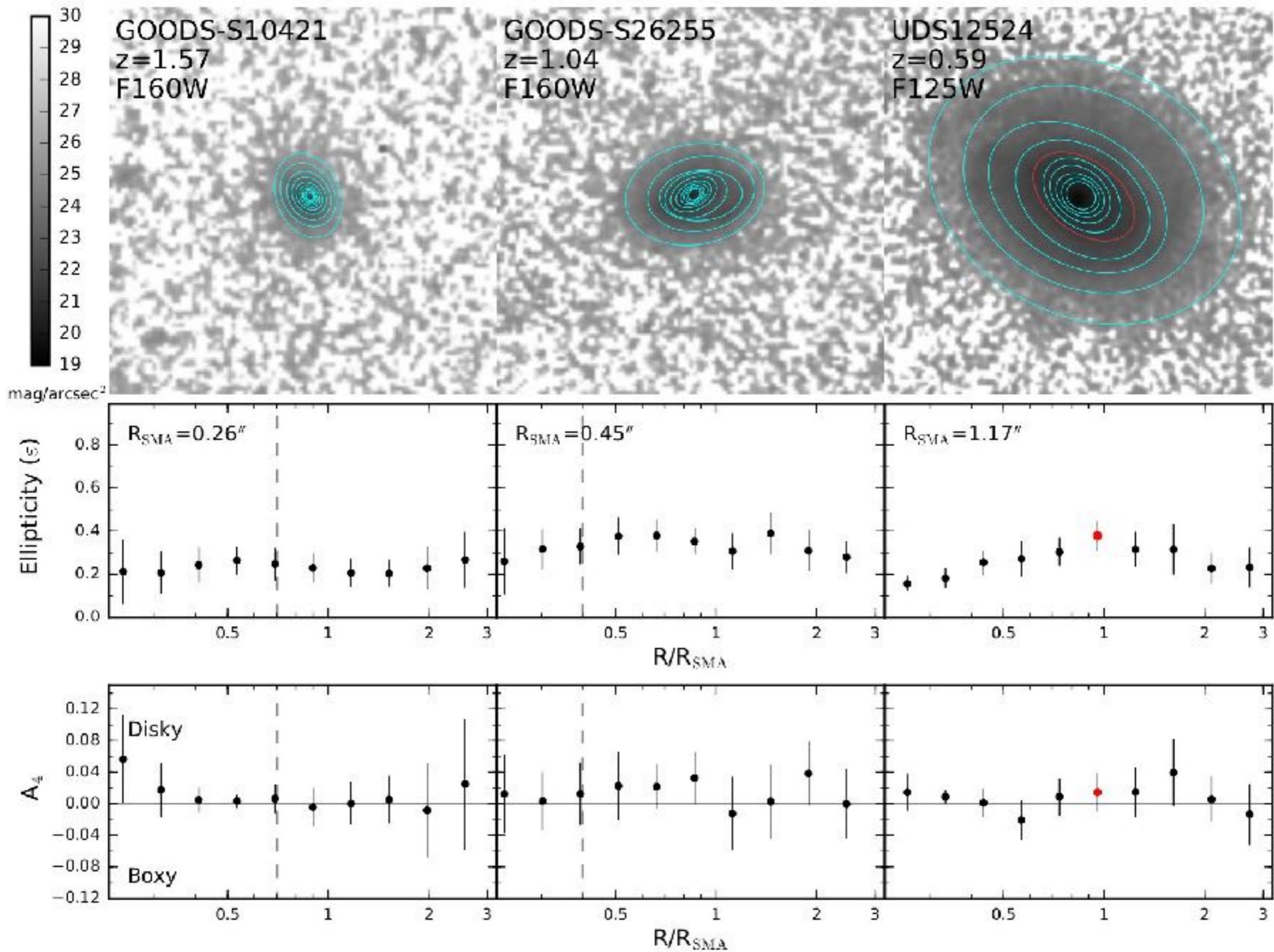
**Axis ratio, Boxy/Disky  $A_4$ , PAs,... as a function of radius**

**(F160w and F125w)**

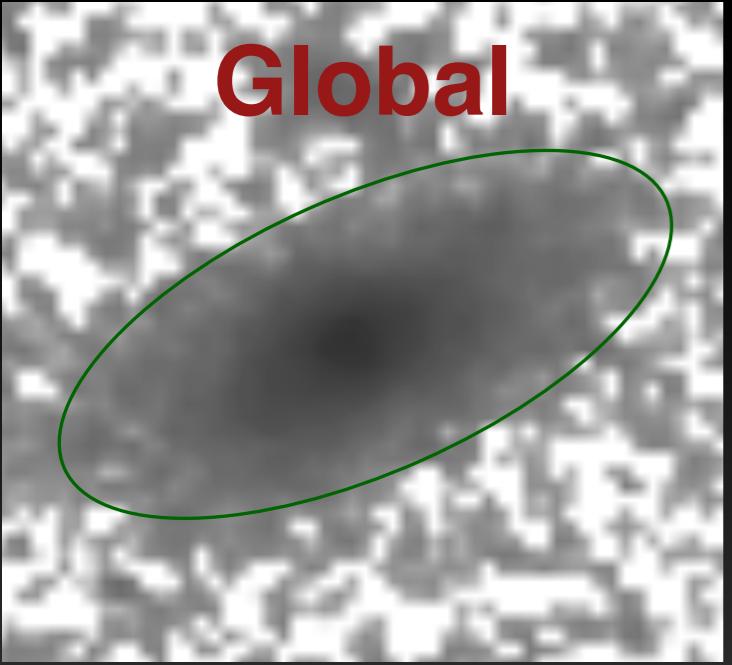
# Part I: Isophotal Photometry (Fourier Expansion)



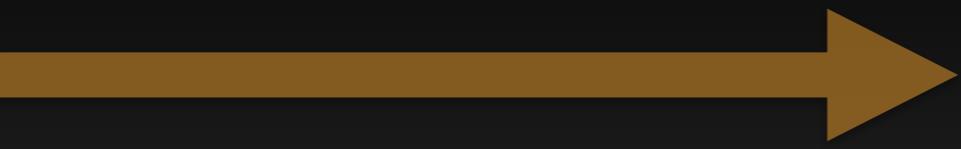
# Part I: Isophotal Photometry (Fourier Expansion)



**Global**

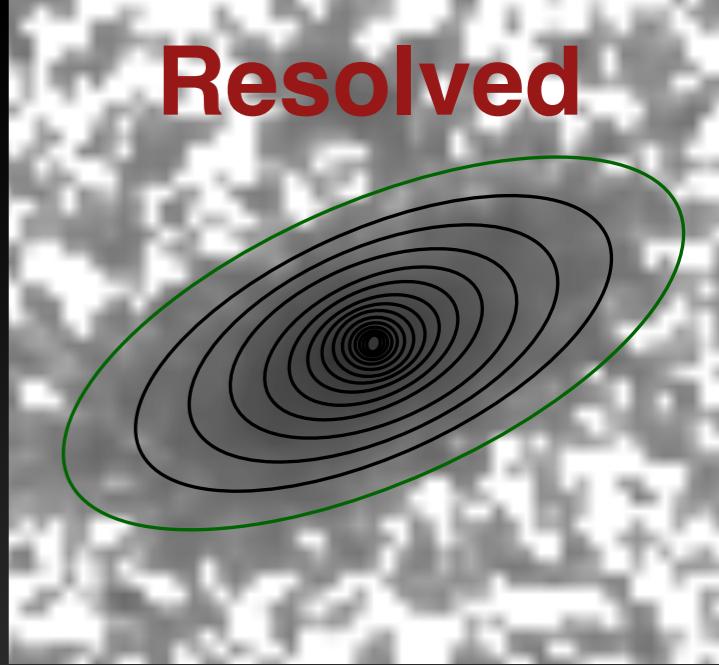


**CANDELS-VAPC: HST Multi-Aperture  
Photometric Catalogs**



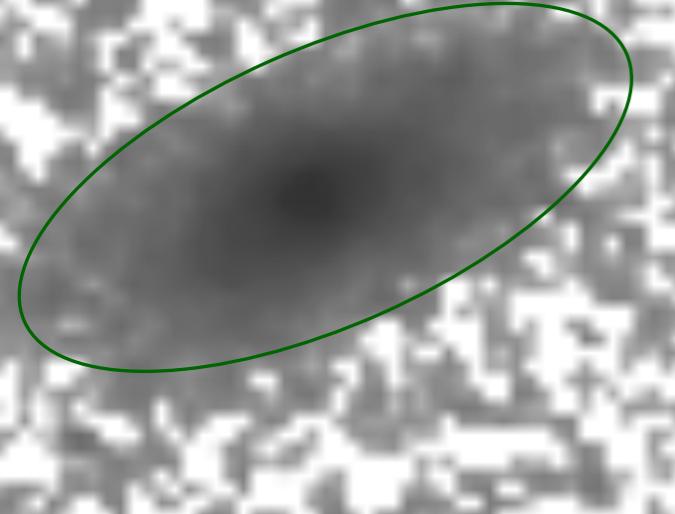
*(Fengshan Liu, Dongfei Jiang,  
David C. Koo, S. M. Faber,  
Yicheng Guo, and CANDELS Team)*

**Resolved**

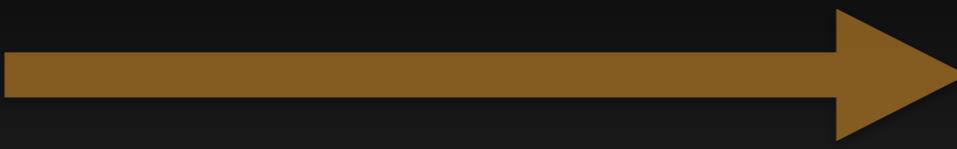


**Part I : Isophotal Aperture Photometry (Fourier Expansion)  
Axis ratio, Boxy/Disky  $A_4$ , PAs,... as a function of radius  
(F160w and F125w)**

**Global**

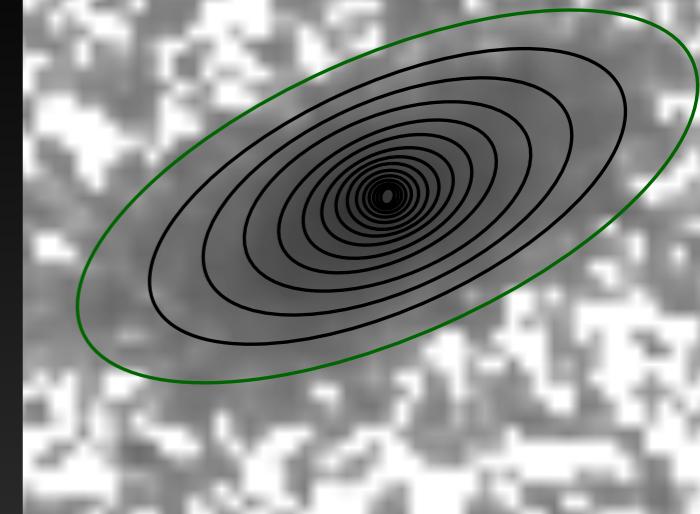


## CANDELS-VAPC: HST Multi-Aperture Photometric Catalogs



*(Fengshan Liu, Dongfei Jiang,  
David C. Koo, S. M. Faber,  
Yicheng Guo, and CANDELS Team)*

**Resolved**



### Part I : Isophotal Aperture Photometry (Fourier Expansion)

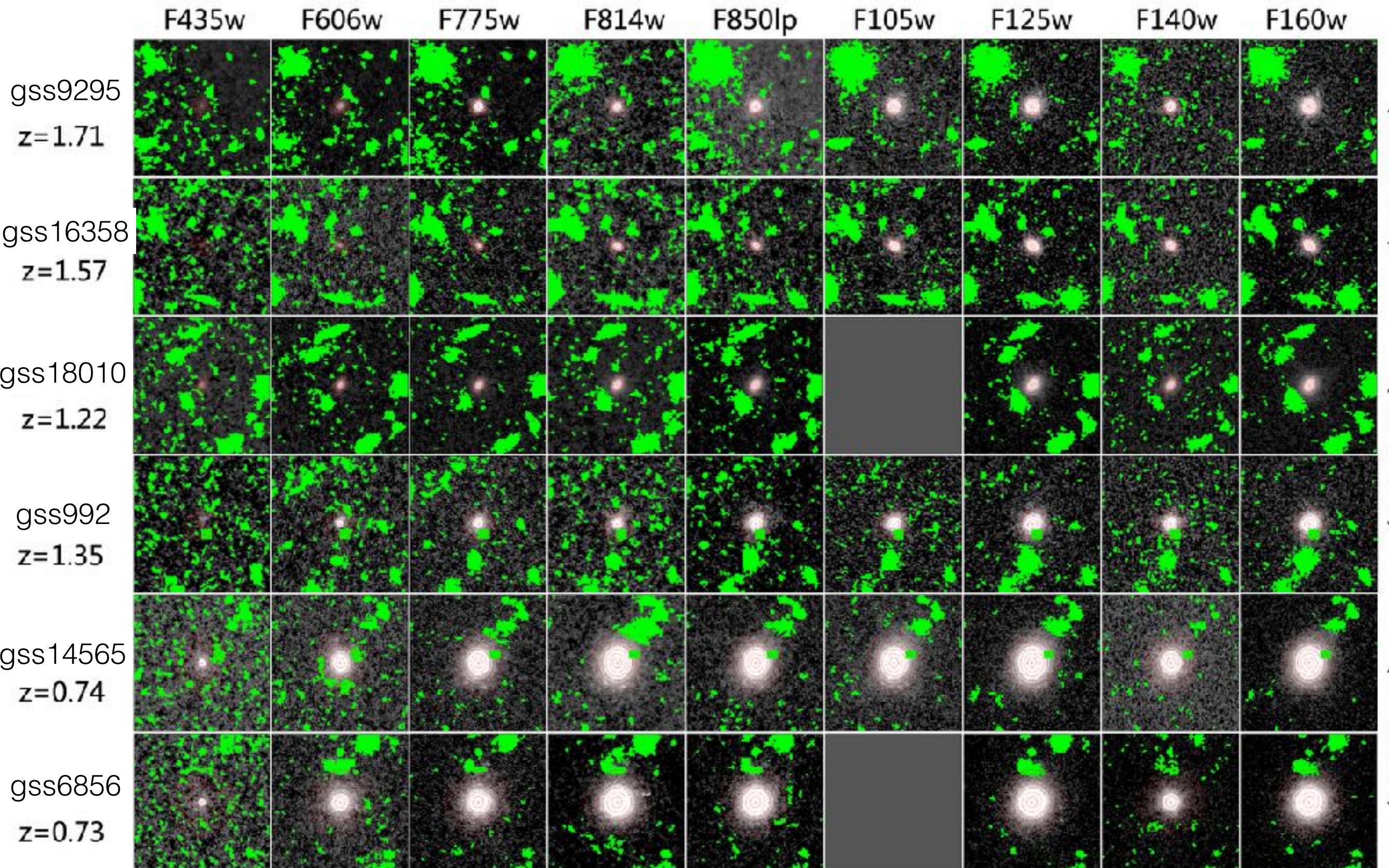
Axis ratio, Boxy/Disky  $A_4$ , PAs,... as a function of radius  
(F160w and F125w)

### Part II: Elliptical Aperture Photometry

(the fixed centre, axis ratio, and PA)

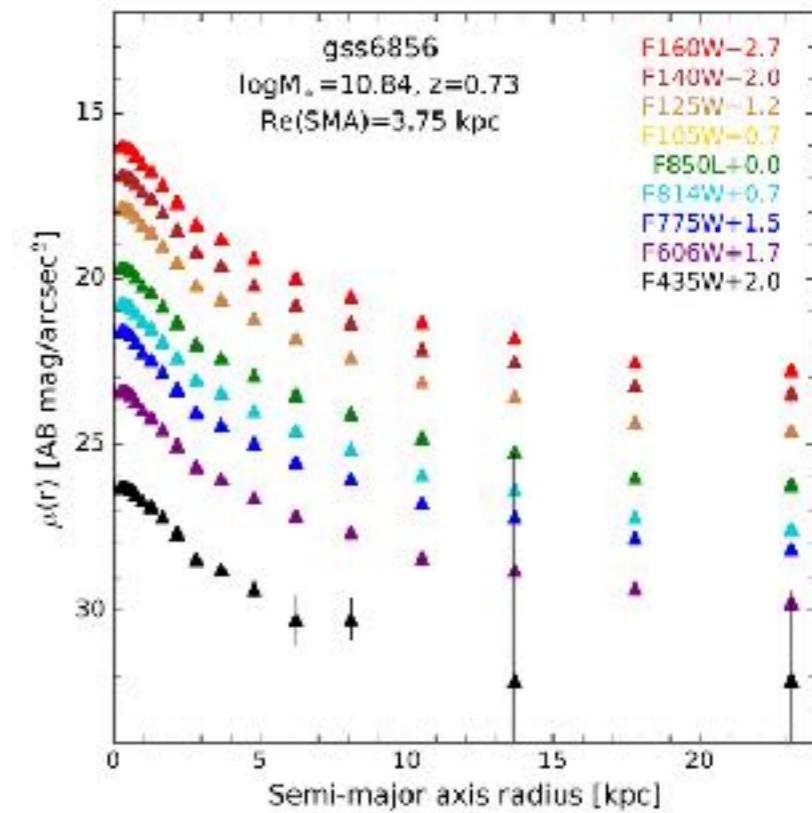
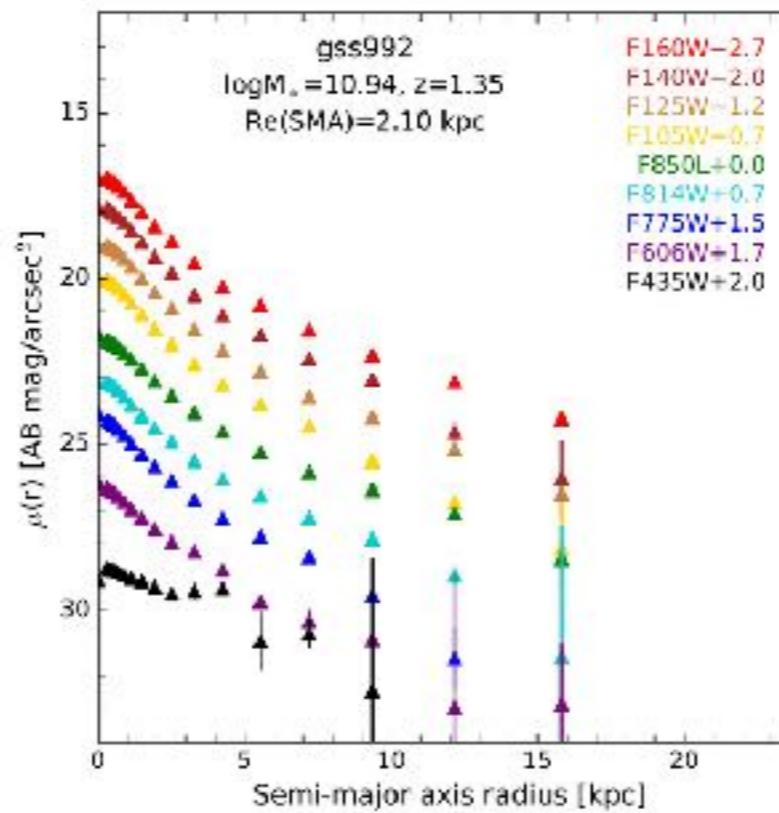
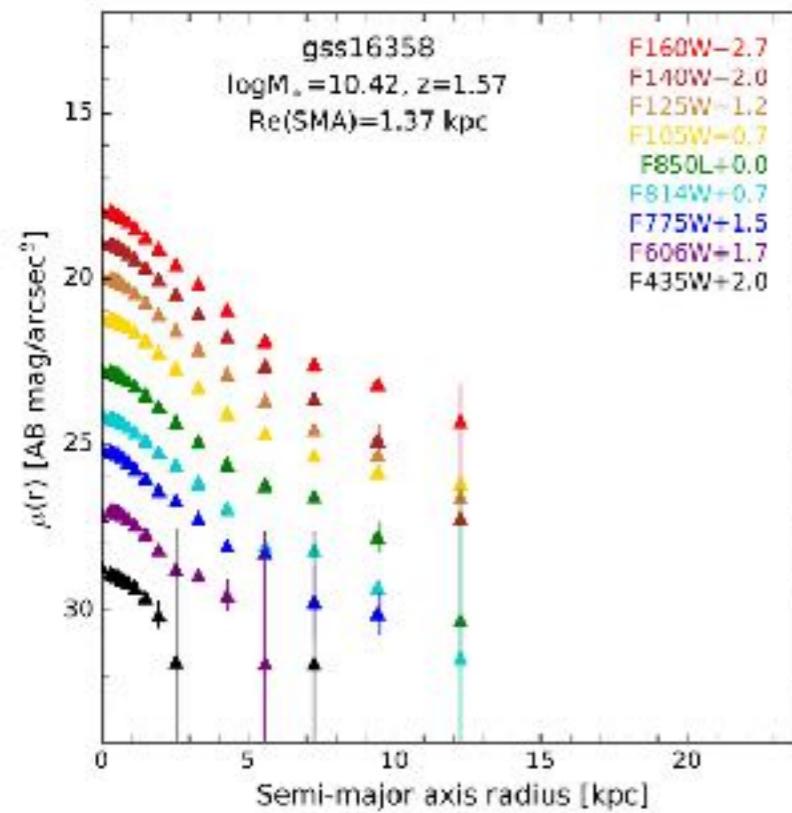
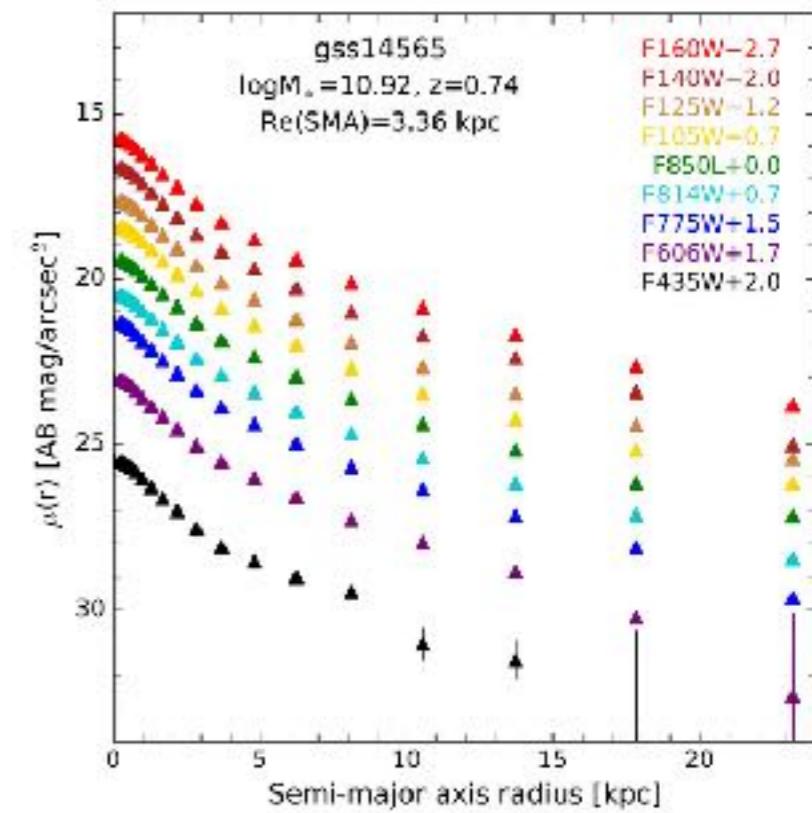
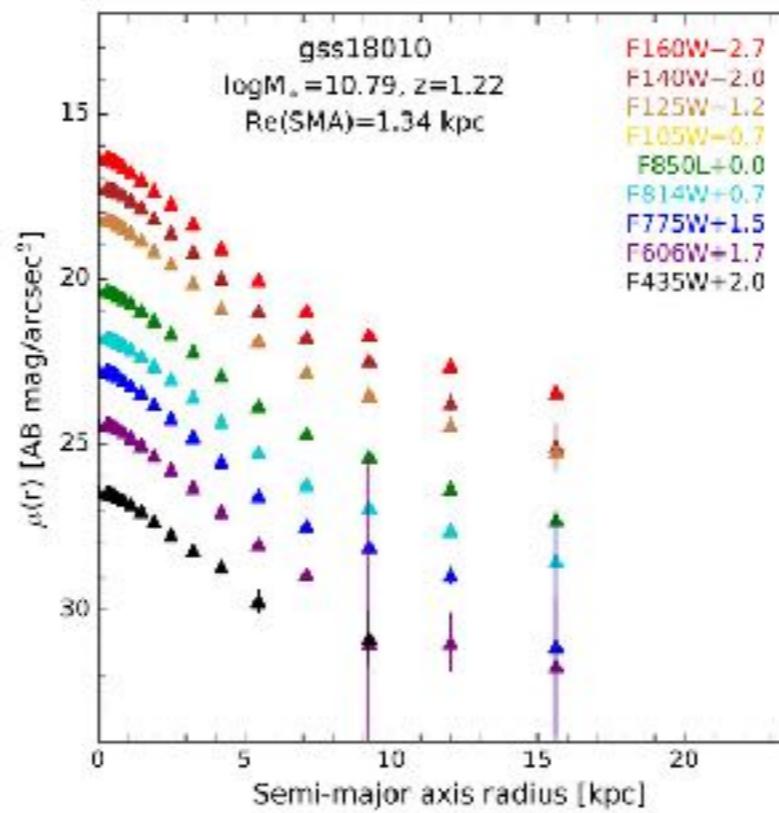
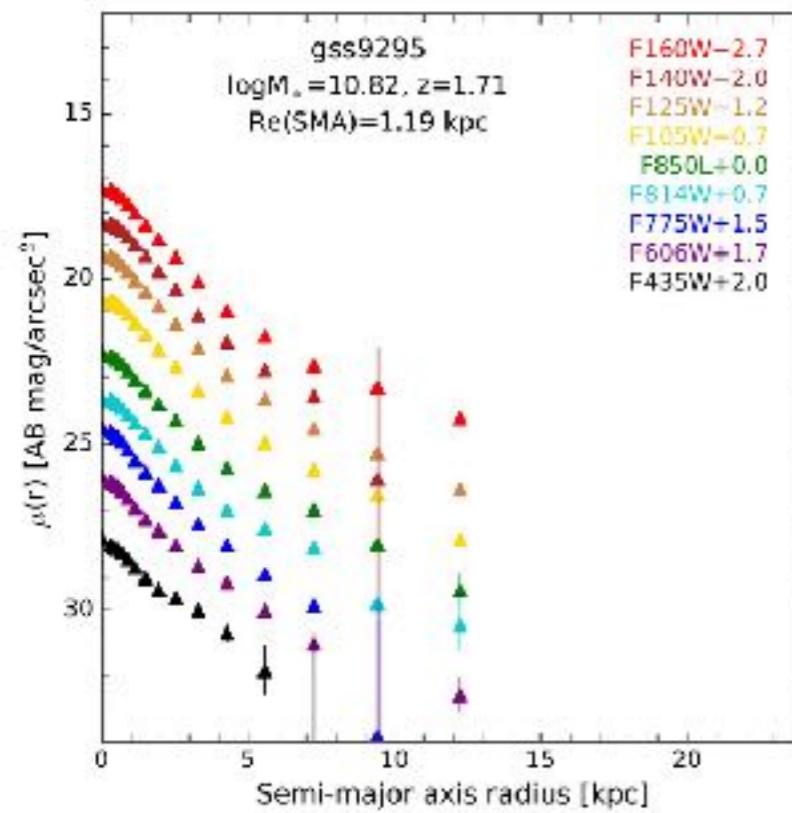
1. Multi-wavelength surface brightness profiles
2. Rest-frame colour profiles from EAZY code
3. Stellar mass & SFR profiles from FAST code
4. Cumulative magnitude and mass profiles  
(GOODS-S & GOODS-N: 5 ACS + 4 WFC3 bands  
UDS,COSMOS,EGS: 2 ACS + 3 WFC3 bands)

## Part II: Elliptical Aperture (fixed centre, ellipticity, PA) (Examples of Quiescent Galaxies)

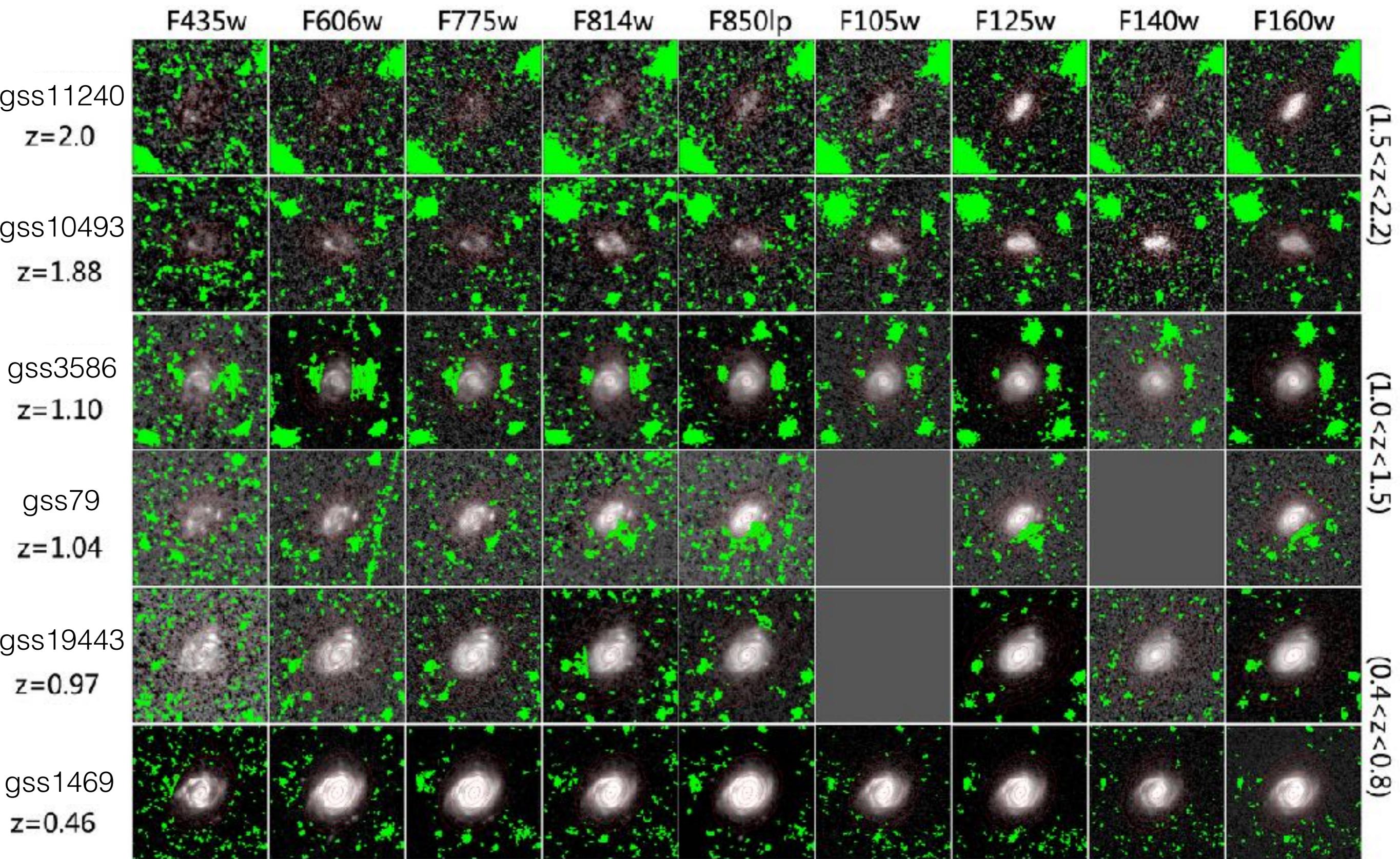


GOODS-S UVJ-classification: quiescent galaxies ( $10.5 < \log M^* < 11.0$ ,  $0.4 < z < 2.2$ ); imaging size: 80kpc X 80kpc

# Part II: Elliptical Aperture (fixed centre, ellipticity, PA) (Examples of Quiescent Galaxies)

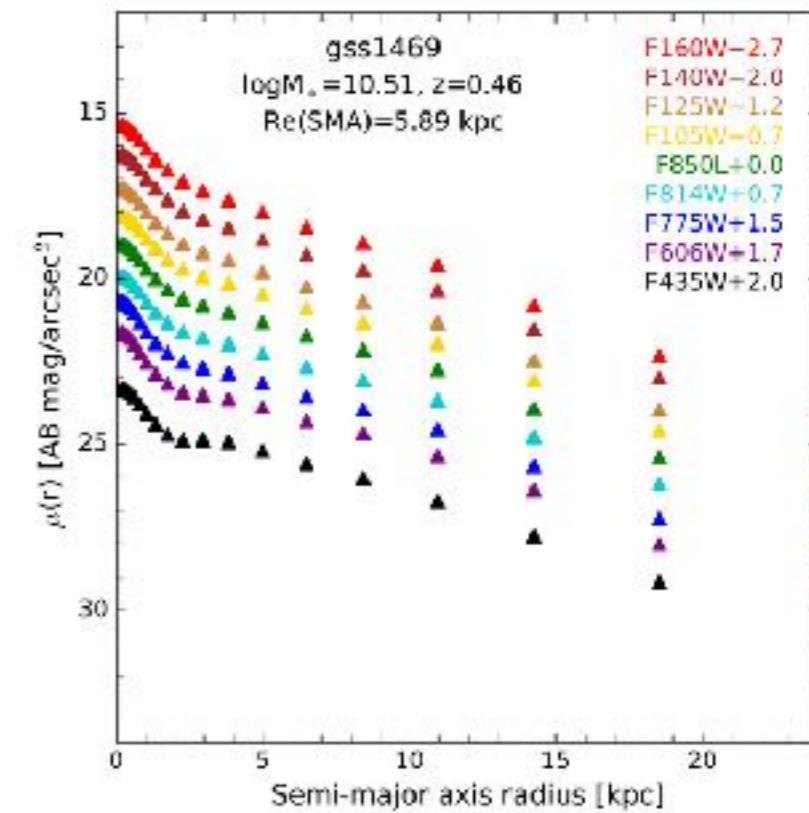
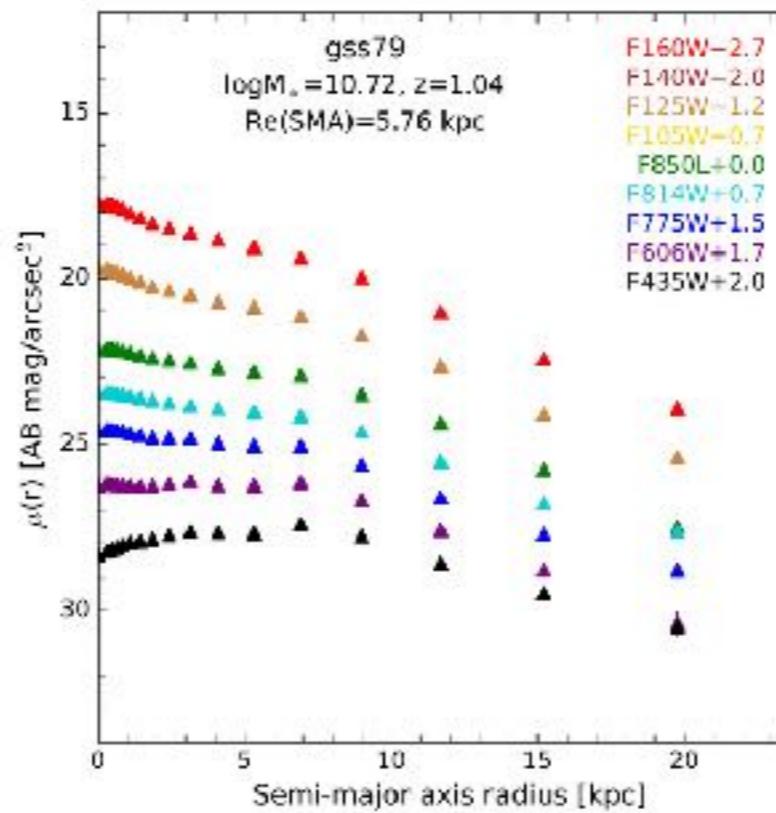
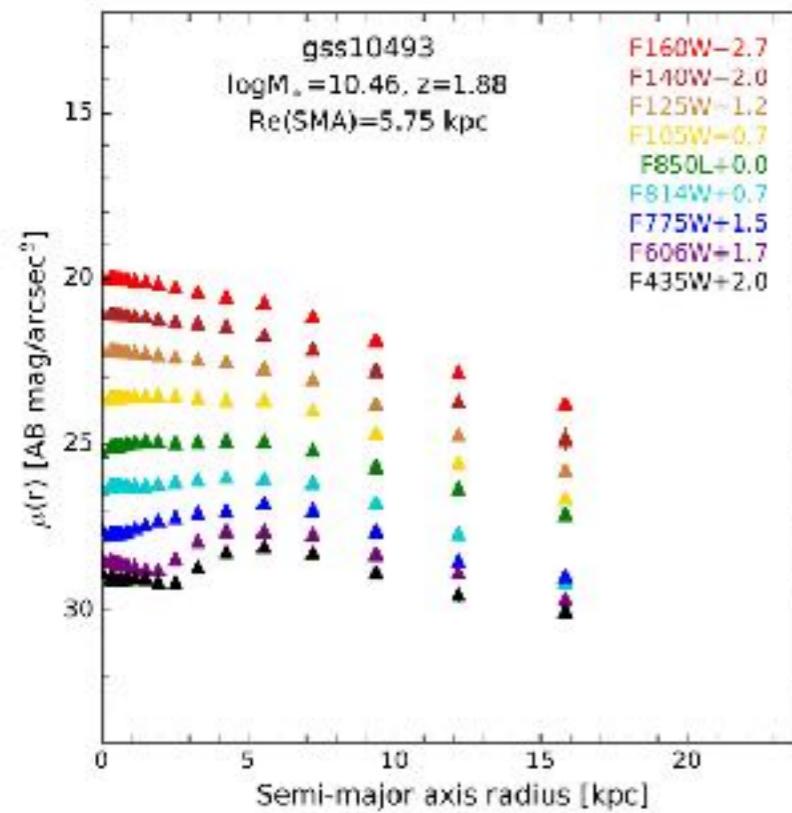
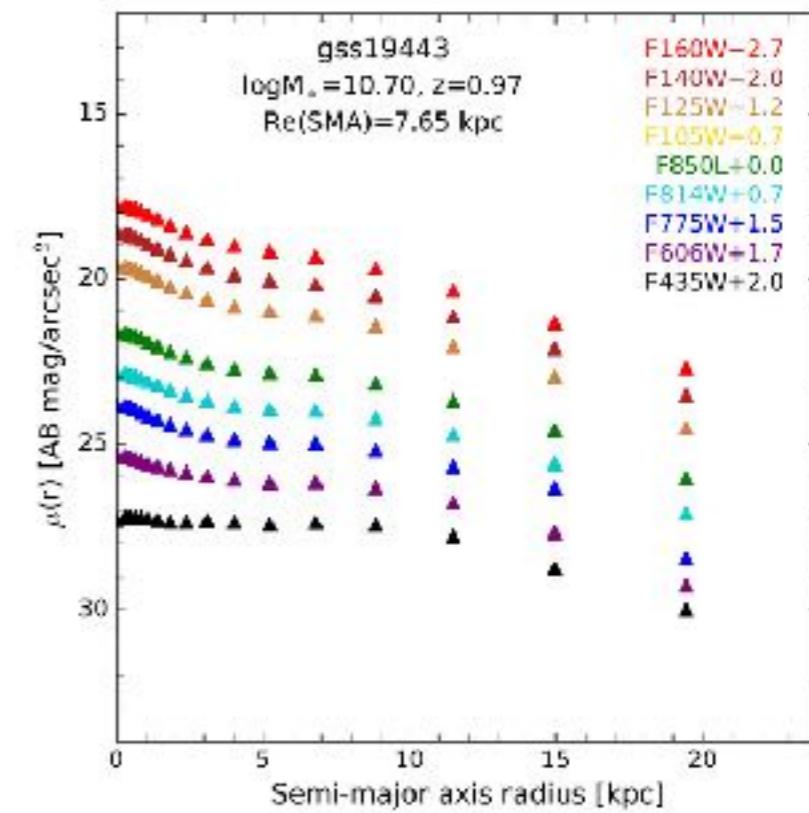
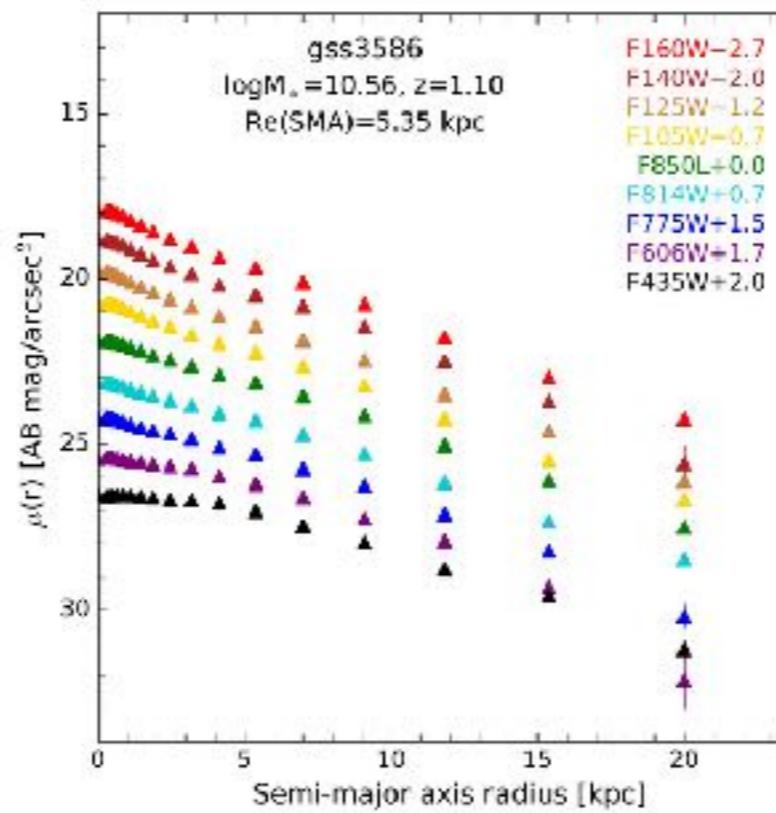
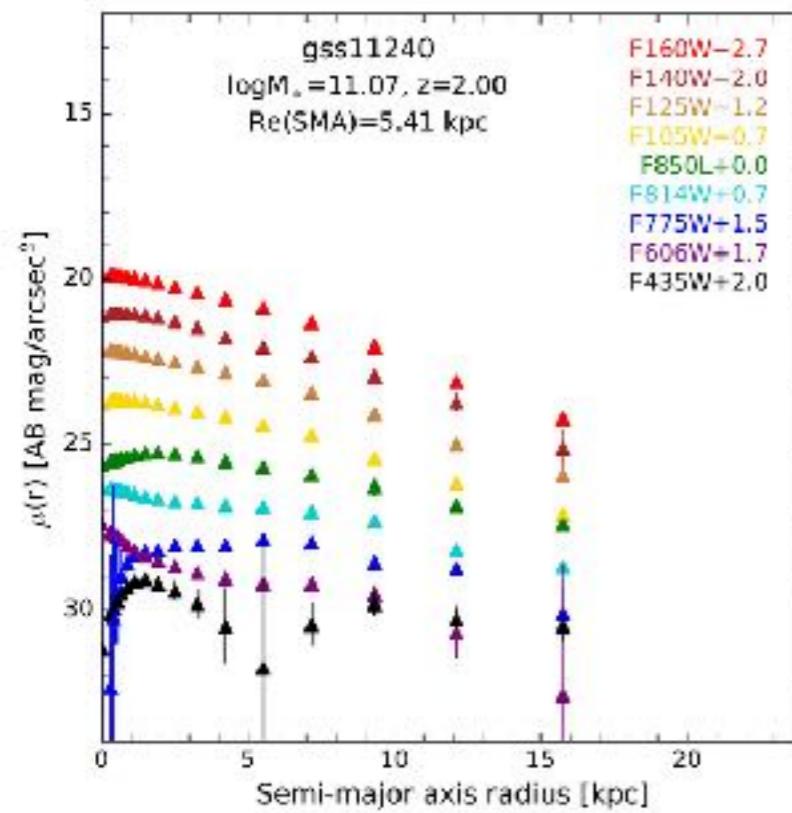


## Part II: Elliptical Aperture (fixed centre, ellipticity, PA) (Examples of Star-Forming Galaxies)

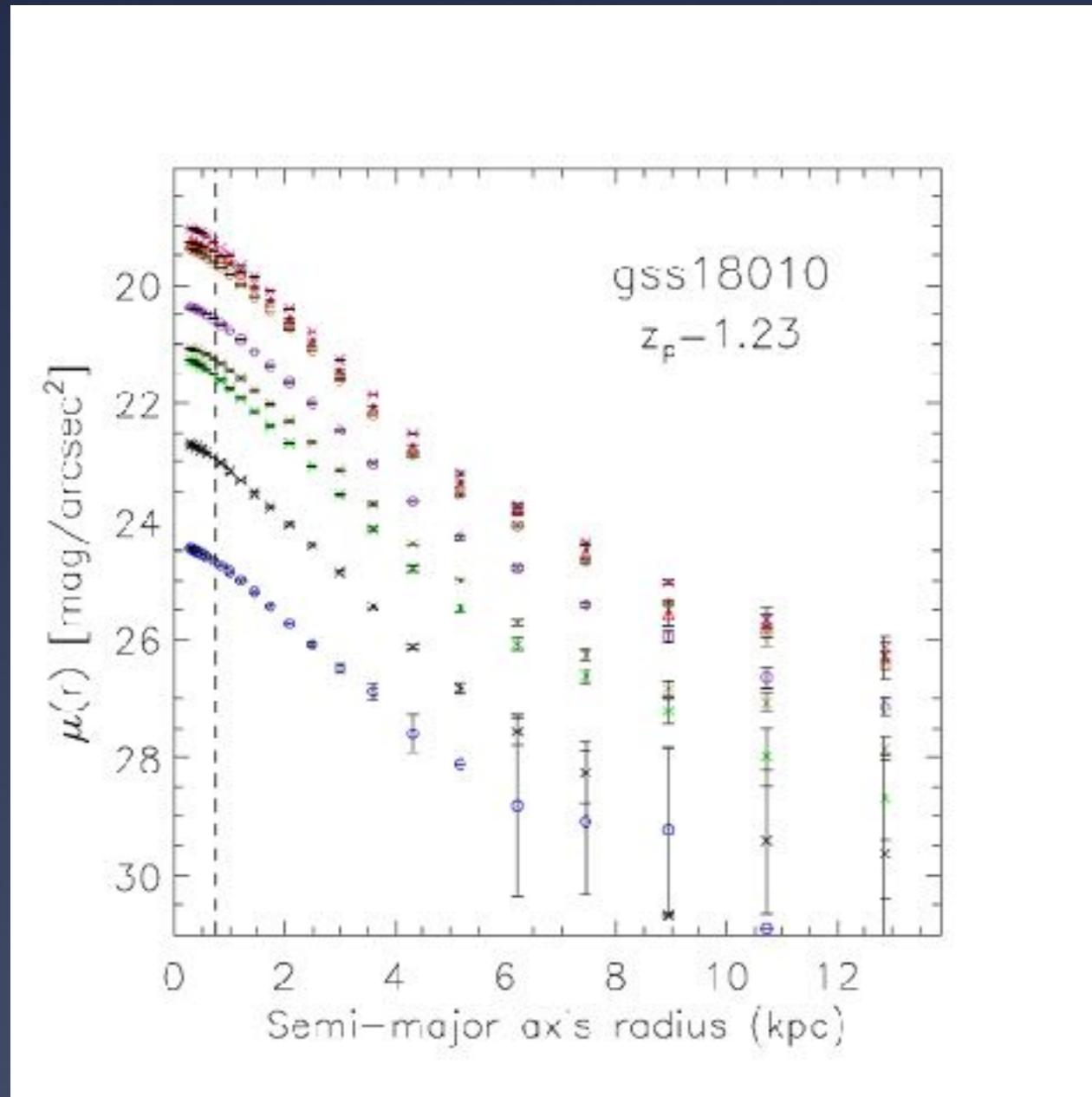


GOODS-S UVJ-classification: star-forming galaxies ( $10.5 < \log M^* < 11.0$ ,  $0.4 < z < 2.2$ ); imaging size: 80kpc X 80kpc

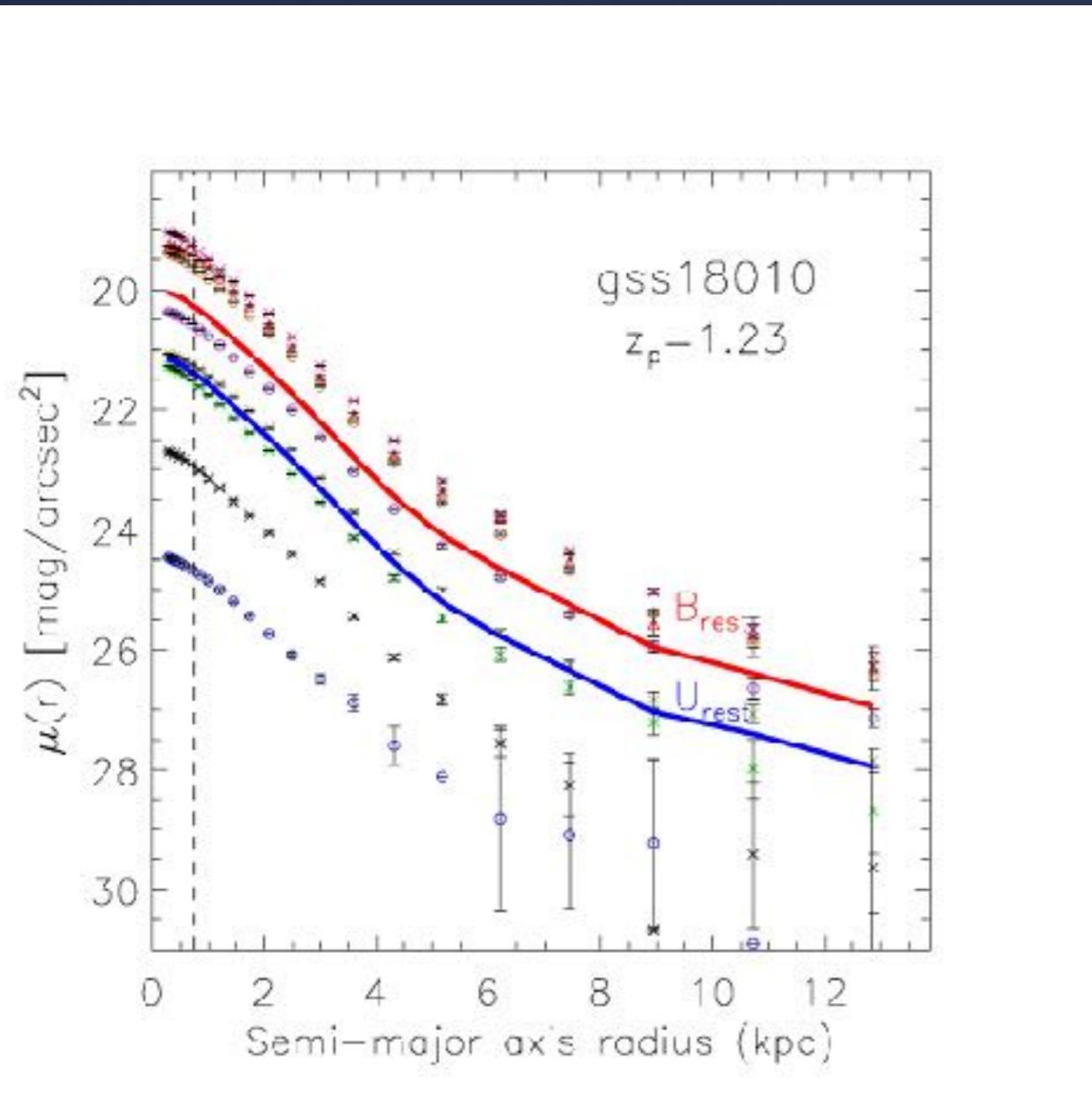
# Part II: Elliptical Aperture (fixed centre, ellipticity, PA) (Examples of Star-Forming Galaxies)



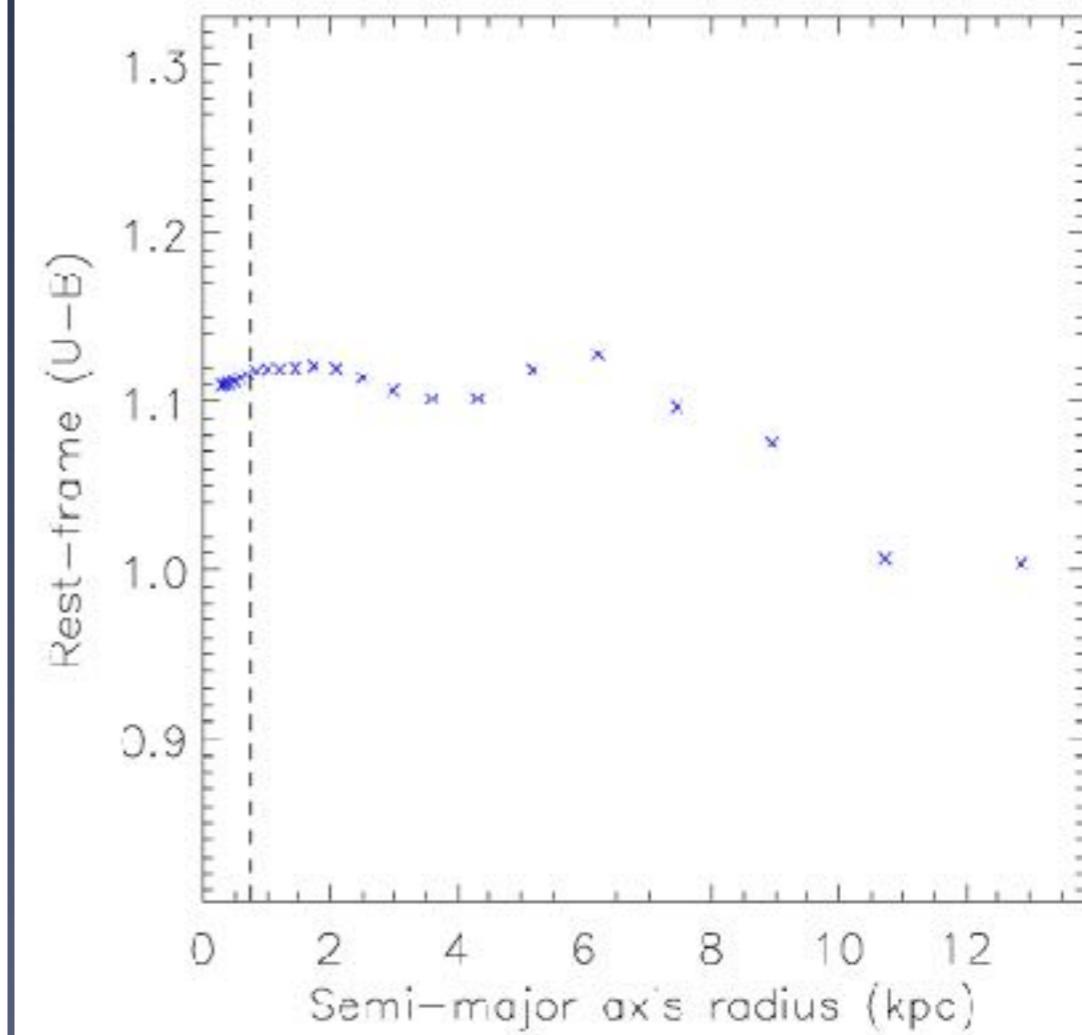
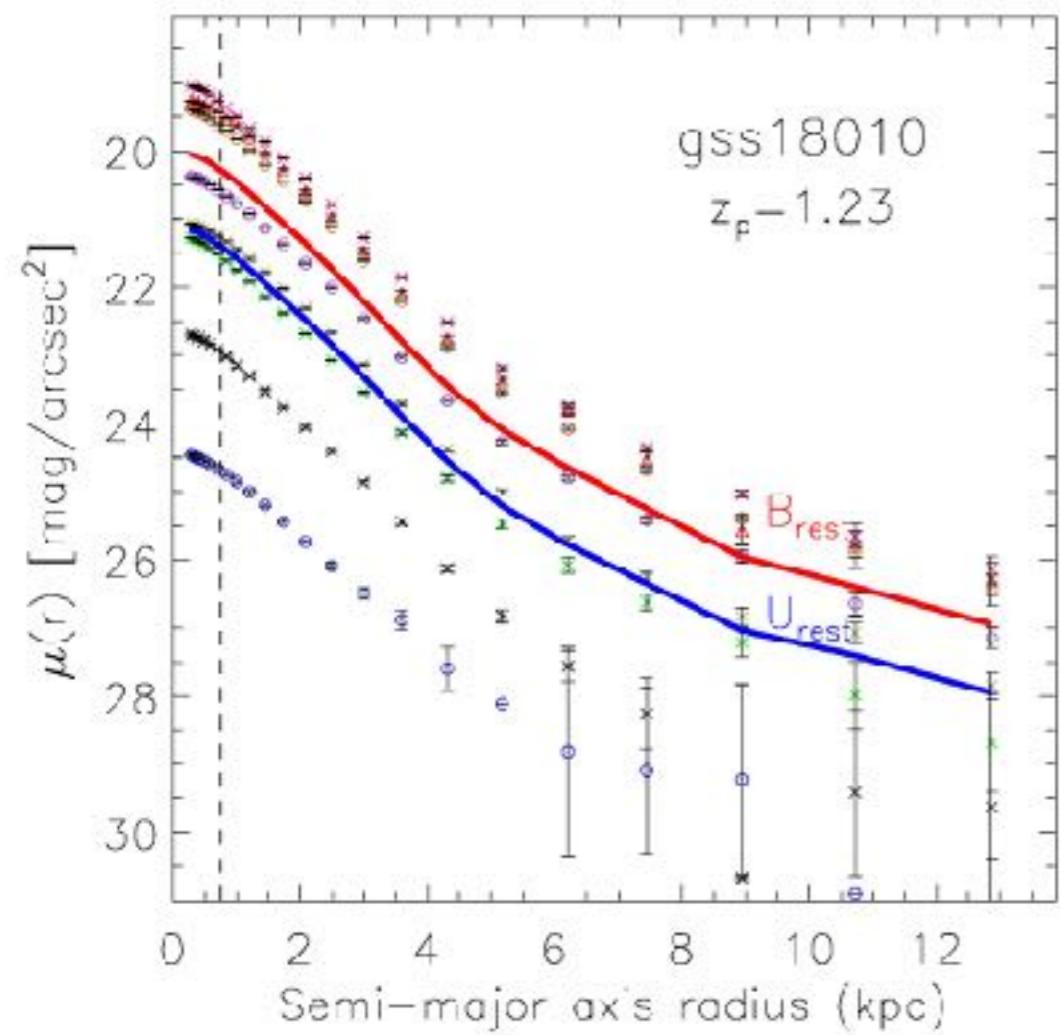
# Rest-frame color profiles from EAZY code



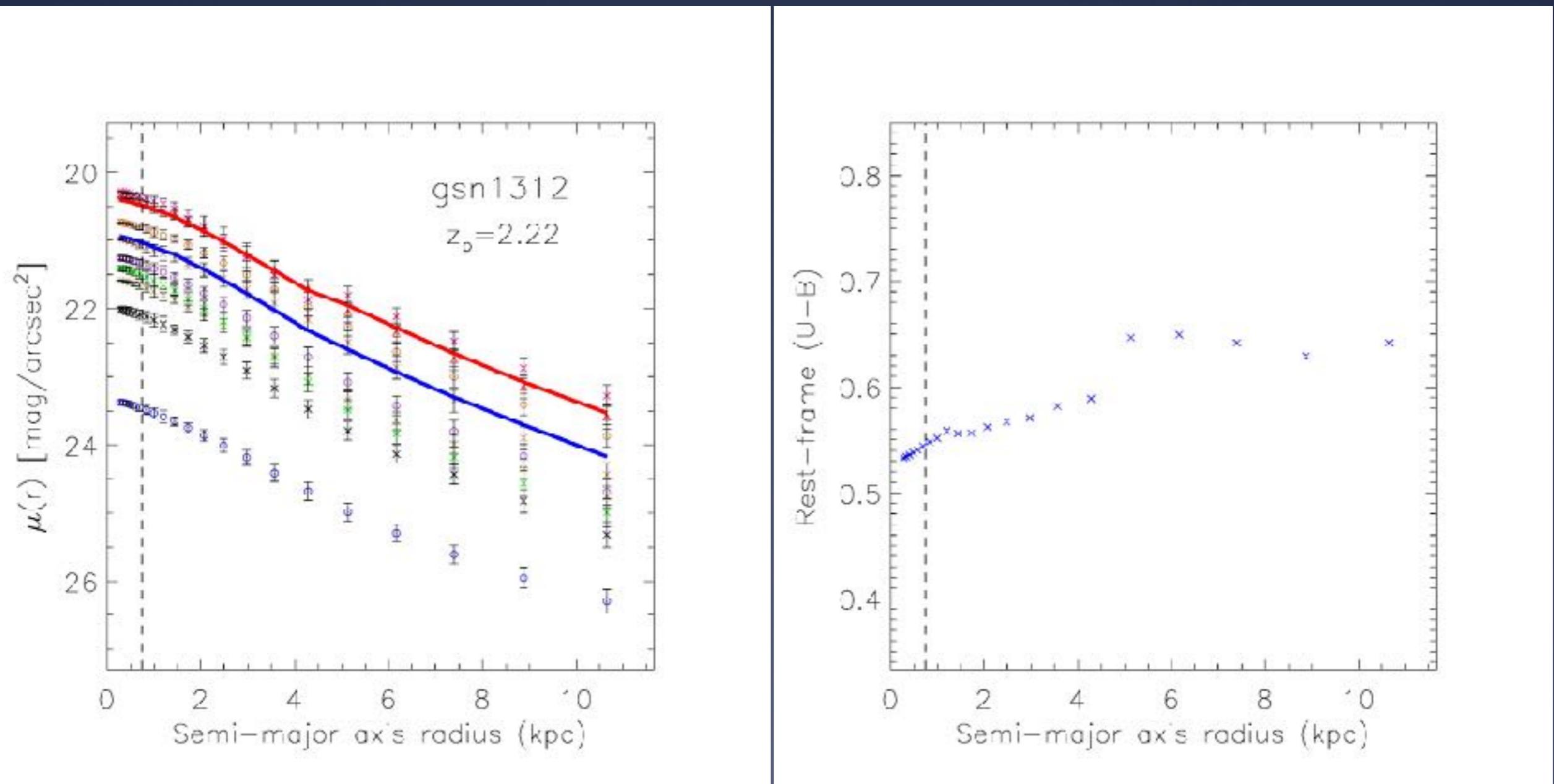
# Rest-frame color profiles from EAZY code



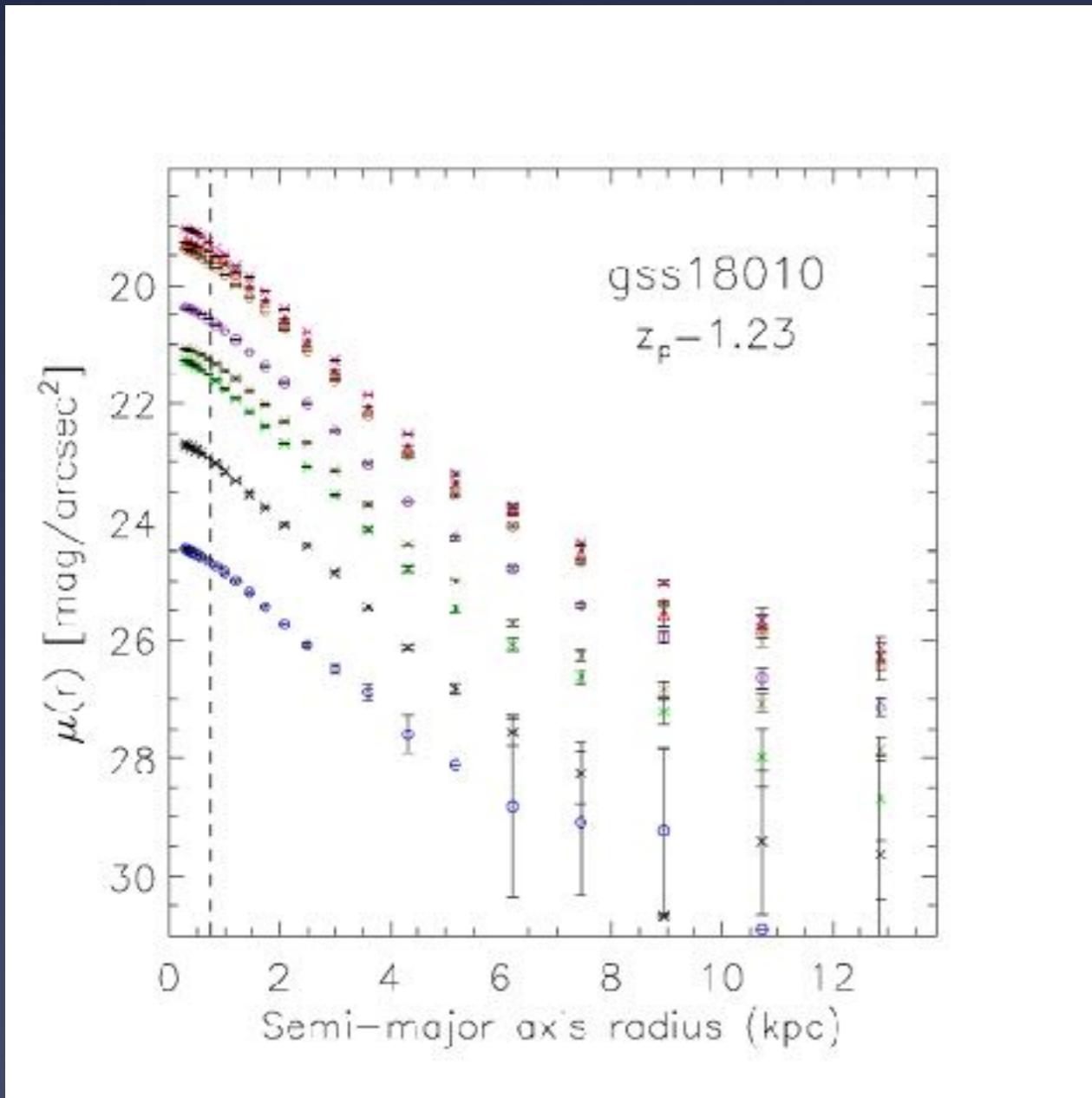
# Rest-frame color profiles from EAZY code



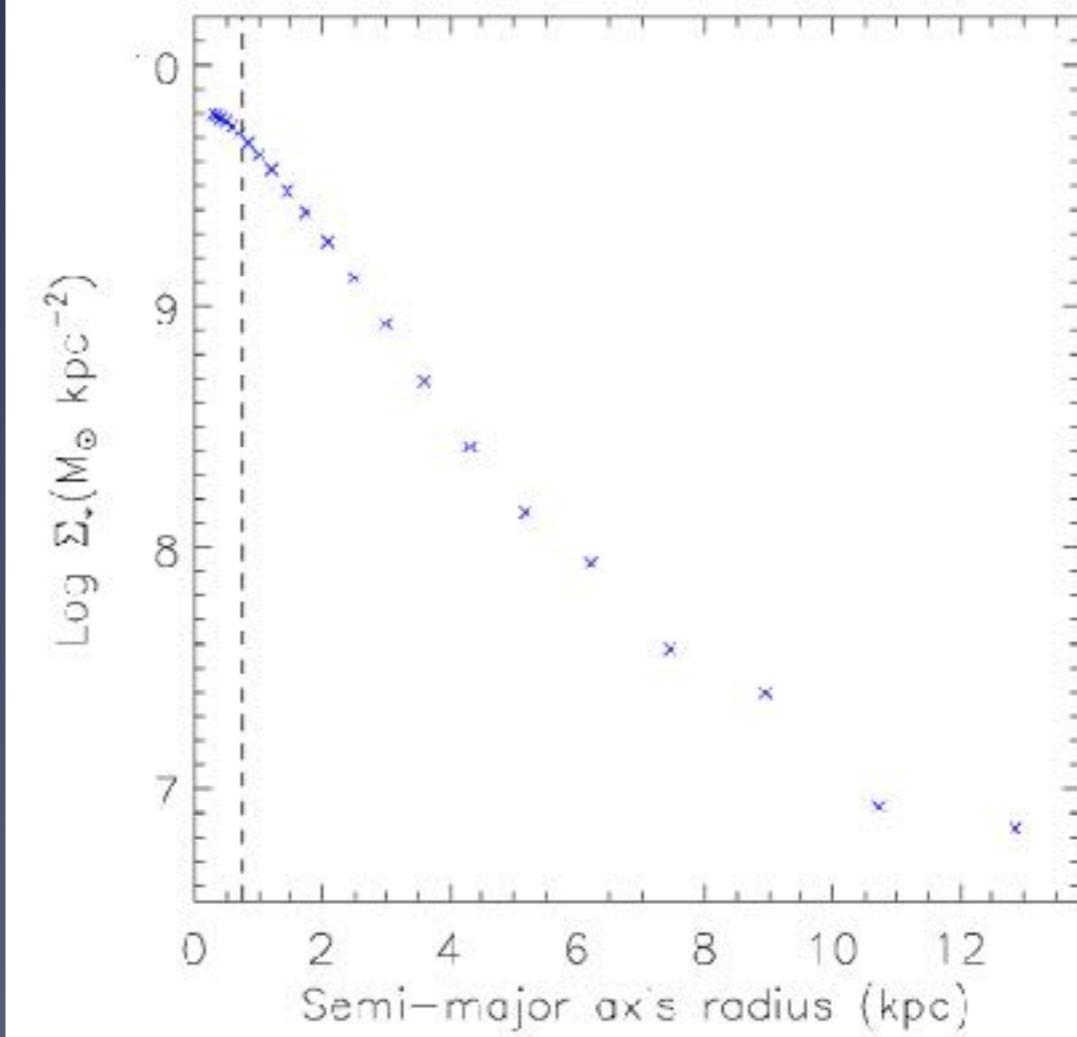
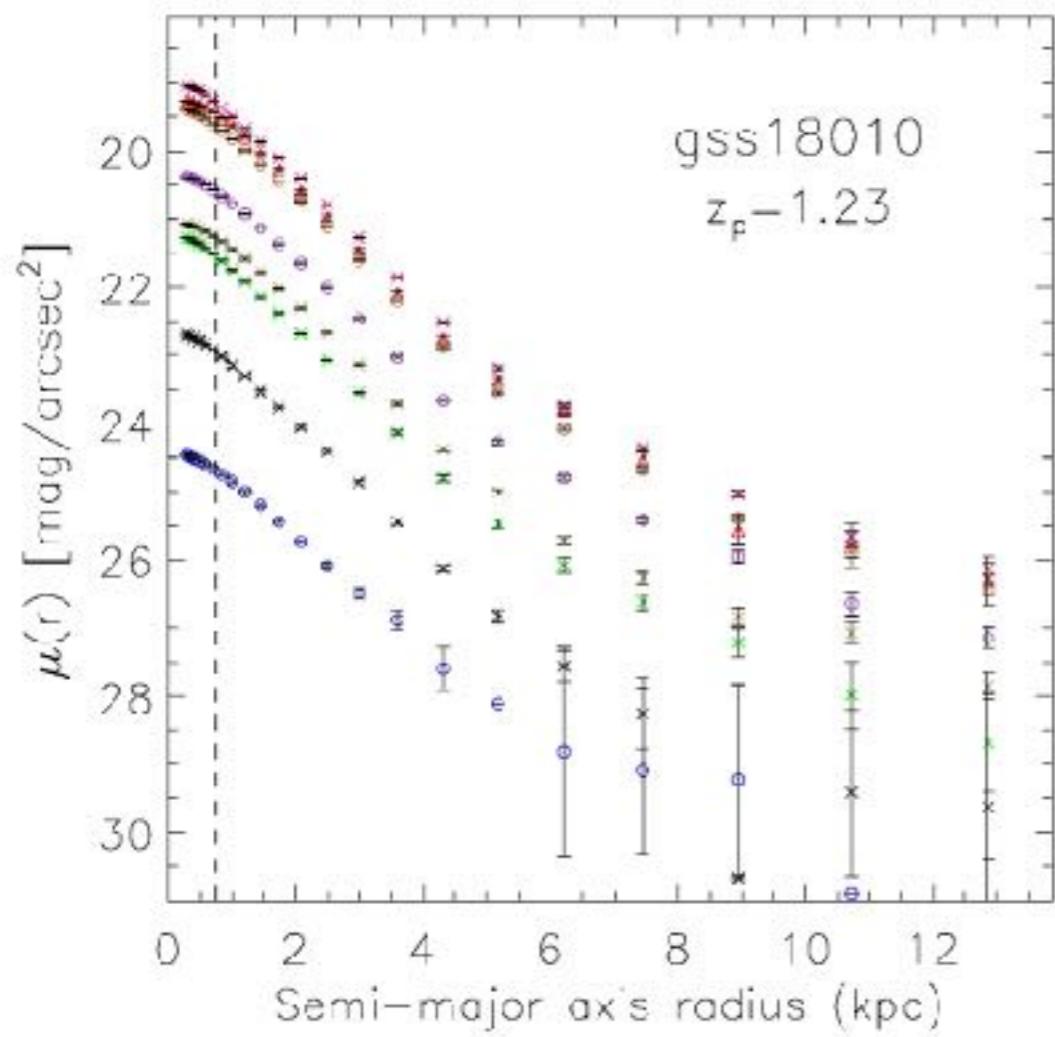
# Rest-frame color profiles from EAZY code



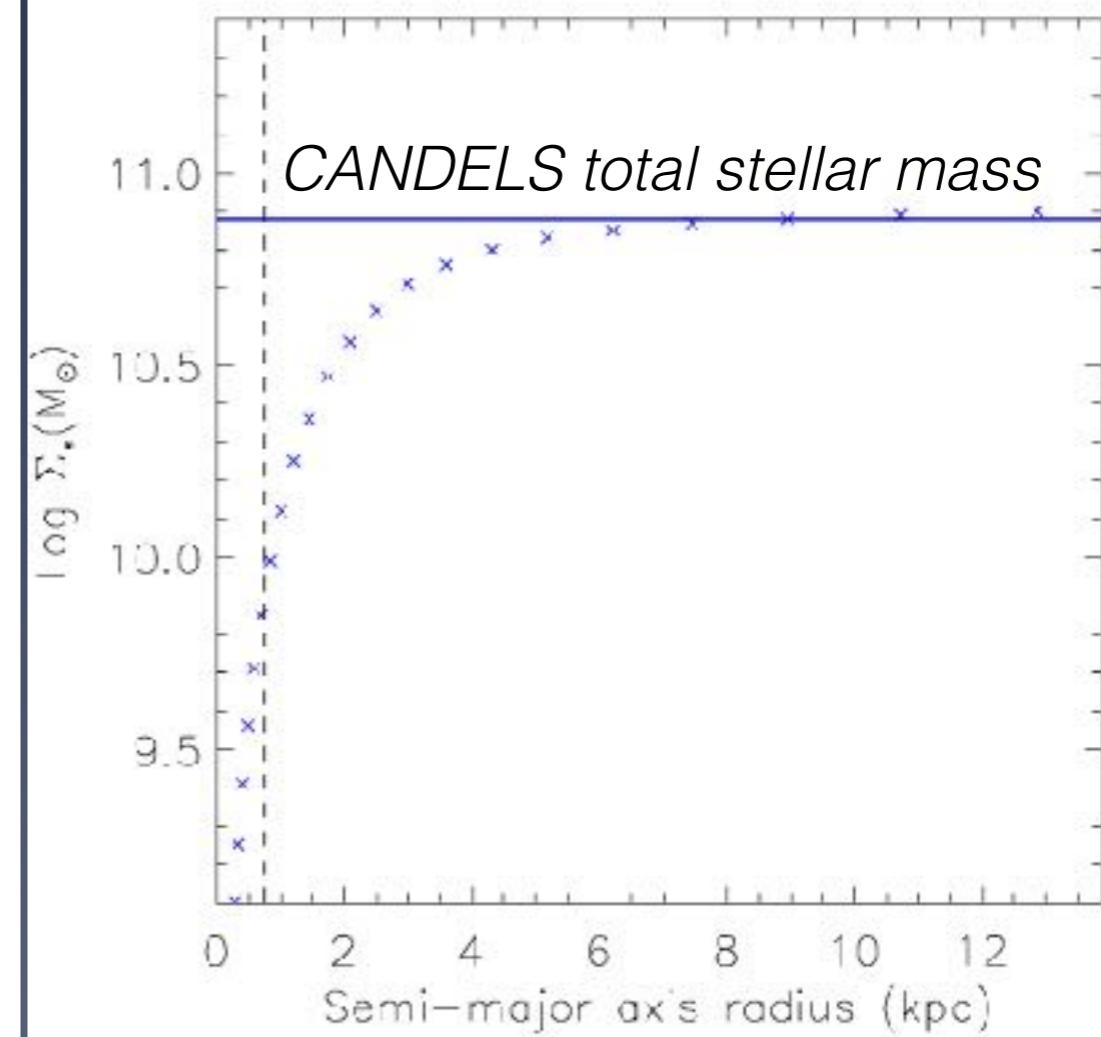
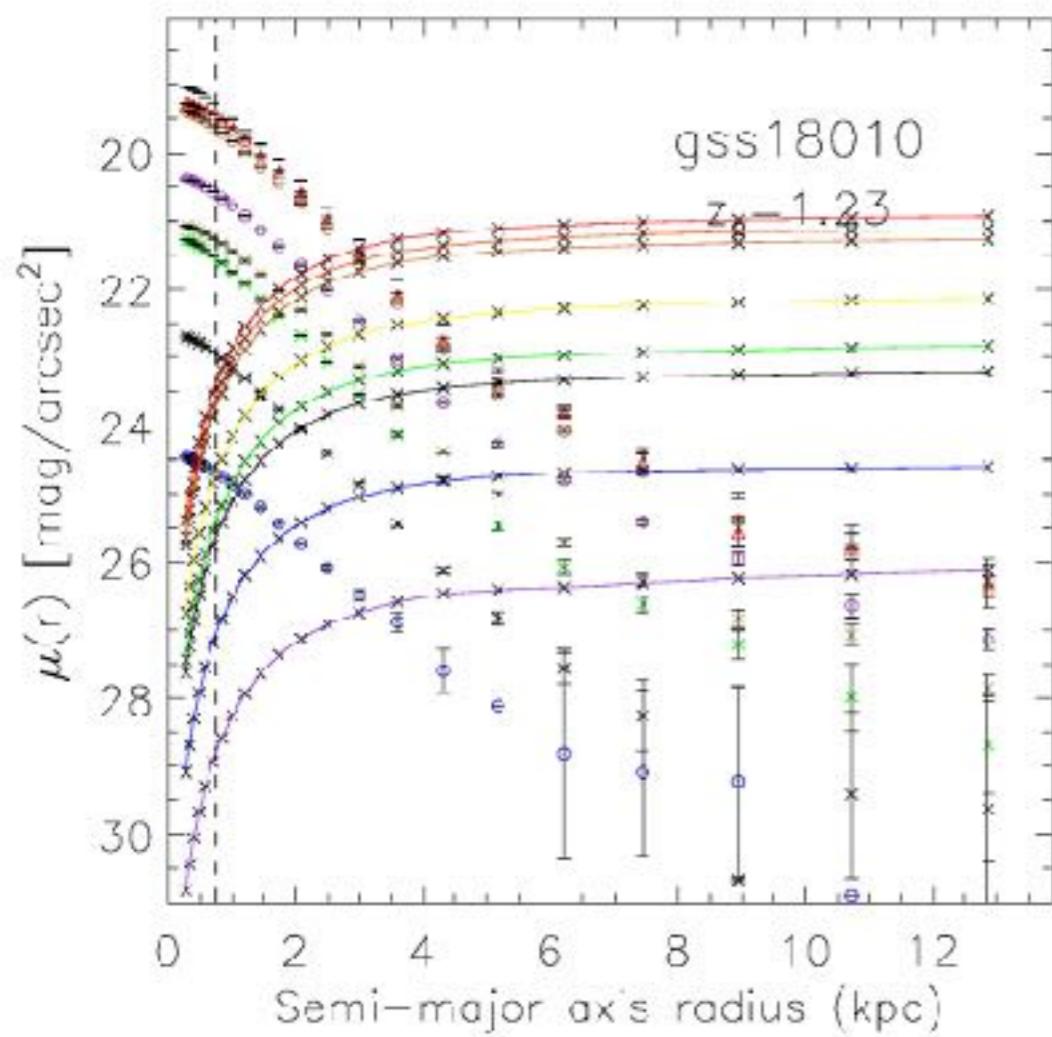
# Stellar mass profiles from FAST code



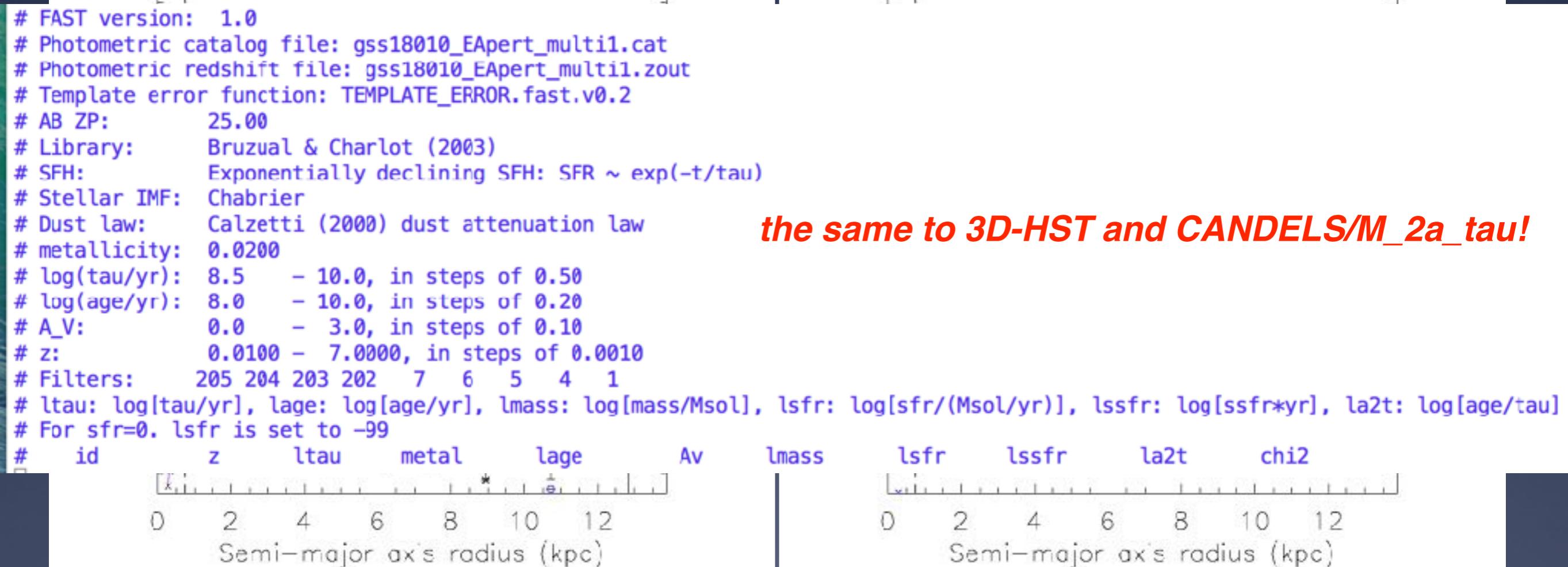
# Stellar mass profiles from FAST code



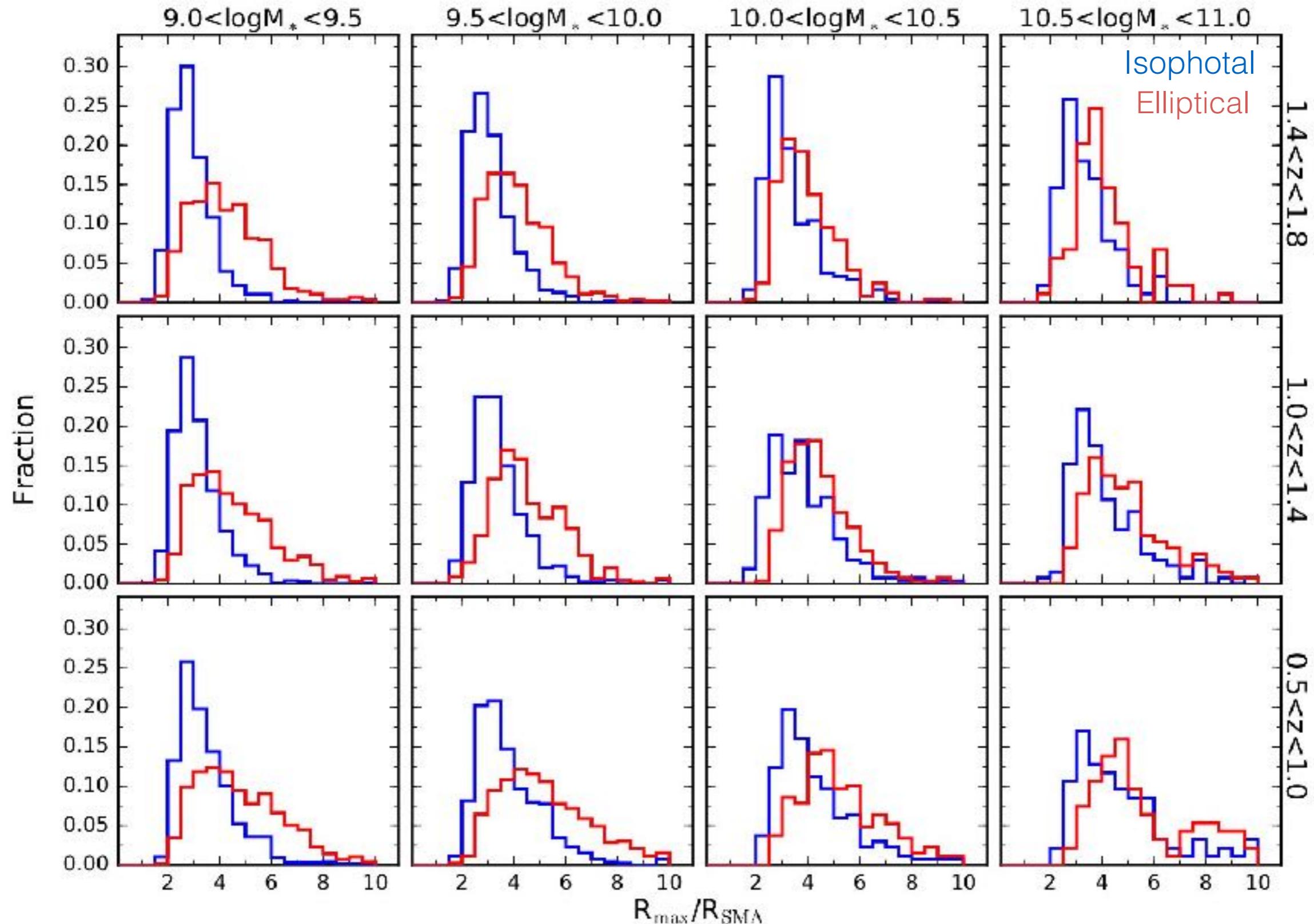
# Stellar mass profiles from FAST code



# Stellar mass profiles from FAST code

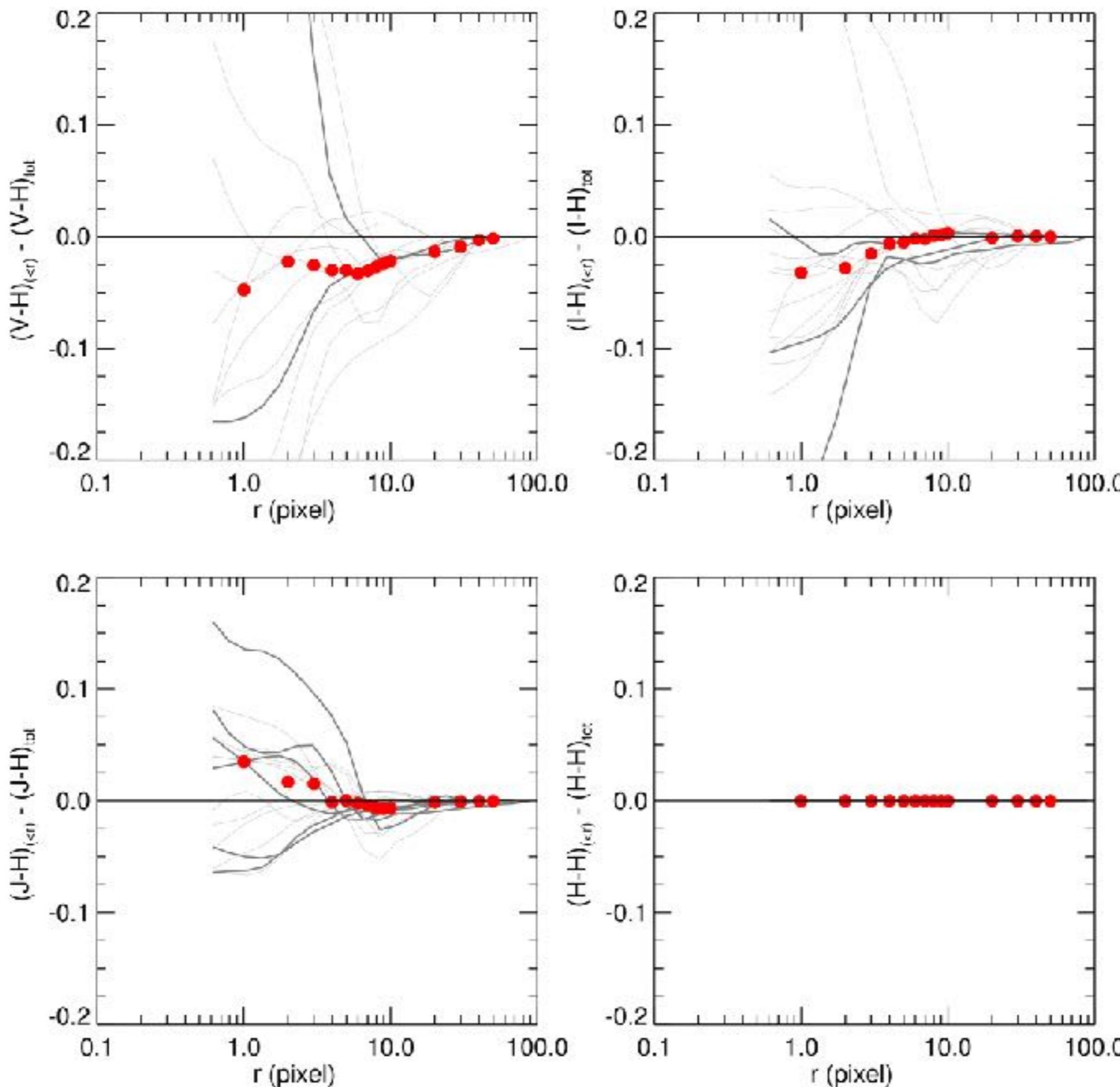


# How far we can measure ( $3^*R_e \sim 4^*R_e$ on average in F160w)



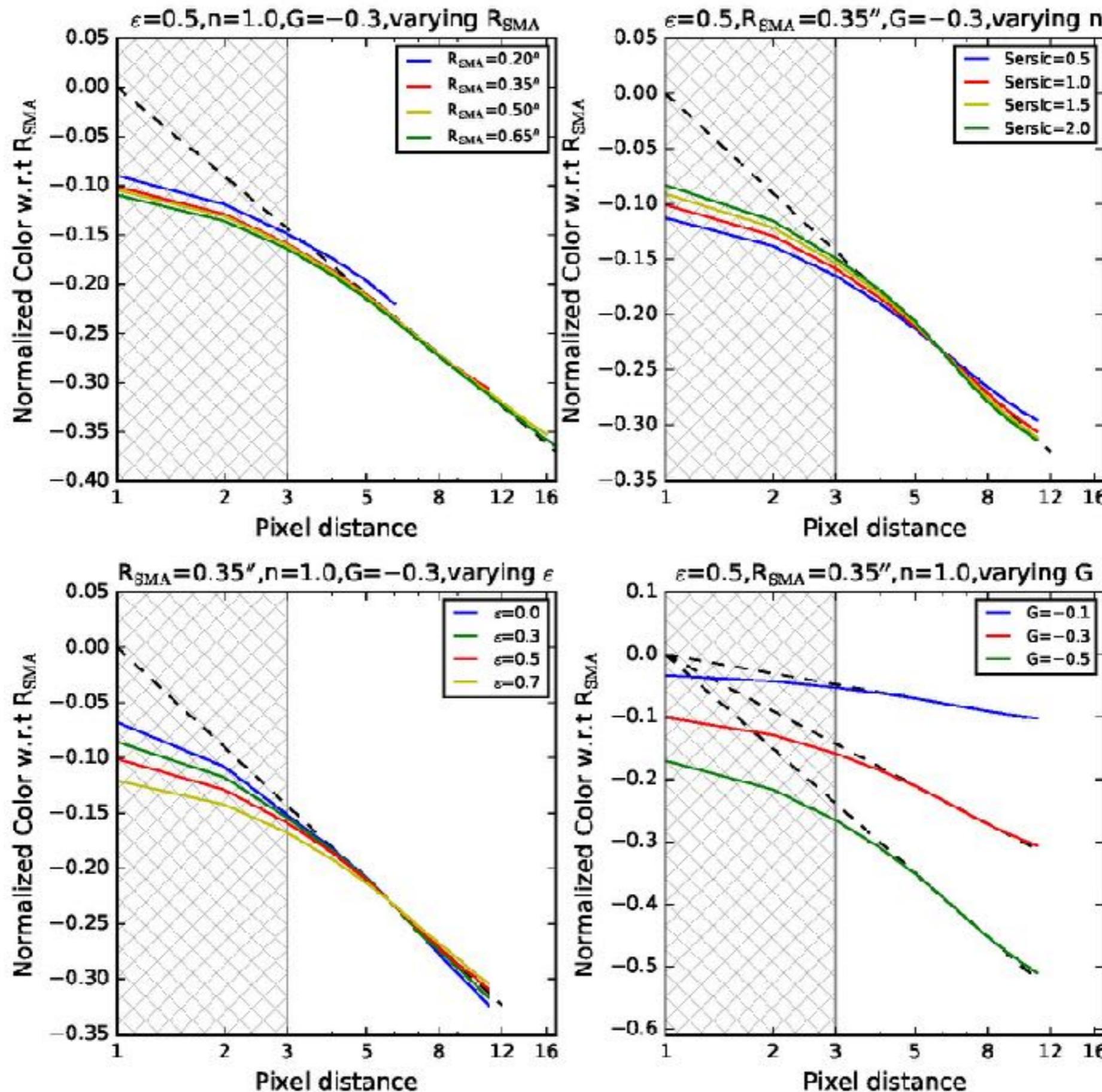
# Effect of PSF mis-matching on colours (unsantaured stars)

credit :Yicheng Guo

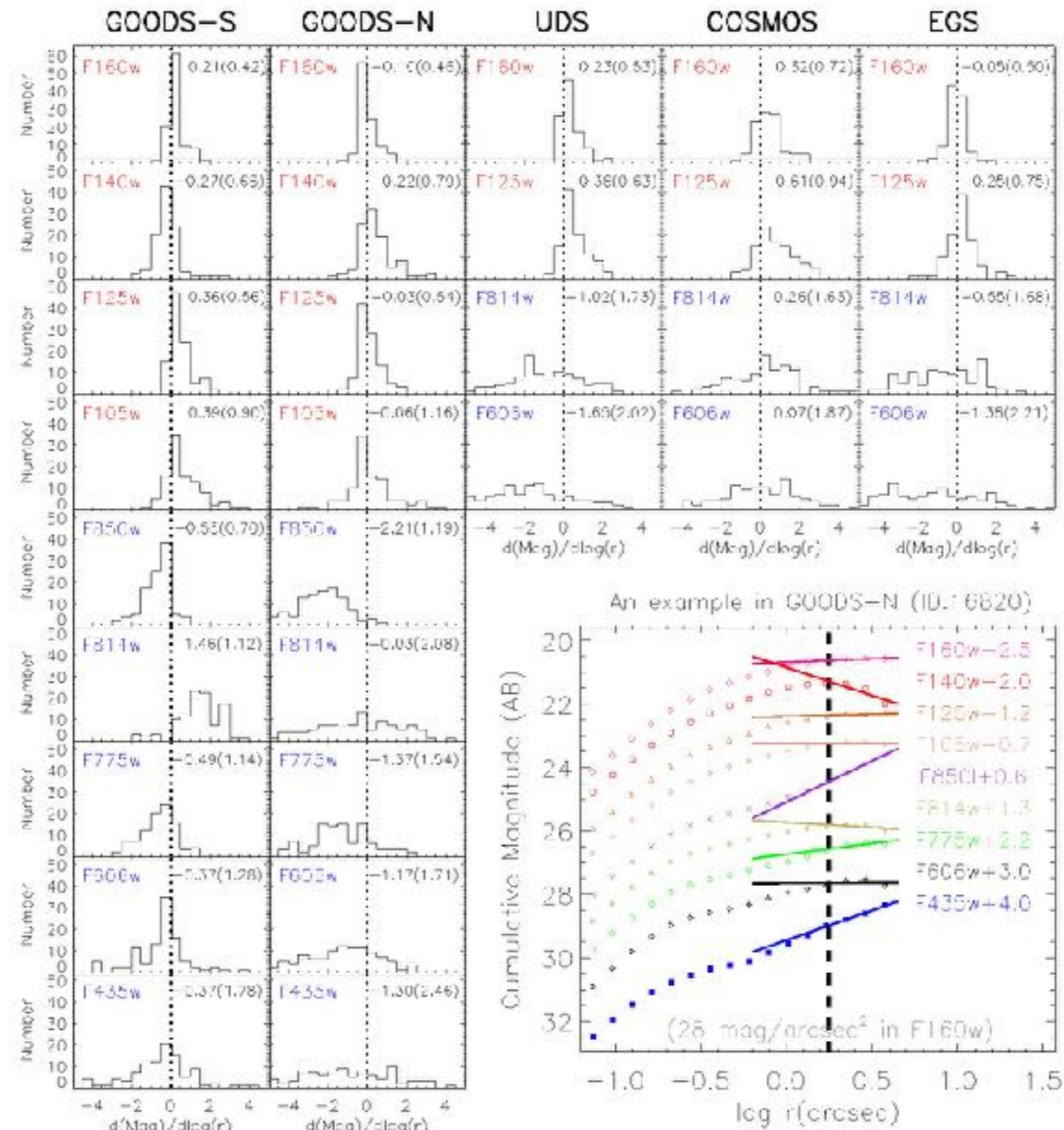


# PSF smearing effect on colours

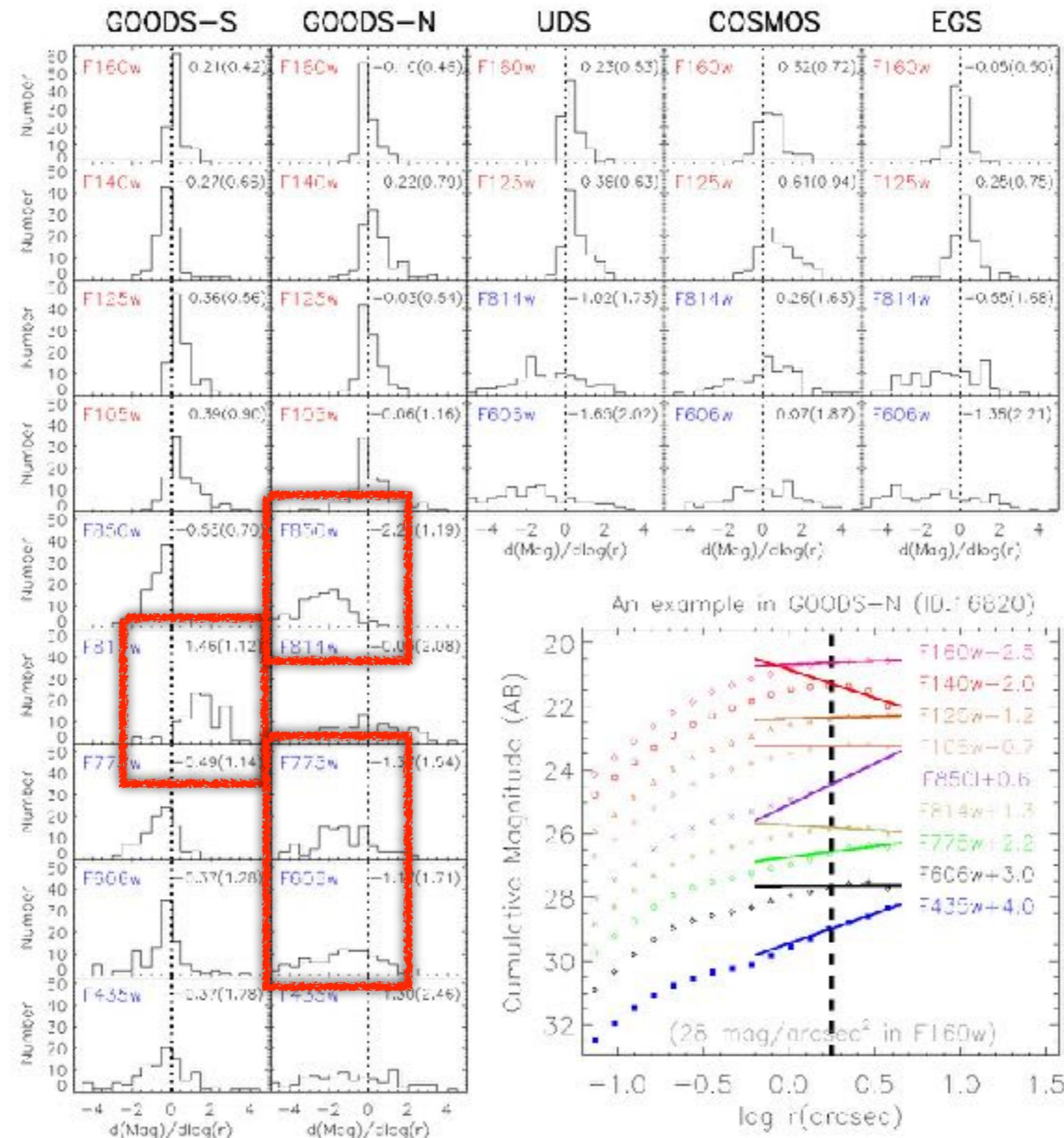
credit :Weichen Wang



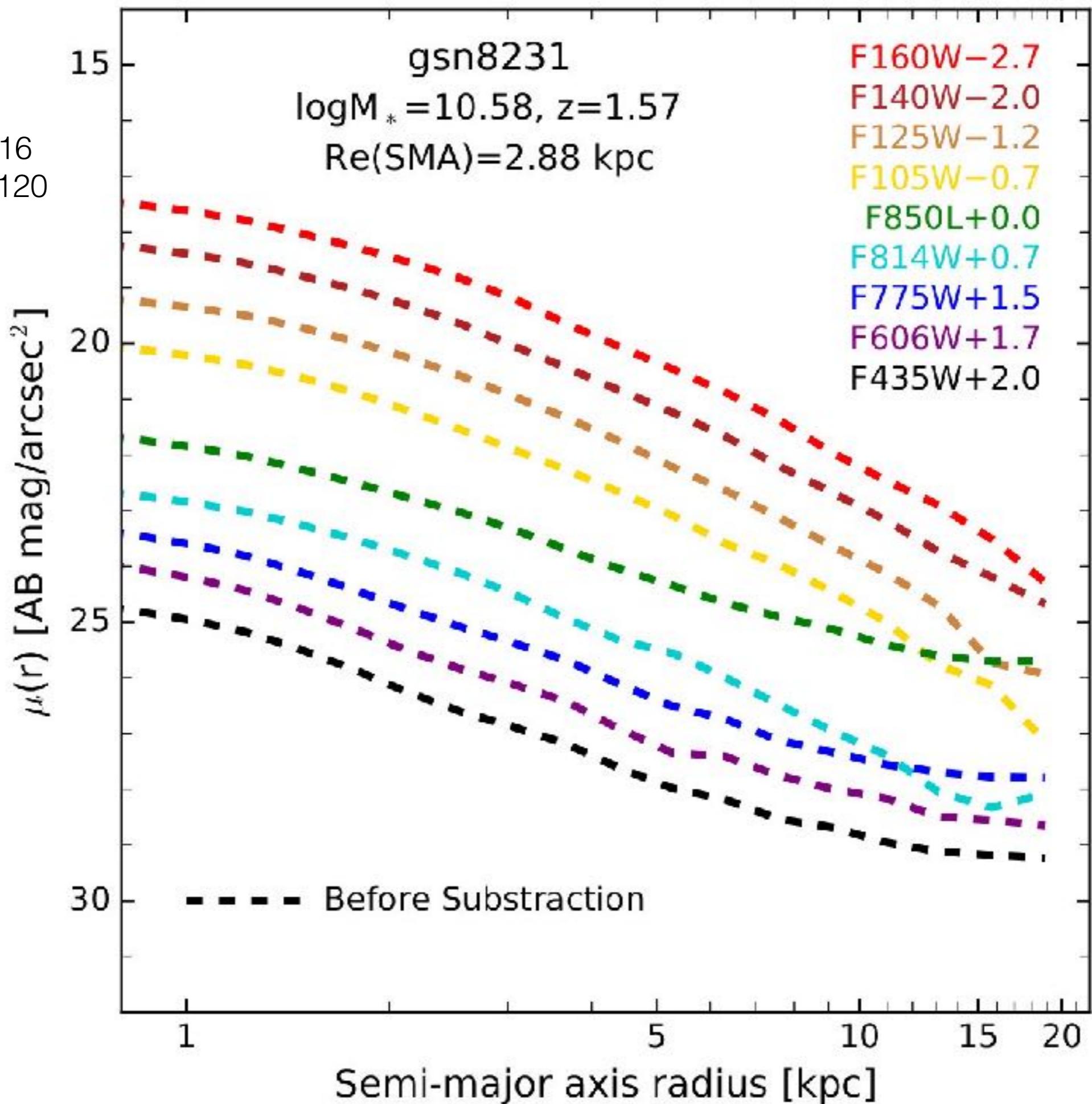
# The uncertain and biased “local background” of CANDELS mosaic imaging



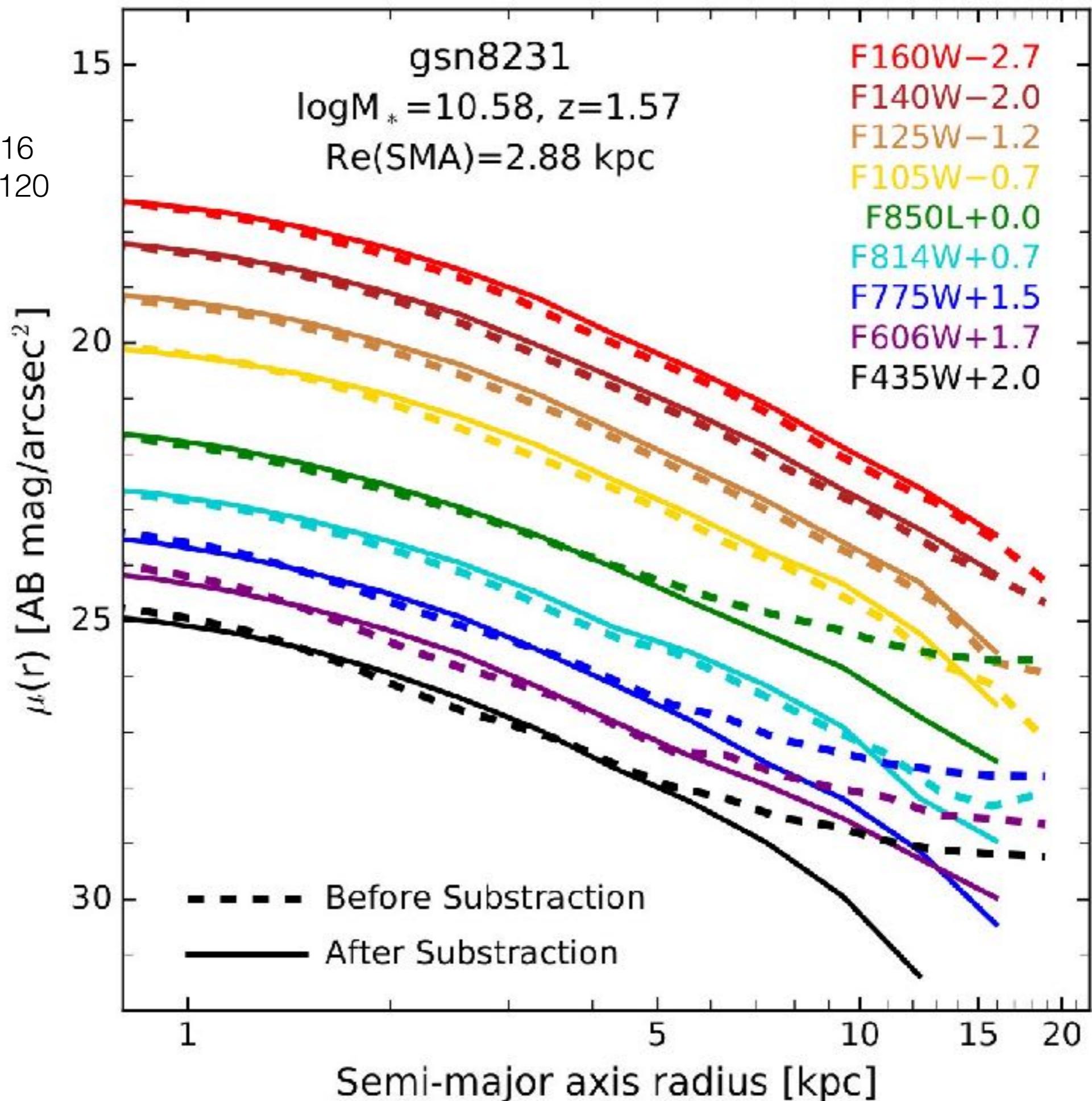
# The uncertain and biased “local background” of CANDELS mosaic imaging



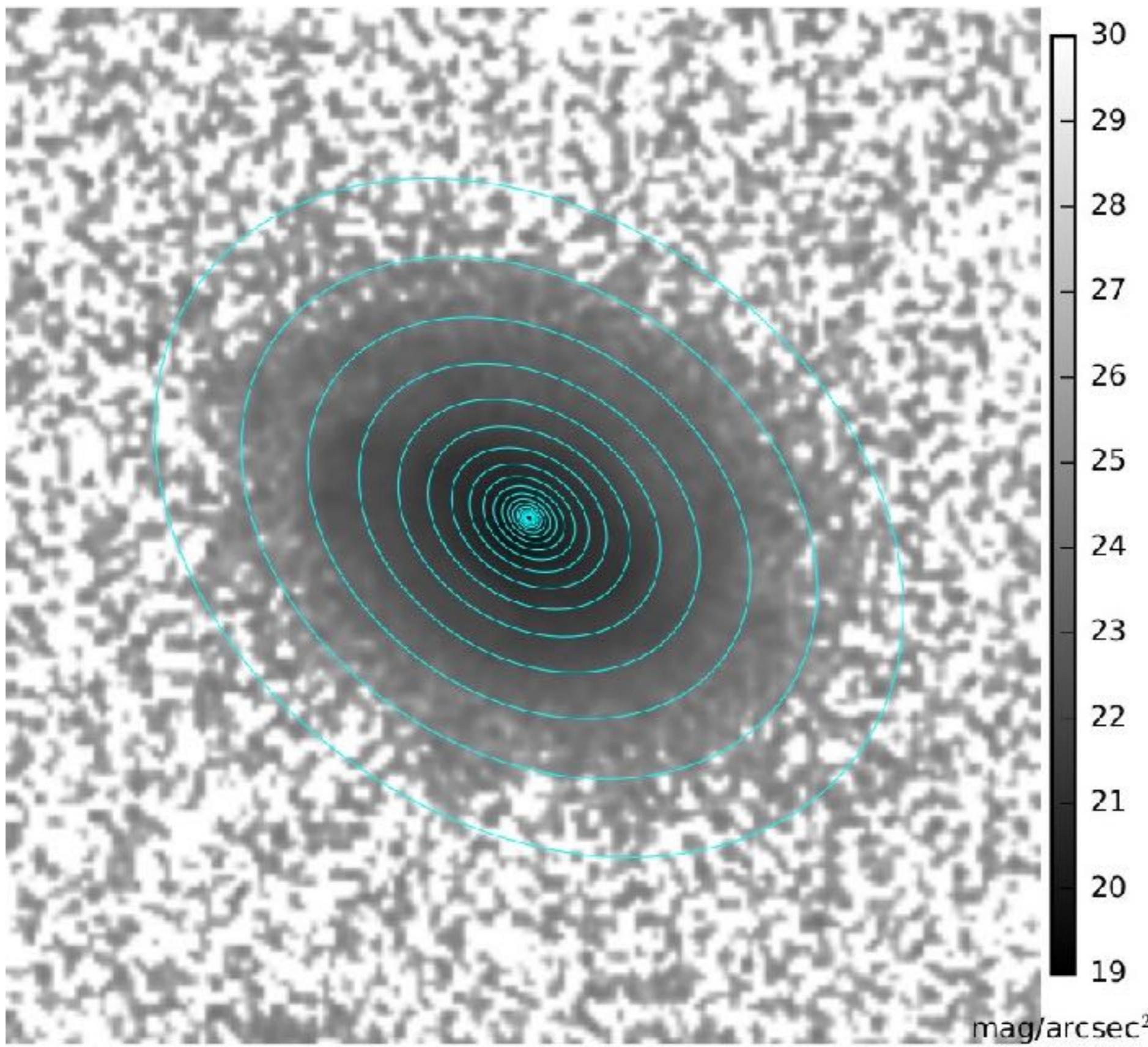
Barro+2016  
ApJ, 820, 120



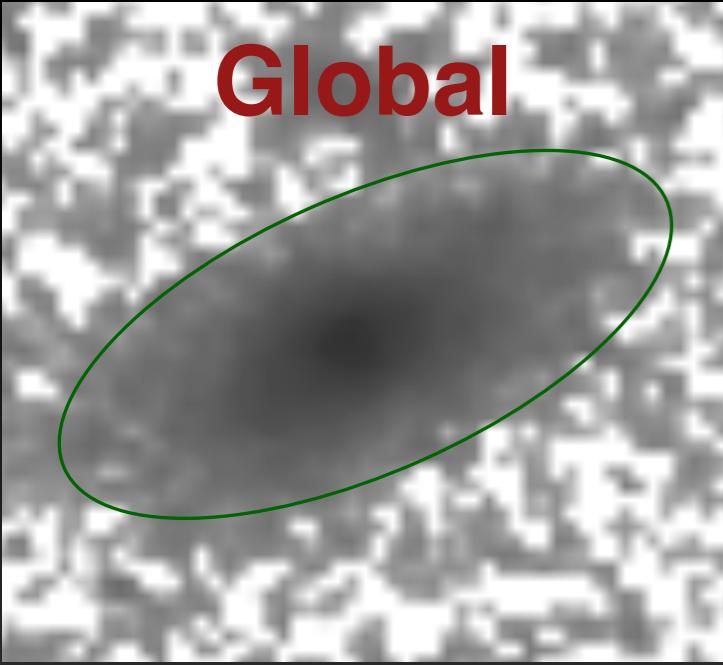
Barro+2016  
ApJ, 820, 120



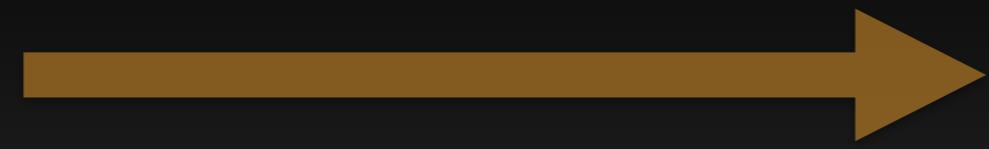
Applying a logarithmic step of 0.3  
(*IRAF/Ellipse*, Jedrzejewski et al. 1987)



**Global**

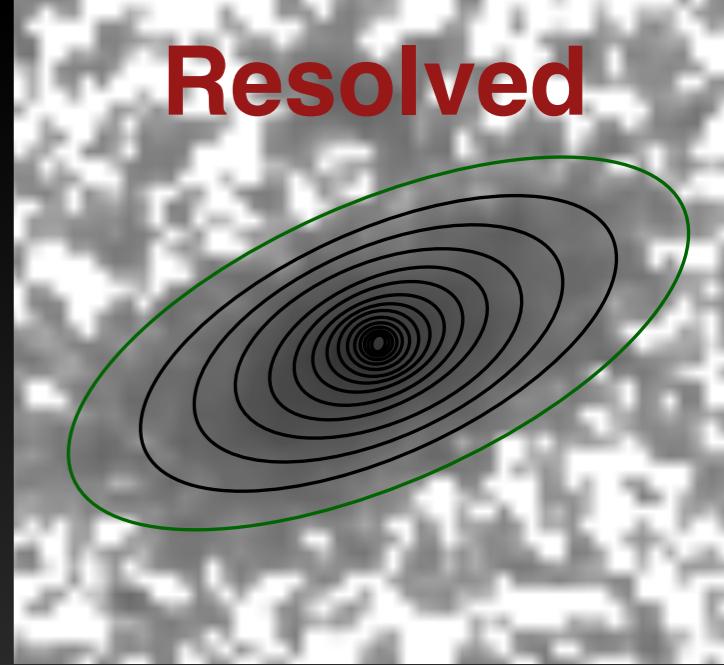


## CANDELS-VAPC: HST Multi-Aperture Photometry Catalogs



(Fengshan Liu, Dongfei Jiang,  
David C. Koo, S. M. Faber,  
Yicheng Guo, and CANDELS Team)

**Resolved**



### Part I : Isophotal Aperture Photometry (Fourier Expansion)

Axis ratio, Boxy/Disky A<sub>4</sub>, PAs,... as a function of radius  
(F160w and F125w)

### Part II: Elliptical Aperture Photometry

(the fixed centre, axis ratio, and PA)

1. Multi-wavelength surface brightness profiles
2. Rest-frame colour profiles from EAZY code
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(GOODS-S & GOODS-N: 5 ACS + 4 WFC3 bands  
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**Shenyang Normal Univ.  
& UC Santa Cruz**

**v0.8 available soon!**  
Contact: [fengshan.liu@yahoo.com](mailto:fengshan.liu@yahoo.com)

# Current catalogs

<b>UDS</b>	wide	<b>v,i<sub>814</sub>,J,JH,H</b>	23642/35932	v0.5
<b>GOODS-S</b>	deep,wide	<b>B,V,i,i<sub>814</sub>,z,Y<sub>105</sub>,J,JH,H</b>	21796/34930	v0.5
<b>COSMOS</b>	wide	<b>v,i<sub>814</sub>,J,JH,H</b>	18183/38671	v0.5
<b>EGS</b>	wide	<b>v,i<sub>814</sub>,J,JH,H</b>	24979/41457	v0.5
<b>GOODS-N</b>	deep,wide	<b>B,V,i,i<sub>814</sub>,z,Y<sub>105</sub>,J,JH,H</b>	23412/35445	v0.5

- CANDELS Value-Added Photometry Catalogs (CANDELS-VAPC): HST Multi-wavelength Multi-aperture Catalogs in all five CANDELS Fields (F. S. Liu et al. in preparation)

Access to v0.5:

<https://www.dropbox.com/sh/ybhfm1qs8afkhwl/AAA9TQQ6u6eZAfVZuzibF9t-a?dl=0>

# What we have done & What we plan to do

1. Jiang Dongfei, Liu F. S., et al. 2017a, submitted to ApJ, "The Isophotal Structure of Star-Forming Galaxies at  $0.5 < z < 1.8$  in CANDELS: Tracing the Buildups of Central Bulges and Outer Stellar Halos"
2. Jiang Dongfei, Liu F. S., et al. 2017b, in preparation, "The Resolved Stellar Population of the Outskirts of Star-Forming Galaxies at  $z \sim 1$ "
3. Jia Men, Liu F. S., et al. 2017, in preparation, "Spatially-resolved Star Formation and Structural Changes from the Blue Cloud to the Red Sequence at  $0.5 < z < 1.4$ "
4. Liu F. S., Jiang Dongfei, Faber S. M., et al. 2017, ApJL, 844, 1, "The Origins of UV-optical Color Gradients in Star-forming Galaxies at  $z \sim 2$ : Predominant Dust Gradients But Negligible sSFR Gradients"
5. Liu F. S., Jiang Dongfei, Guo Y., 2016, ApJL, 822, 25, "The UV–Optical Color Gradients in Star-forming Galaxies at  $0.5 < z < 1.5$ : Origins and Link to Galaxy Assembly"
6. Wang Weichen, Faber S. M., Liu F. S., et al. 2017, MNRAS, 469, 4063, "UVI colour gradients of  $0.4 < z < 1.4$  star-forming main-sequence galaxies in CANDELS: dust extinction and star formation profiles"
7. Barro G., Faber S. M., Koo David C., et al. 2017, ApJ, 840, 47, "Structural and Star-forming Relations since  $z \sim 3$ : Connecting Compact Star-forming and Quiescent Galaxies"
8. Barro G., Faber S. M., Dekel Avishai, et al. 2016, 820, 120, "Caught in the Act: Gas and Stellar Velocity Dispersions in a Fast Quenching Compact Star-Forming Galaxy at  $z \sim 1.7$ "

These spatially-resolved datasets would be potentially a valuable and new resource to study galaxy evolution over cosmic time with the CANDELS, which will be useful for a wide range of science!

*Welcome to use!*