AY1 Announcements

- Quiz 4: Thursday June 8 (last day of class)
- Final: June 14, 4pm -7pm

June 1: Mandel Lecture

Mandel Lecture June 2017

Einstein, Gravitational Waves and Black Holes

Dr. Gabriela Gonzales

Thursday, June 1, 2017

7:30 pm

Rio Theatre - 1205 Soquel Avenue, Santa Cruz

Admission is FREE and open to the general public.

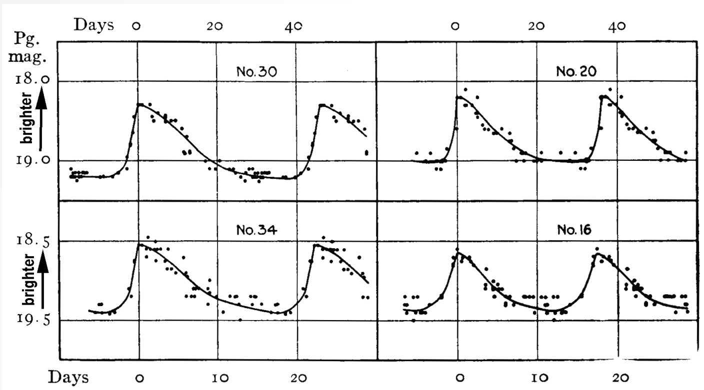
No tickets or reservations will be required for entry. Seating availability is on a first come basis. Theater will be open for seating starting at 7:00 pm.

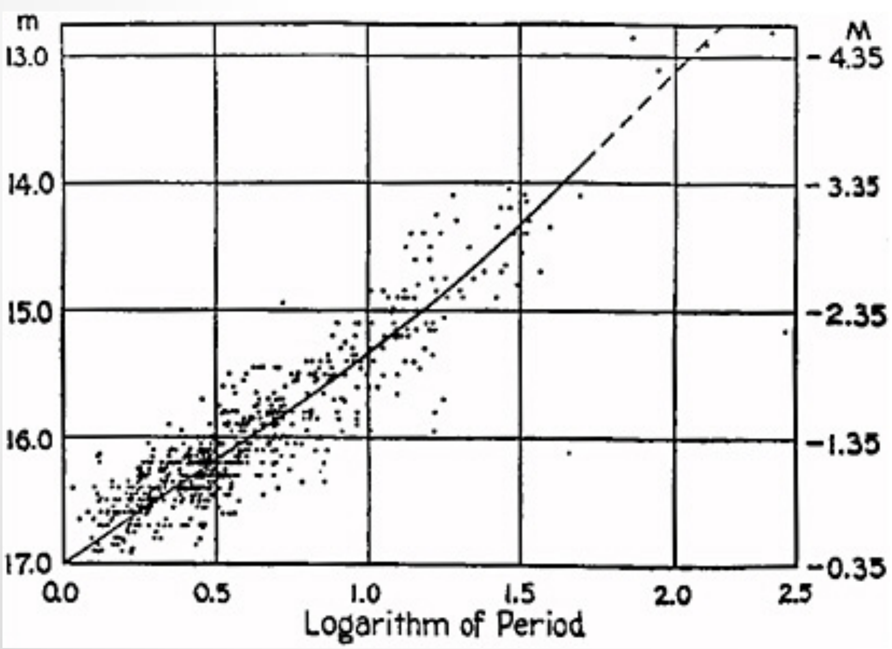
From Last Time

Motions of Galaxies

- 1912-1917 Vesto Slipher at Lowell Observatory made the first Doppler shift measurements of extra-galactic "spiral nebula"
- 22 out of 25 had redshifted spectra, some implying very large motions away from the Earth and Galaxy

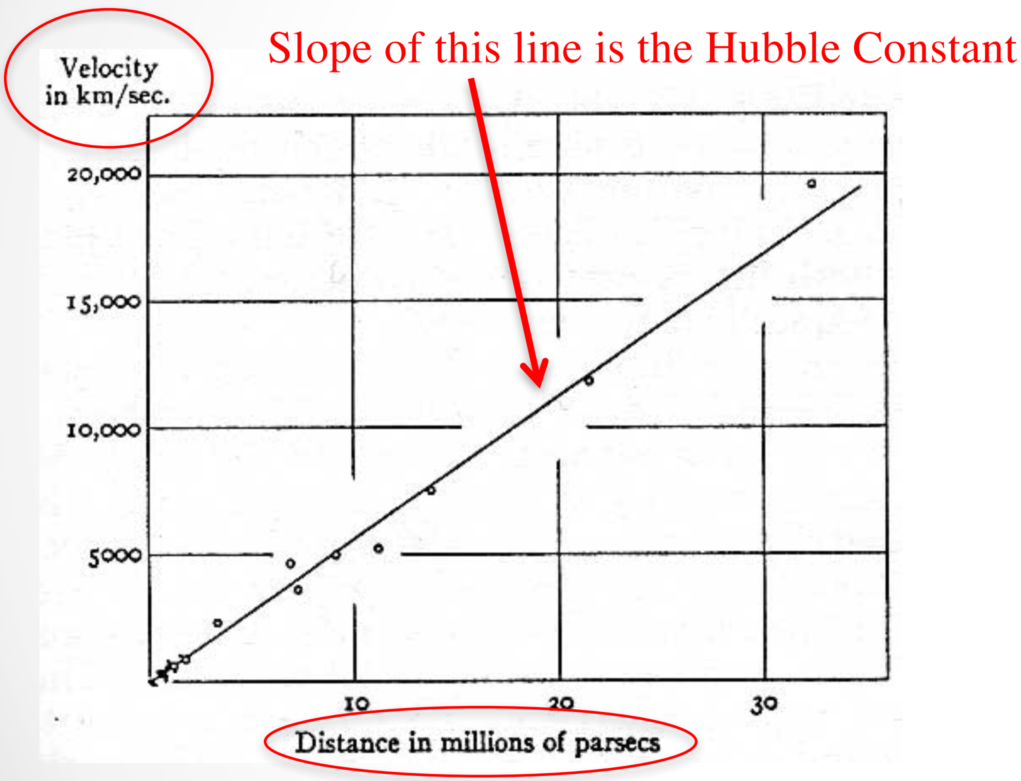
Cepheid Variables





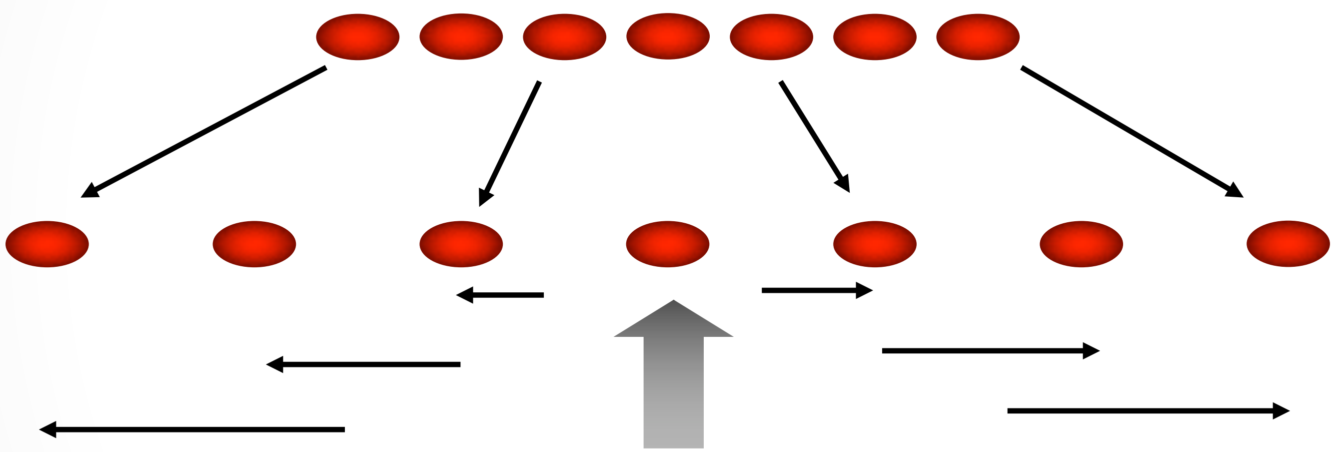
- 1908 Henrieta Swope at Harvard College published a paper noting that for a particular type of star called a Cepheid Variable the period of variability was related to the maximum brightness of the stars

Expanding Universe



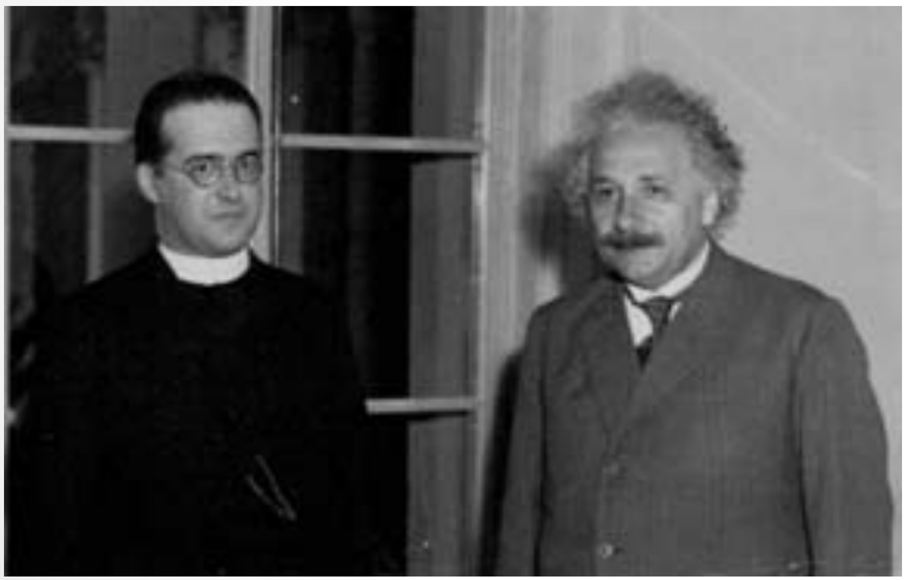
In 1931, Hubble and Humason published a second paper on the distances and velocities of galaxies using additional distance indicators calibrated by the Cepheid work

The Expanding Universe



These people measure their nearest neighbors to have moved one unit, the next galaxies to have moved 2 units…

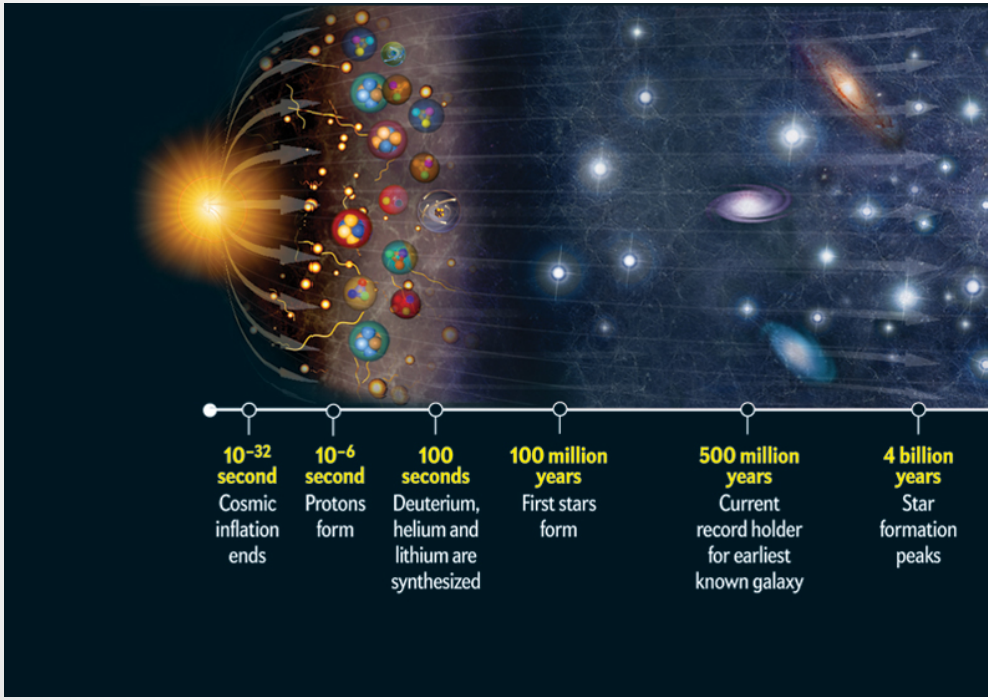
Expanding Universe Implications



- Back to Lemaitre. 1931 he published a paper that suggested the Universe was once contained in a single point, the "primeval atom"
- The expansion of the Universe was not matter expanding into pre-existing space but rather the creation of space and time

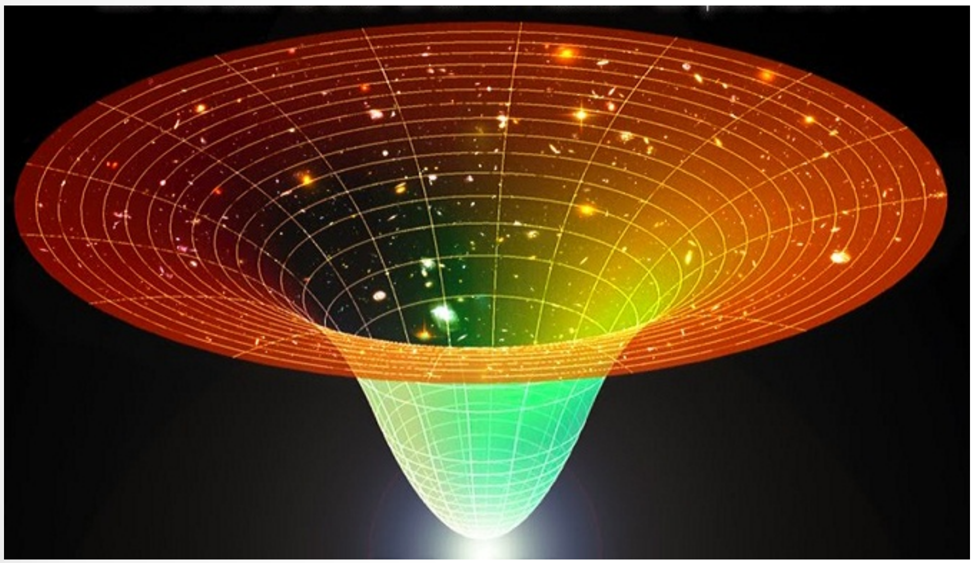
Section 4 Material

The Big Bang thought experiment

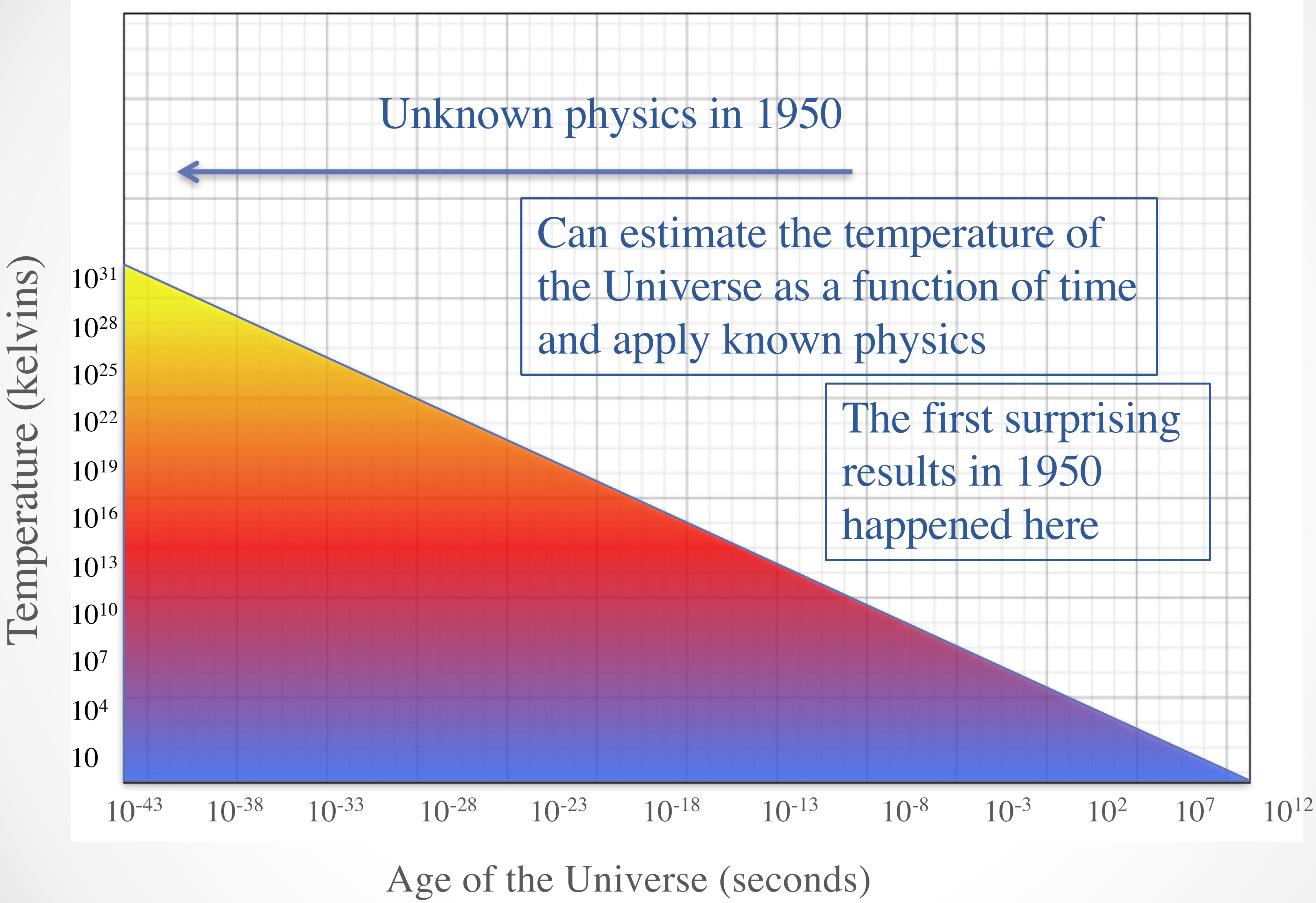


Lemaitre in 1927 discussed the concept of a Universe that started as an extremely dense and hot point and expanded, cooled and evolved into the Universe we live in today

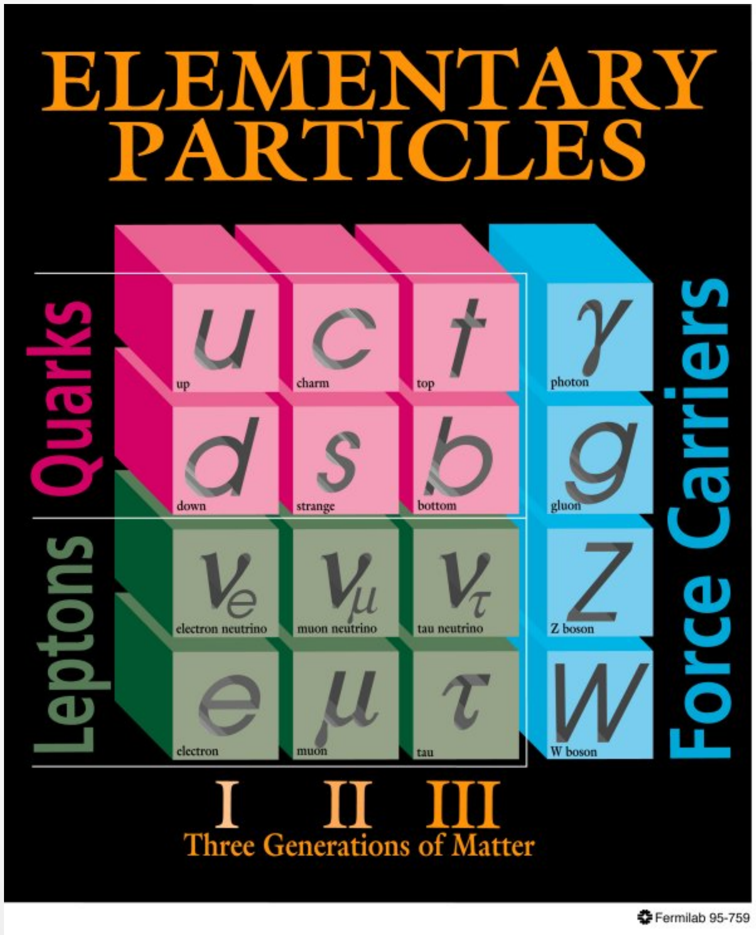
Moving back to the Beginning



- Run the Universe timeline backward
- Density and temperature go up, apply physics as we know it
- Need to think about "phase transitions"
- This first was considered in the 1950's when much of the early Universe physics had still not been developed

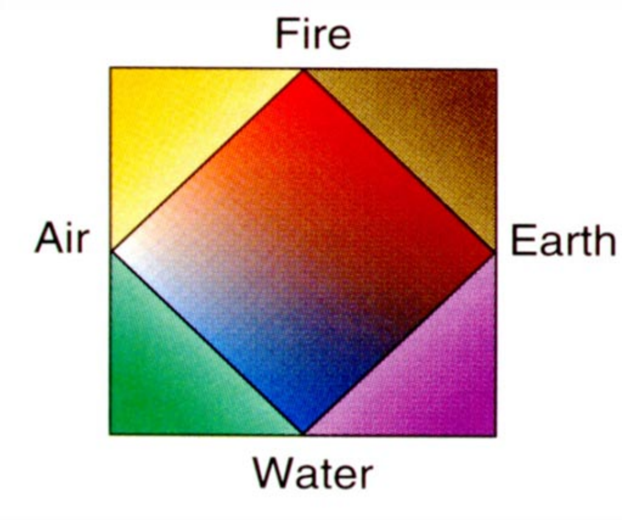

Unknown physics in 1950
Can estimate the temperature of the Universe as a function of time and apply known physics
The first surprising results in 1950 happened here
Temperature (kelvins)
10^31
10^28
10^25
10^22
10^19
10^16
10^13
10^10
10^7
10^4
10
10^-43
10^-38
10^-33
10^-28
10^-23
10^-18
10^-13
10^-8
10^-3
10^2
10^7
10^12
Age of the Universe (seconds)

Background: Fundamental Particles



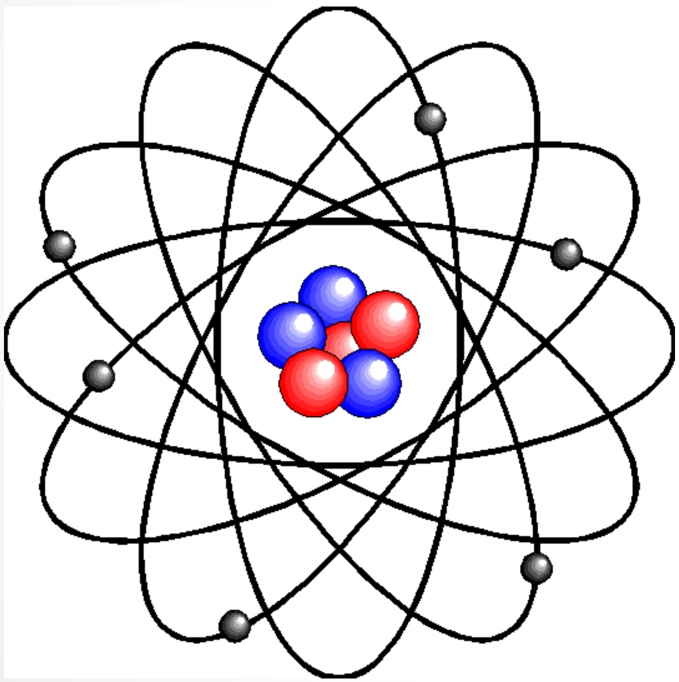
- Identifying the fundamental building blocks of the Universe is a long-time activity of humans
- Like the history of cosmology, there are striking similarities in history and cultures

Earth, Wind, Fire, Water



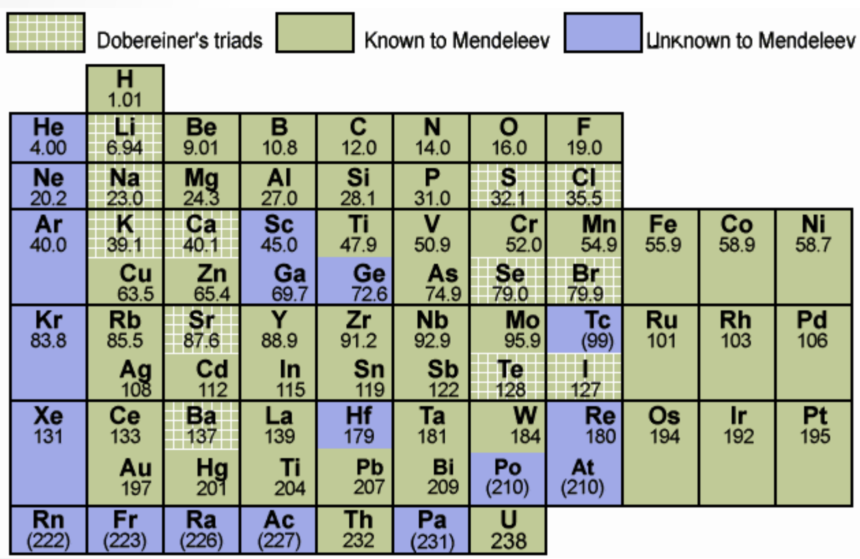
- Greeks: EWFW
- Hindu: Earth, air, fire, water, void (nothing)
- China: Earth, wood, metal, fire, water
- Japan: EWFW, spirit
- Buddhism: EWFW

Atoms



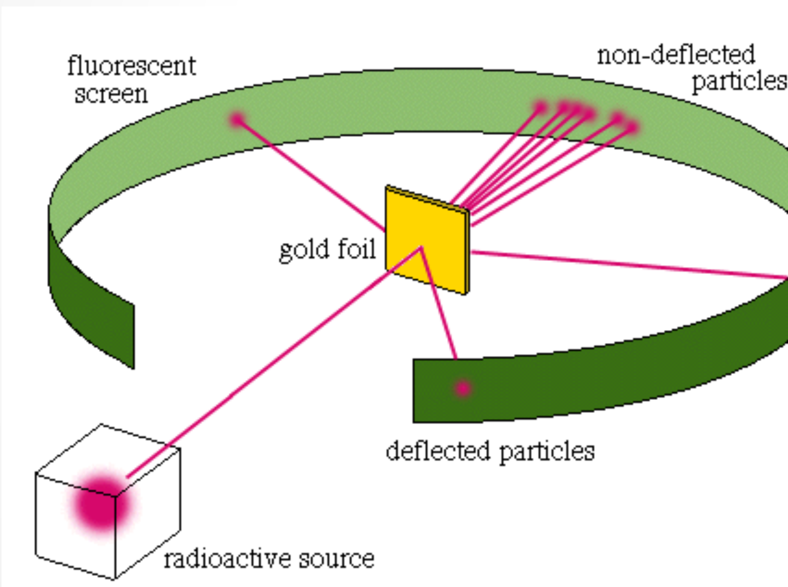
Democritus (greek philosopher and contemporary of Aristotle) considered what would happen if you took matter, divided it in two, then again and again and again, eventually you would have an "atom" that could not be divided further: atomos "not to be cut"

Atoms II: inferences from chemistry



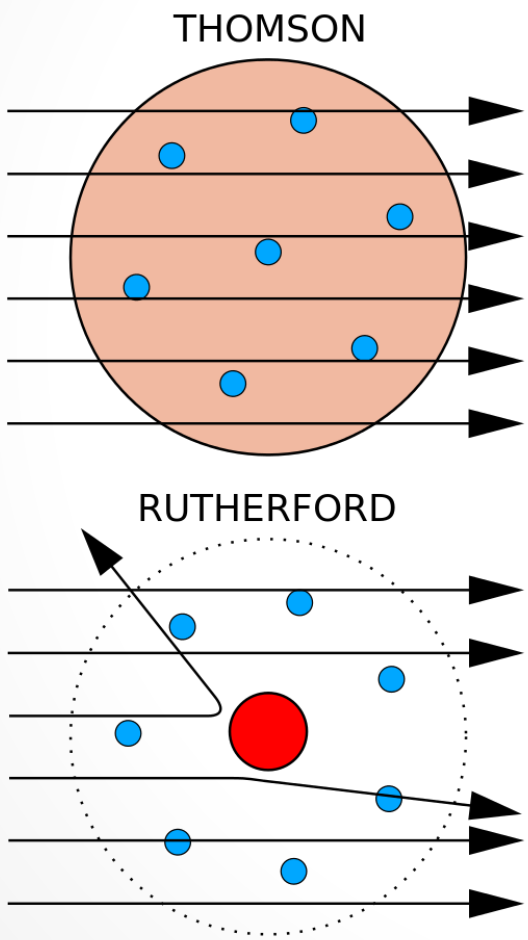
- 1830 John Dalton identified elements with particular types of atom and compounds as combinations of elements
- 1869 Mendeleev proposed the periodic table
- 1897 Thomson discovered electrons and it was realized that atoms were not fundamental but could be further divided

Rutherford and the modern theory



- 1911 Rutherford inferred that atoms had a strongly centralized mass distribution: he discovered the nucleus
- This is all experimentally difficult because of the tiny size of atoms
  - 10^{-10} meters in diameter
  - Human hair is 10^6 atoms across
  - Water drop contains ~10^{21} atoms of oxygen

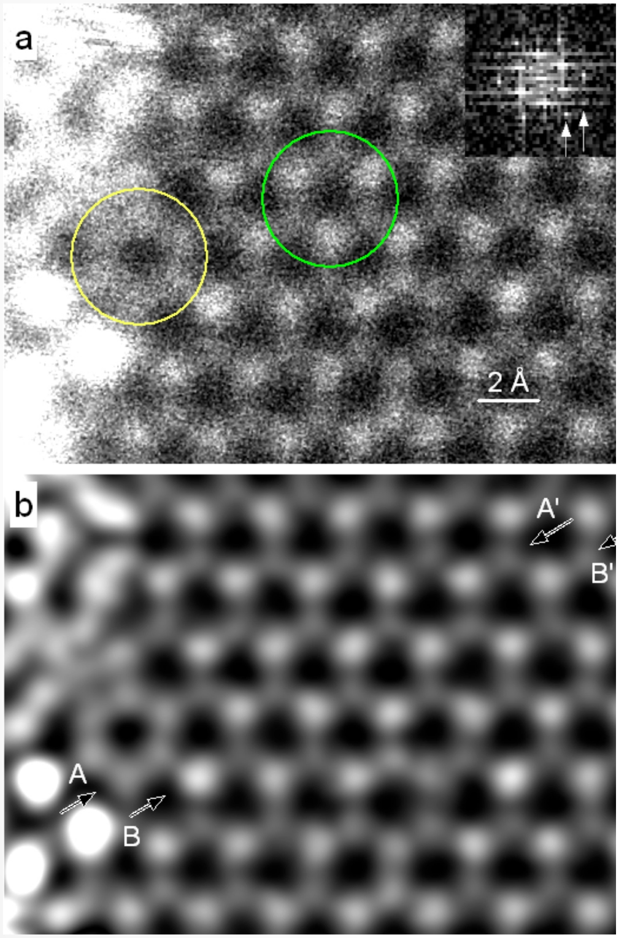
Rutherford Model



Nucleus not to scale!

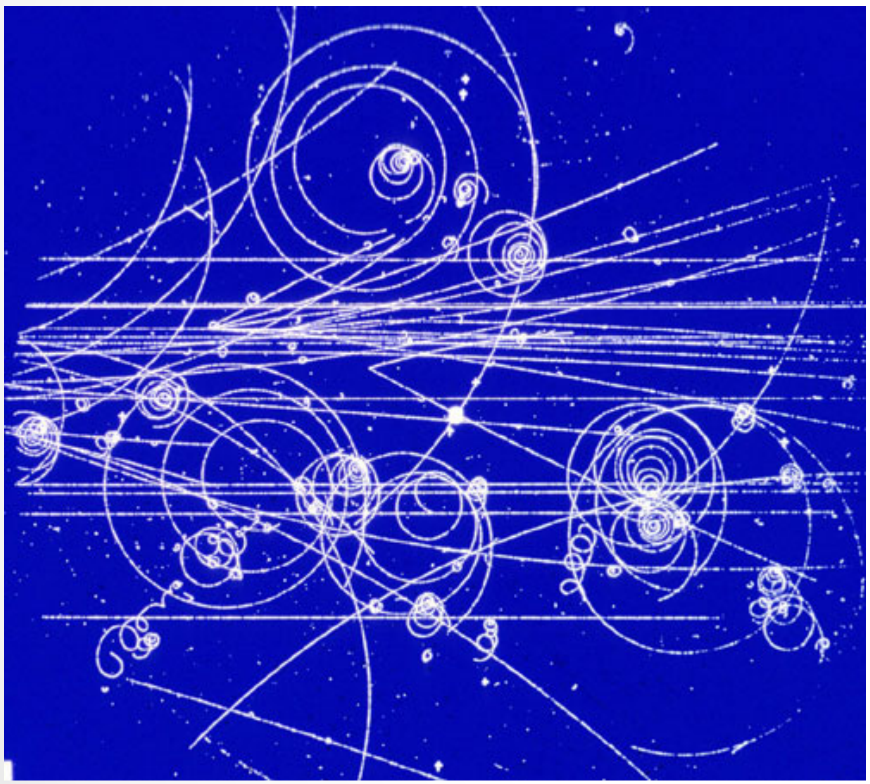
- Before Rutherford, the structure of the atom was thought the be like a “plum pudding” with electrons milling around in a diffuse charged medium
- The large-angle deflections of alpha particles showed there was a dense, charged center of the atom

Quantum Physics

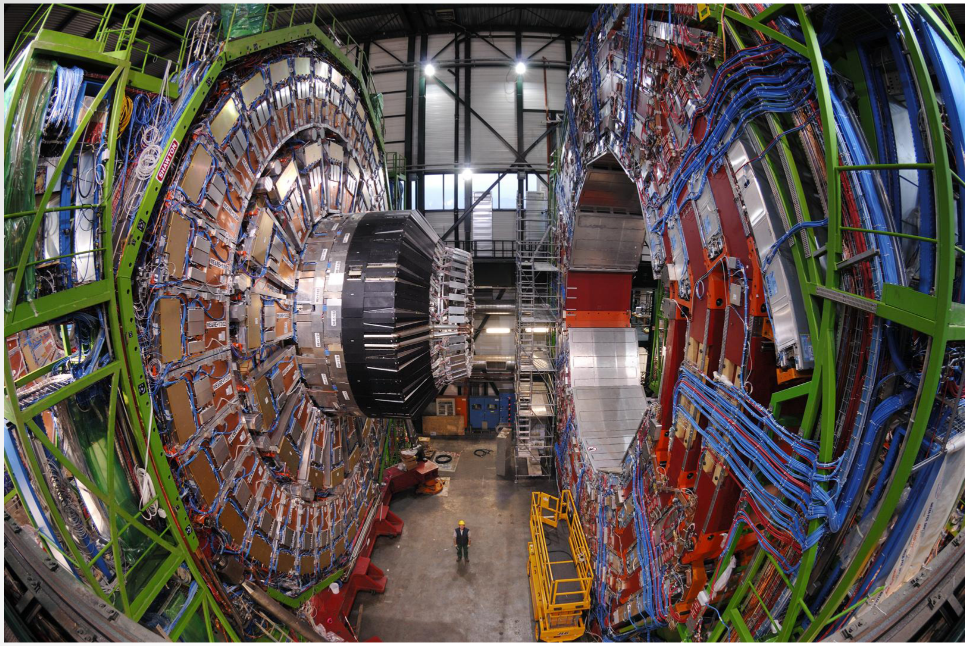


- From 1911 – 1935 our knowledge of the structure of atoms and matter was the subject of intense study "quantum mechanics"
- Fundamental building blocks of everything were thought to be electrons, protons and neutrons

Three Quarks for Muster Mark



- Cosmic ray studies and particle accelerators had been demonstrating a whole slew of unexpected massive particles coming out of energetic collisions
- Particle physics theorists proposed all could be explained by three (then four) truly fundamental particles that got names "quarks" (1964)



- Modern accelerators are incredible machines
- Large Hadron Collider accelerates protons to 0.9999991c
- Super-cooled, vacuum of space
- 10,000 scientists, 100 countries
- 14 x 10^{12} eV (TeV)collisions
- 80,000 computers on private internet
- $6.5B (US)

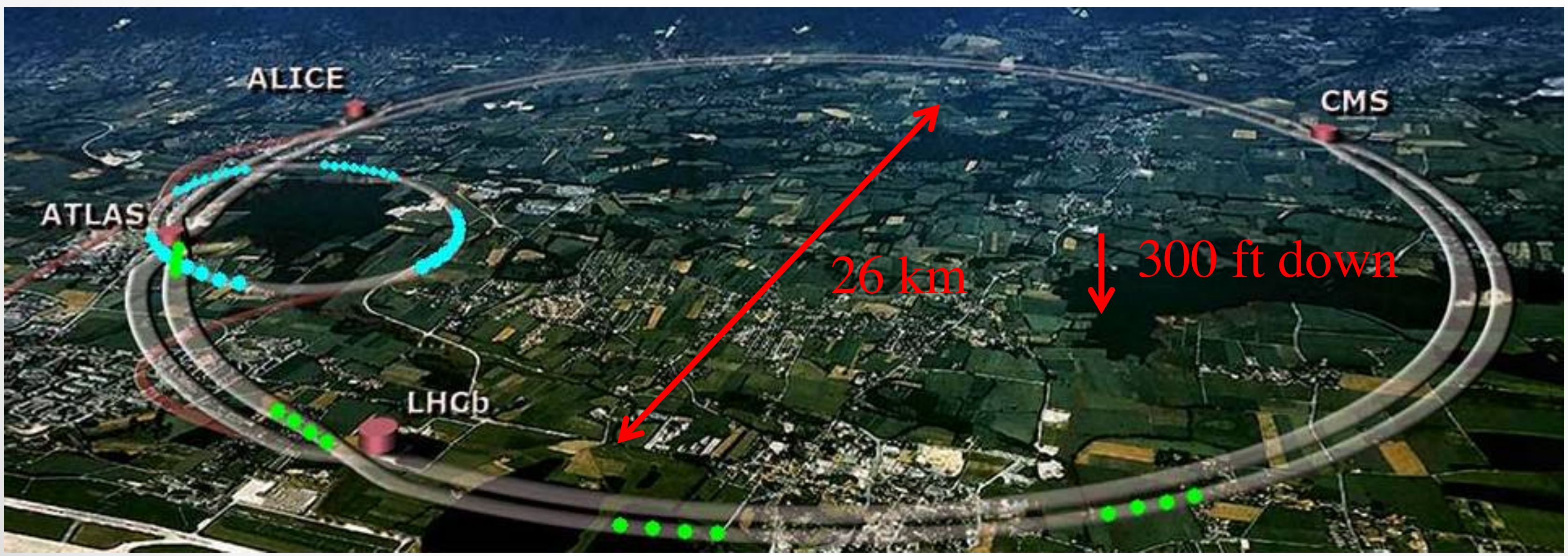
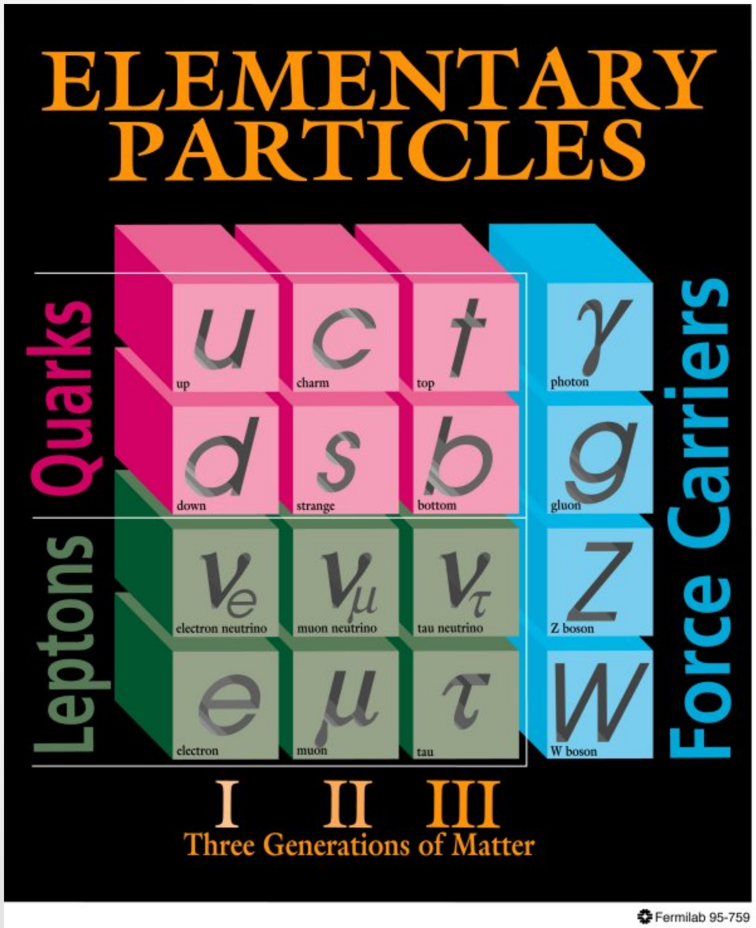
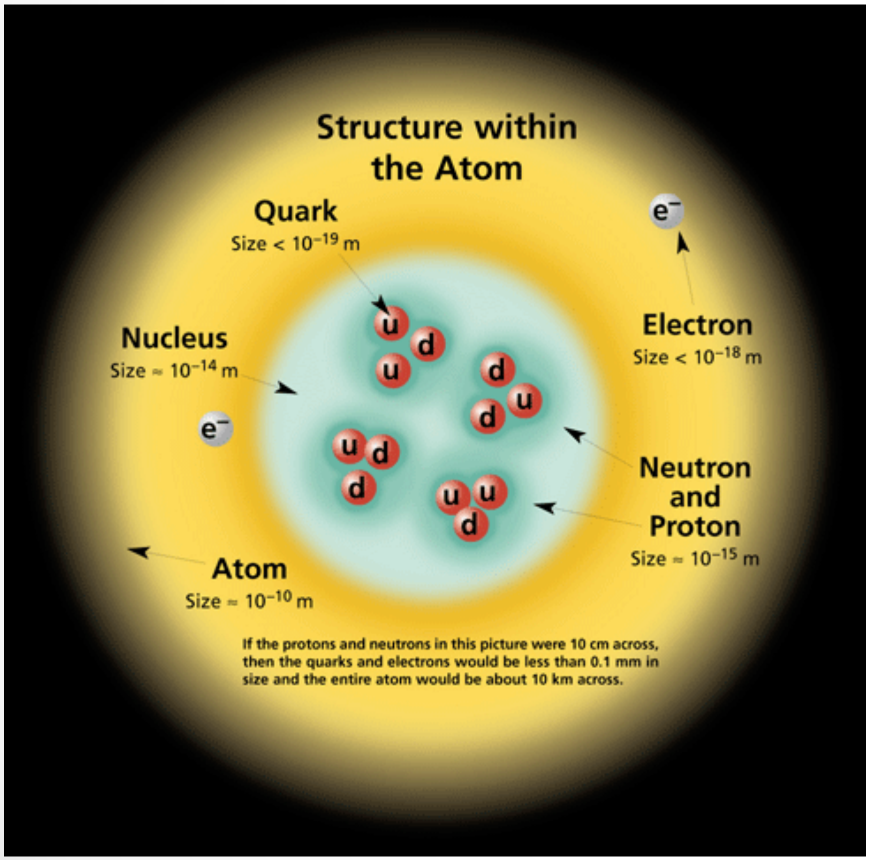


Table of Fundamental Building Blocks



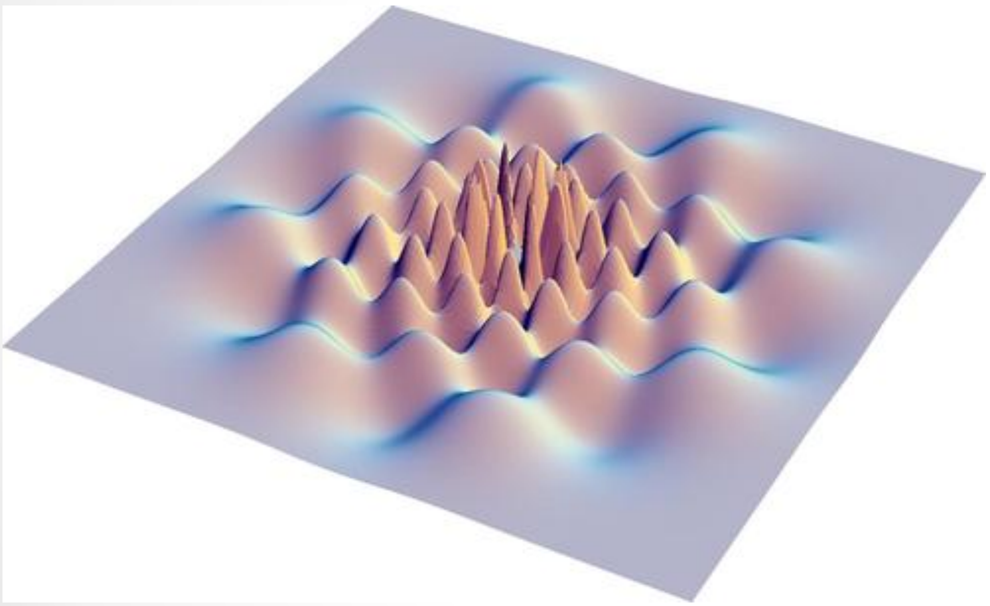
- Fermions: fundamental particles.
  - Leptons
  - Quarks
- Hadrons: combinations of Fermions (i.e. proton, neutron)
- Bosons: particles that exchange forces

Quarks



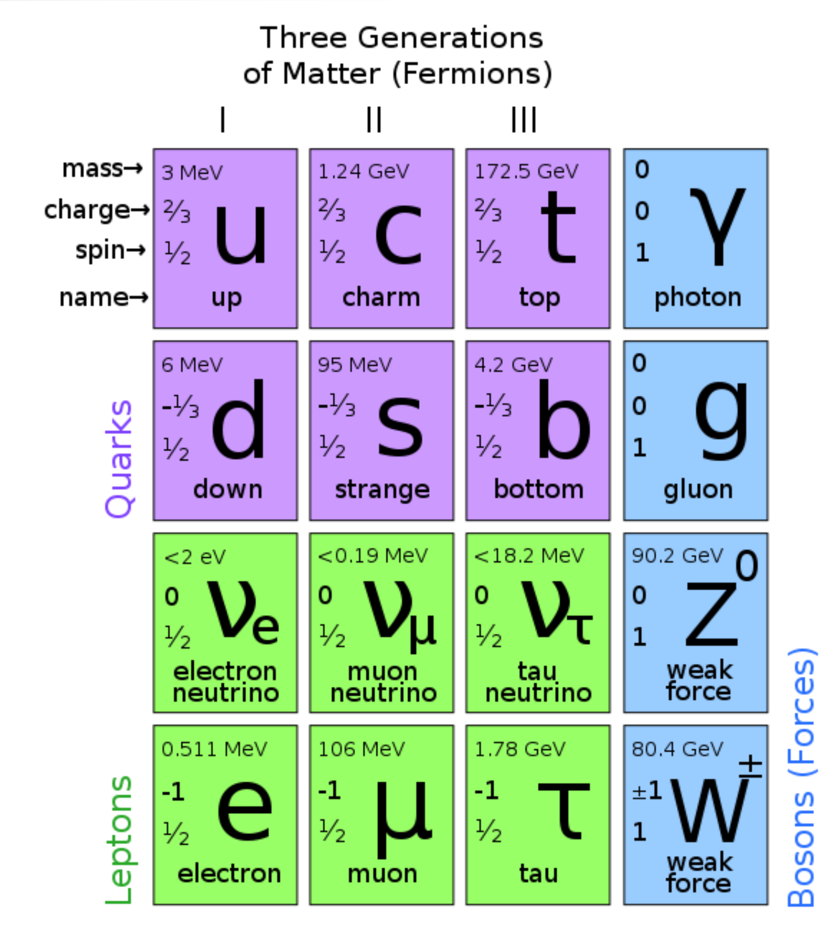
- Postulating the existence of quarks allowed physicists to understand the zoo of new particles
- Six in total with the two lowest mass (up and down) the most stable
- Always confined in nuclei
- In 1968 at the Stanford Linear Accelerator first observational evidence was gathered
- https://en.wikipedia.org/wiki/List_of_baryons

Leptons



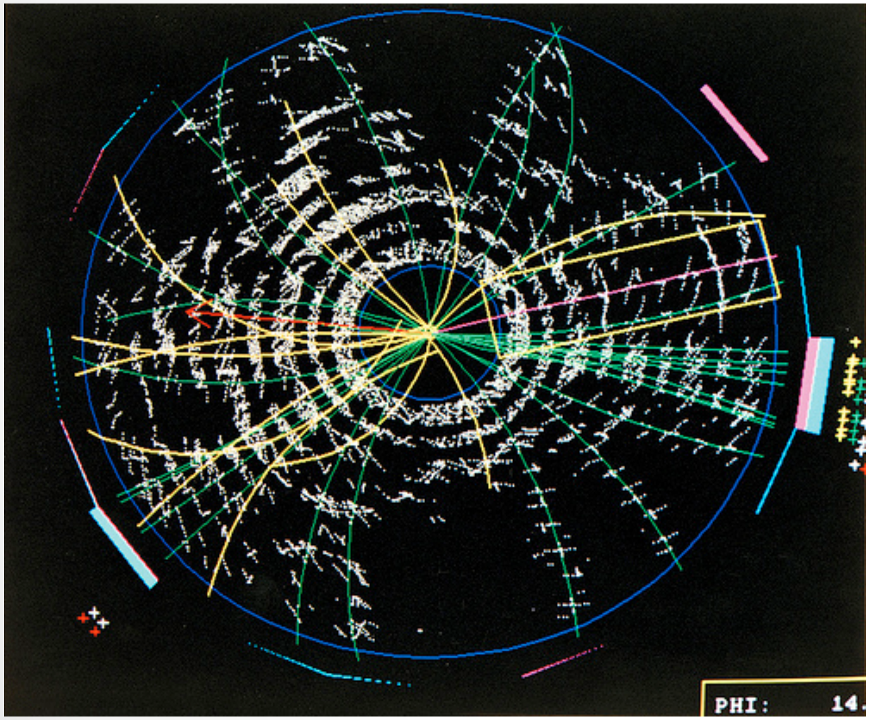
- Near massless, no structure, unknown, but very, very small size
  - Electron, electron neutrino
  - Muon, muon neutrino
  - Tau, Tau neutrino
  - Anti-matter partners

Fermions: Quarks



- Proton: 2u + 1d quarks (uud)
  - Charge: [(2 x 2/3) + (-1/3)]=+1
  - Held together by 3 Gluons
  - Mass=1.672×10^{-27}kg=938.3Mev/c^2
- Neutron: 1u + 2d quarks (udd)
  - Charge: [2/3 + 2 x (-1/3)]=0
  - Held together by 3 Gluons
  - Mass=1.675×10^{-27}kg=939.6Mev/c^2
  - "free" neutrons are unstable and decay via βdecay to become protons with a half-life of 15 minutes

Why do we believe this?



- All this is part of the "Standard Model" for particle physics
- Model calculations (very complex) get made for the production of particles at different energies and the last 40 years of accelerator experiments have shown the models to be remarkably accurate

Energy and Mass

- Common to give particle mass in electron volts (eV), e.g. proton mass=938.3 MeV (should have $/c^2$)
  - Useful for understanding the conditions in which massive particles are created (in accelerators or in the Big Bang)
  - Useful for reminding us about $E=mc^2$
  - Useful for reminding us about binding energy
  - 1eV=1.602 x 10^{-19} Joules

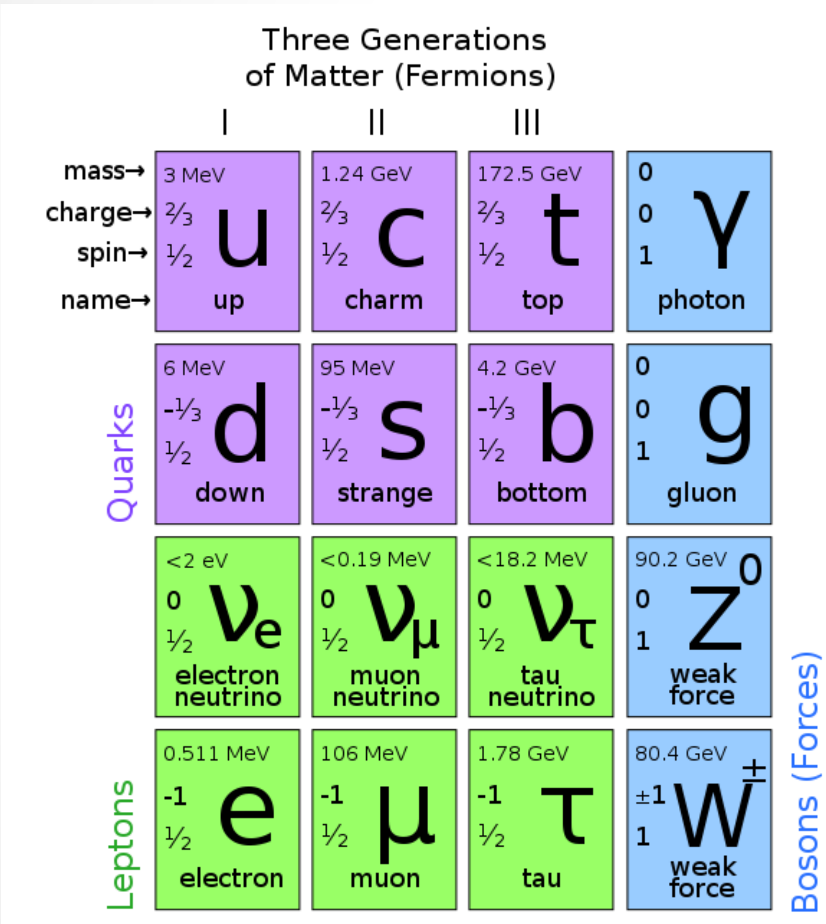
Example: Proton mass = 1.672 x 10^{-27} kg

$E=mc^2=(1.672 \times 10^{-27}) \times (3 \times 10^8)^2 = 1.5 \times 10^{-10}$ J

1.5×10^{-10} J x (1eV)/(1.602 x 10^{-19} J)

=938.3 x 10^6 eV=938.3 MeV (energy equivalent of mass)

An Aside about Proton Mass



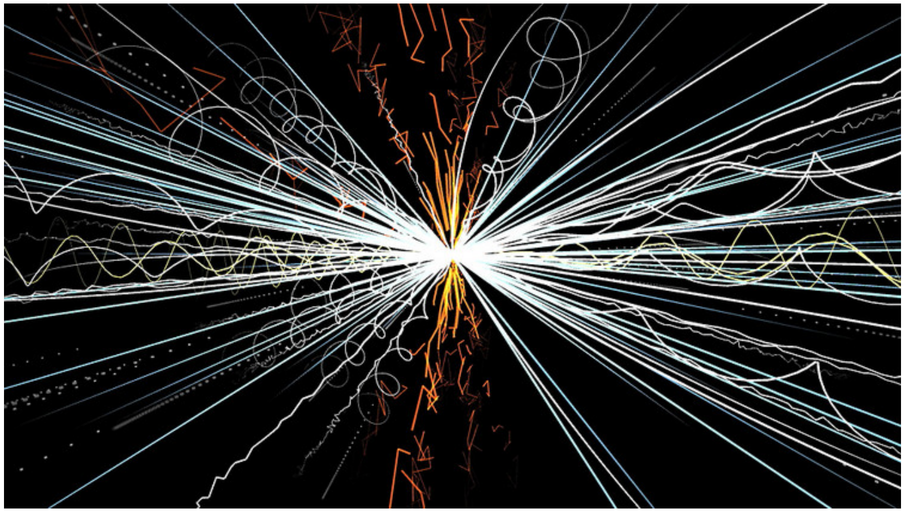
- Proton mass is 938.3 MeV/c^2
  - Up quark mass is 3 MeV/c^2
  - Down quark mass is 6 MeV/c^2
  - Gluon mass is zero
- So, components of proton (uud) only add up to 12 MeV/c^2. The remainder of the rest mass is in the binding energy bonds and some kinetic energy of the quarks in the proton

iclicker

Which of the following are fundamental particles?

A. Atom
B. Electron
C. Proton
D. Allof the above

Creation of Matter

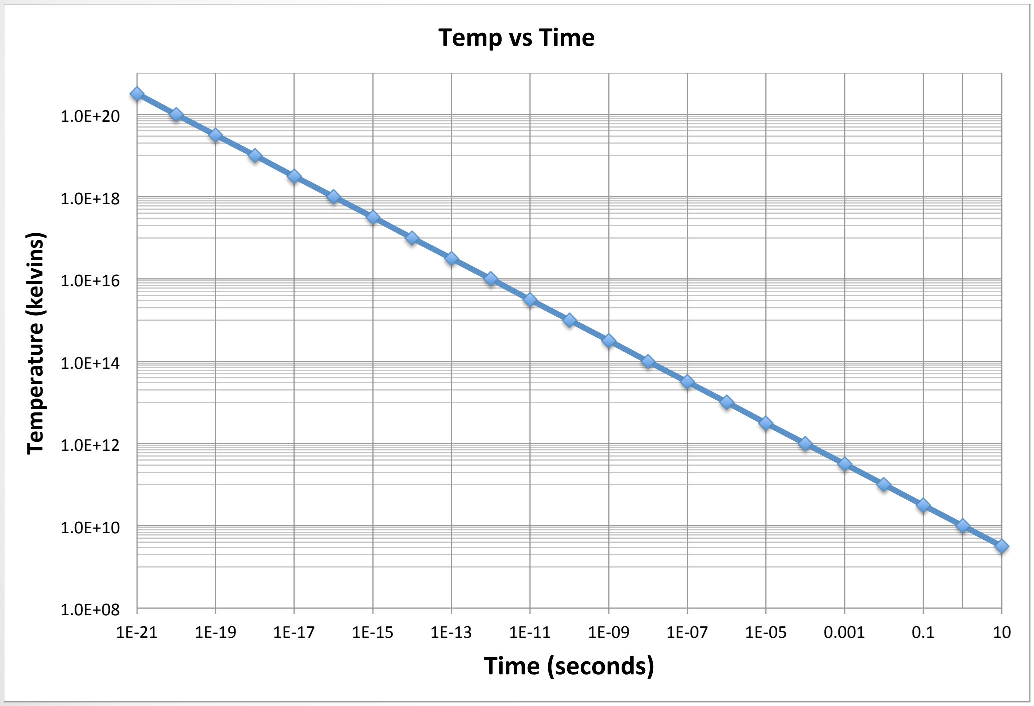


- The threshold energy for creating a particle is its "rest mass energy" given by $E=mc^2$
- In an accelerator, as the beam energy goes up, so does the mass of the particles created in the collisions
- In the early universe, as the temperature goes up so does the mass of particles created

Creation of Matter

- At very high temperatures, radiation is more stable than matter. Very massive matter/anti-matter particle pairs get created and destroyed (broken back into radiation or quarks) regularly: Quark-Gluon Plasma (quark soup)
- The formation of matter and its ability to survive subsequent energetic encounters and native stability is what determines the matter content of the Universe today

Putting these ideas to work



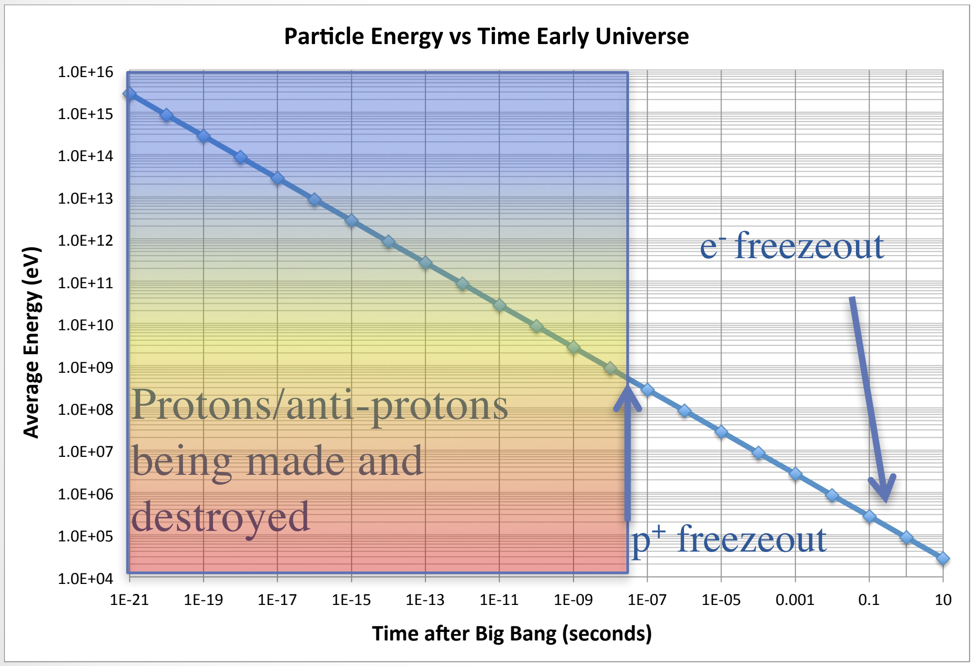
- The temperature of the Universe after the Big Bang is given by:

$$T(\text{kelvin}) = \frac{10^{10}}{\sqrt{t\ (\text{seconds})}}$$

- The average energy of a particle in the gas is given by:

$$E(\text{eV}) = \frac{8.6 \times 10^{5}}{\sqrt{t(\text{seconds})}}$$

Formation of protons

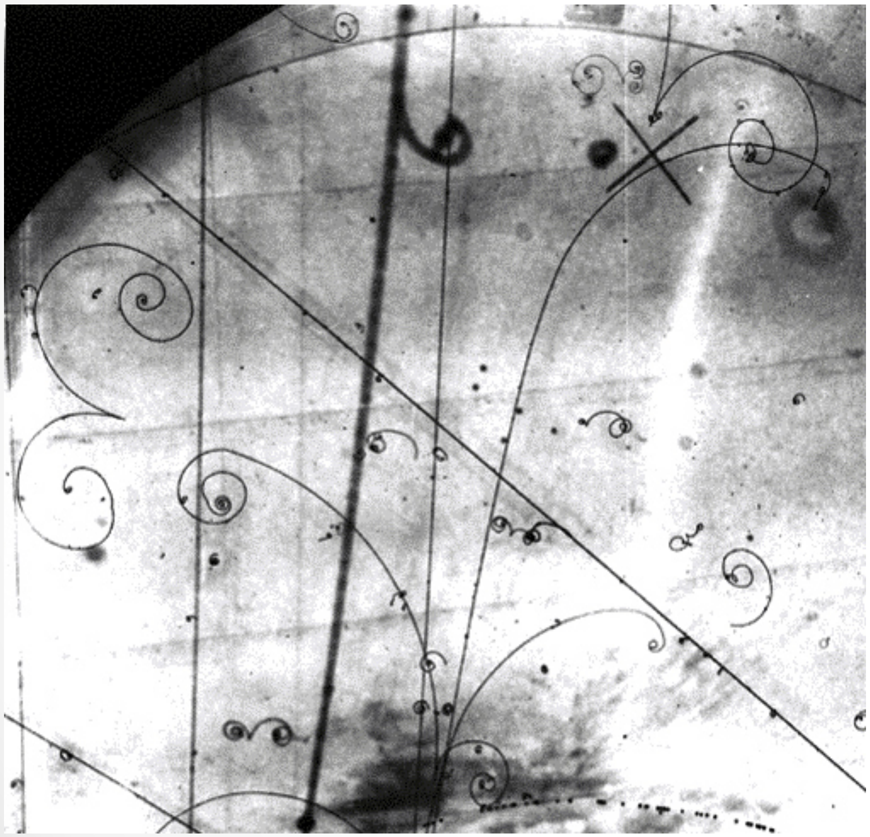


- When is the last time after the Big Bang that proton-antiprotons can be formed?
  - Mass equivalent is twice 938.3 MeV: ~2 x 10^9 eV
  - $E(eV)=(8.6x10^5)/sqrt(t)$
  - $t=(860000)^2/(2 \times 10^9)^2$

    $=1.84 \times 10^{-7}$ seconds

Note: only true in a statistical mean sense

A Really Good Question



- Every particle was produced with an anti-particle
- Yet, we live in a matter (not anti-matter) Universe
- Somehow for every billion anti-particles formed, there were a billion and one particles
- Why and how?

- How close to the Big Bang have we probed with the LHC accelerator if it has achieved beam energies of 1.5 TeV (1.5×10^{12} eV)?

$$E(\mathrm{eV}) = \frac{8.6 \times 10^{5}}{\sqrt{t(\mathrm{seconds})}}$$

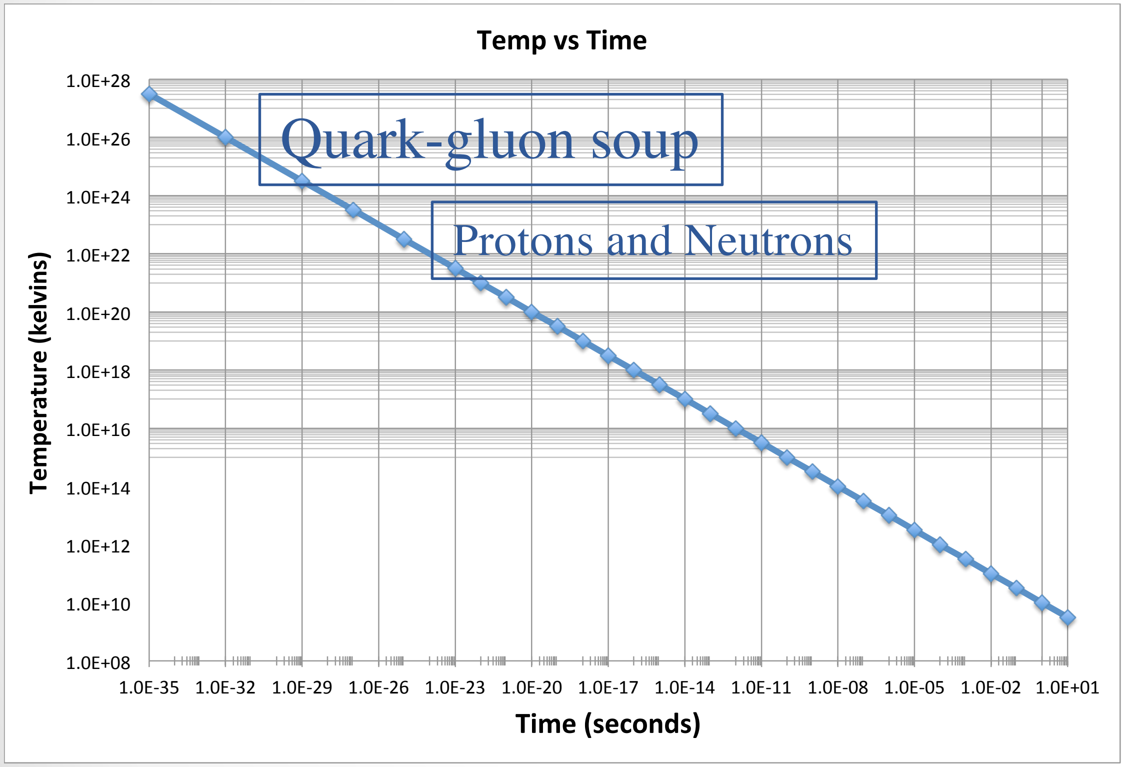
$$t = \frac{(8.6 \times 10^{5})^{2}}{E^{2}} = \frac{7.4 \times 10^{11}}{E^{2}}(\mathrm{sec}) = \frac{7.4 \times 10^{11}}{(1.5 \times 10^{12})^{2}}$$

$$= \frac{7.4 \times 10^{11}}{1.5^{2} \times 10^{24}} = \frac{7.4}{1.5^{2}} \times 10^{(11-24)} = \frac{7.4}{1.5^{2}} \times 10^{(-13)}\mathrm{s}\ (\mathrm{C})$$

Hot Big Bang Nucleosynthesis

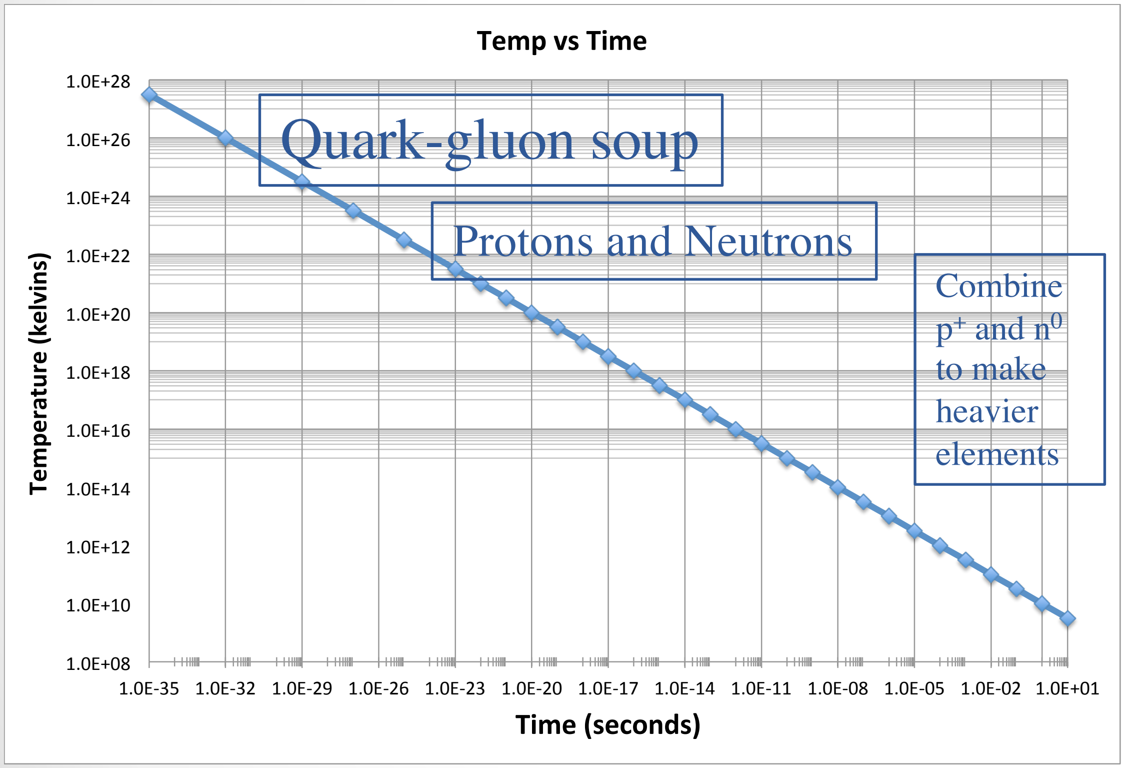
- George Gamov was a nuclear physicist political refugee from Russia who looked in detail at the early times in an expanding Universe run backward
- Made the first predictions of the era of nucleosynthesis and got a big surprise

Early Universe



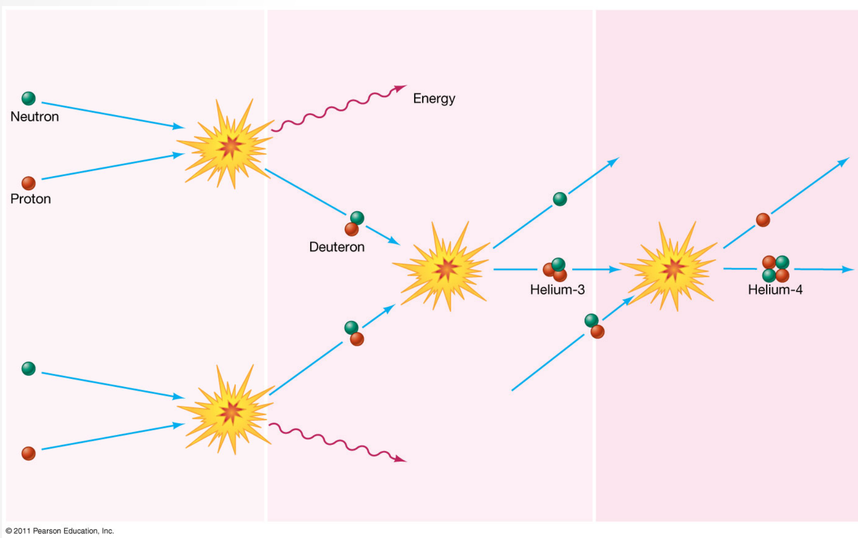
- Have all the background in place to finally get to the Gamow HBB nucleosynthesis predictions
- At very early times, quark soup produces hadrons, they get broken apart, till the temperature is low enough that p^+ and n^0 can survive
- Neutrons/protons/electrons and positrons combine and decay to produce an equilibrium ration of p^+ to n^0

Early Universe



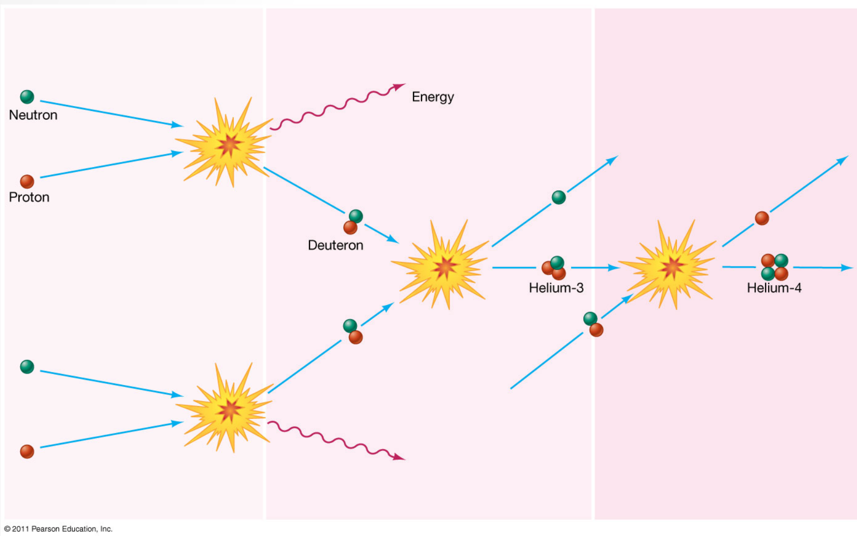
- After 10^{-7}seconds, stop forming p^+ and n^0 and have confined the quarks in these nucleons
- At around 1 second after the BB, the neutron-proton-e^--e^+ reaction stops and the building blocks for "primordial nucleosynthesis" are in place
- Note! Free neutrons are unstable to decay with a half life of 10 minutes which is the window to produce the heavier elements

Primordial Nucleosynthesis



- Gamov and Alpher started with the conditions at a few tenth's of a second after the BB
- Need: temperature, density, # of protons and neutrons, *nuclear physics cross-sections for reactions*

Primordial Nucleosynthesis



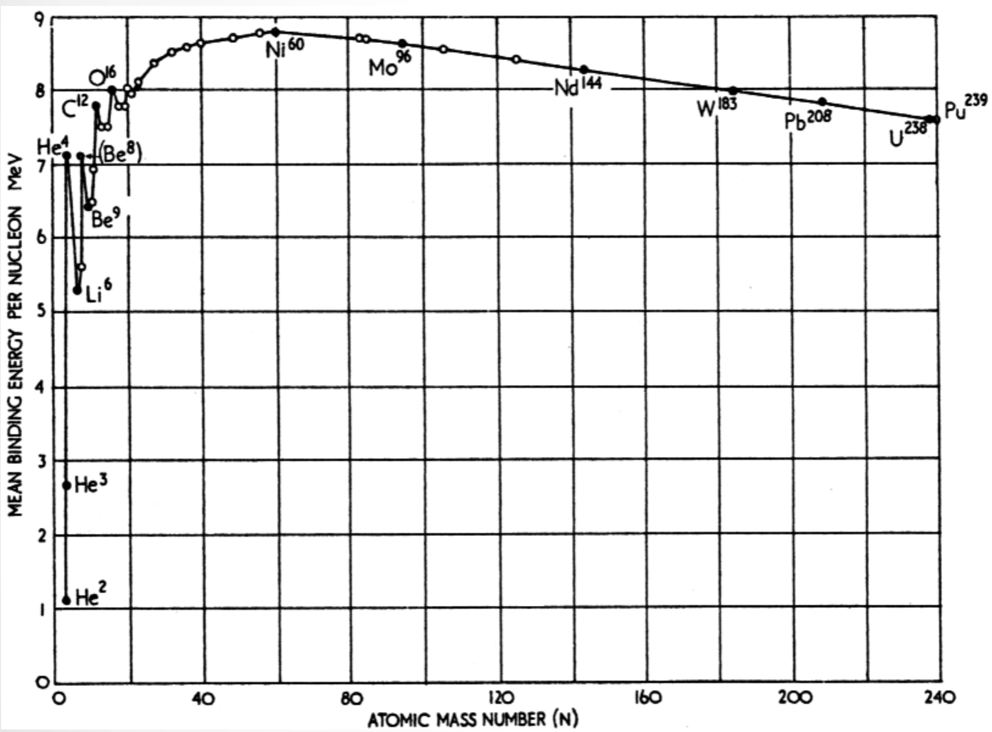
Principle fusion chain:

- $p^+ + n^0 \rightarrow H^2$ (deuterium)
  - No electrical repulsion
- $H^2 + H^2 \rightarrow H^3 + p^+$
- $H^3 + H^2 \rightarrow H^4 + n^0$

H^2 is relatively easily broken apart by energetic photons.

This is what starts the nucleosynthesis chain going: Universe cools to the point where very energetic photons are rare and H^2 is available

The Surprise



Natural next steps:

- $He^4 + He^4 \rightarrow Be^8$

Be^8 is unstable with a half life of 7×10^{-17} seconds!

- $He^4 + p^+ \rightarrow Li^5$

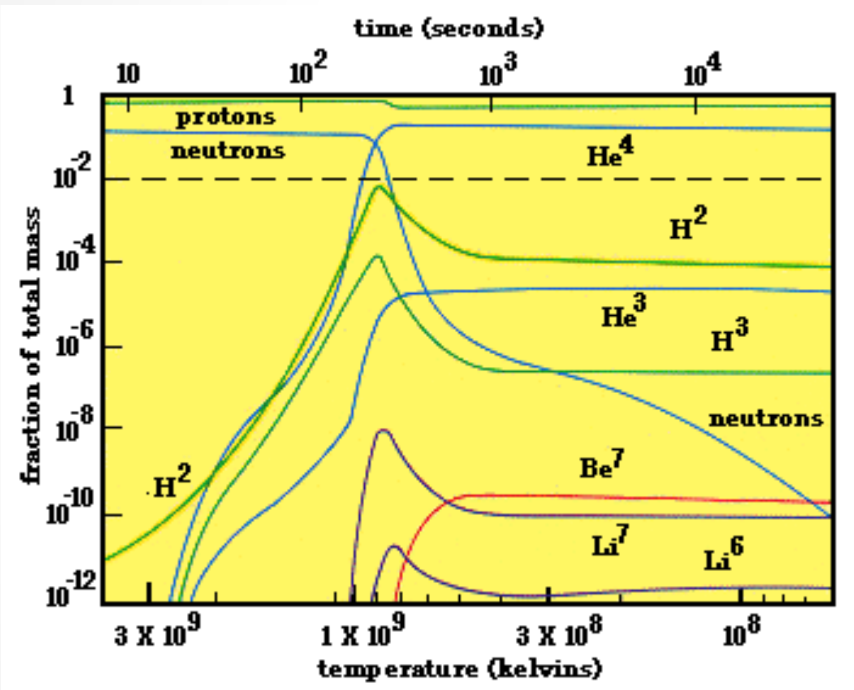
Unstable again

- Other isotopes of Be and Li are stable, but energetically not favored. Makes these in small amounts

The Surprise

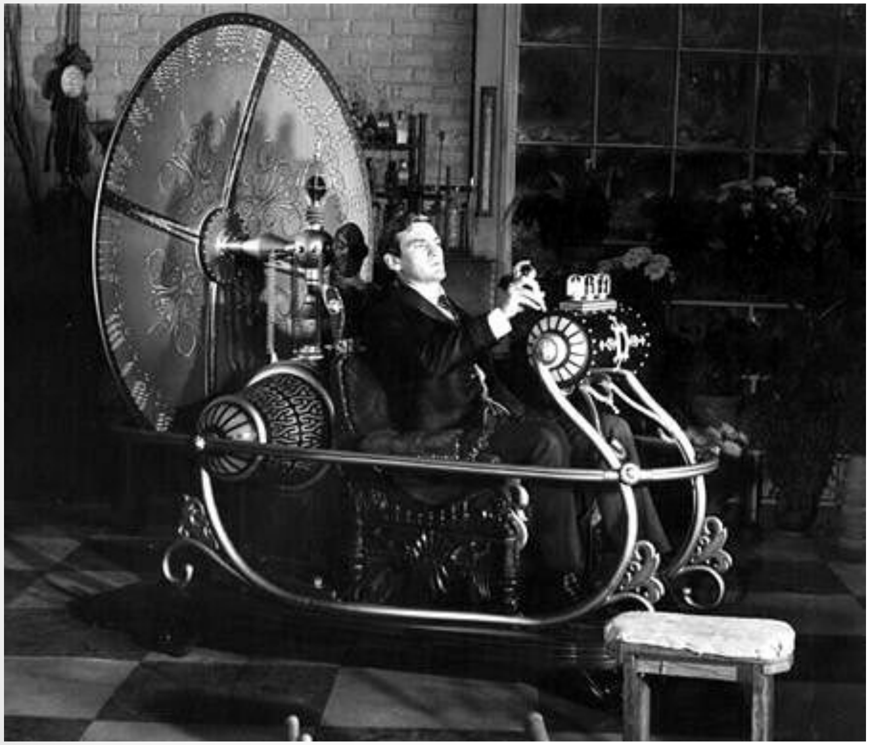
- By 10^3 seconds after the BB, the temperature of the Universe has dropped to values too low to support fusion reactions
- Produced lots of hydrogen, 25% by mass of He^4, some H^2, $H^{3,}$ He^3, Li^7, Li^6, Be^7
- This was used as an argument against the Hot Big Bang model: where did all the other elements come from?
- Required a slight asymmetry in matter vs anti-matter (1 extra matter particle per billion)

Big Bang Nucleosynthesis

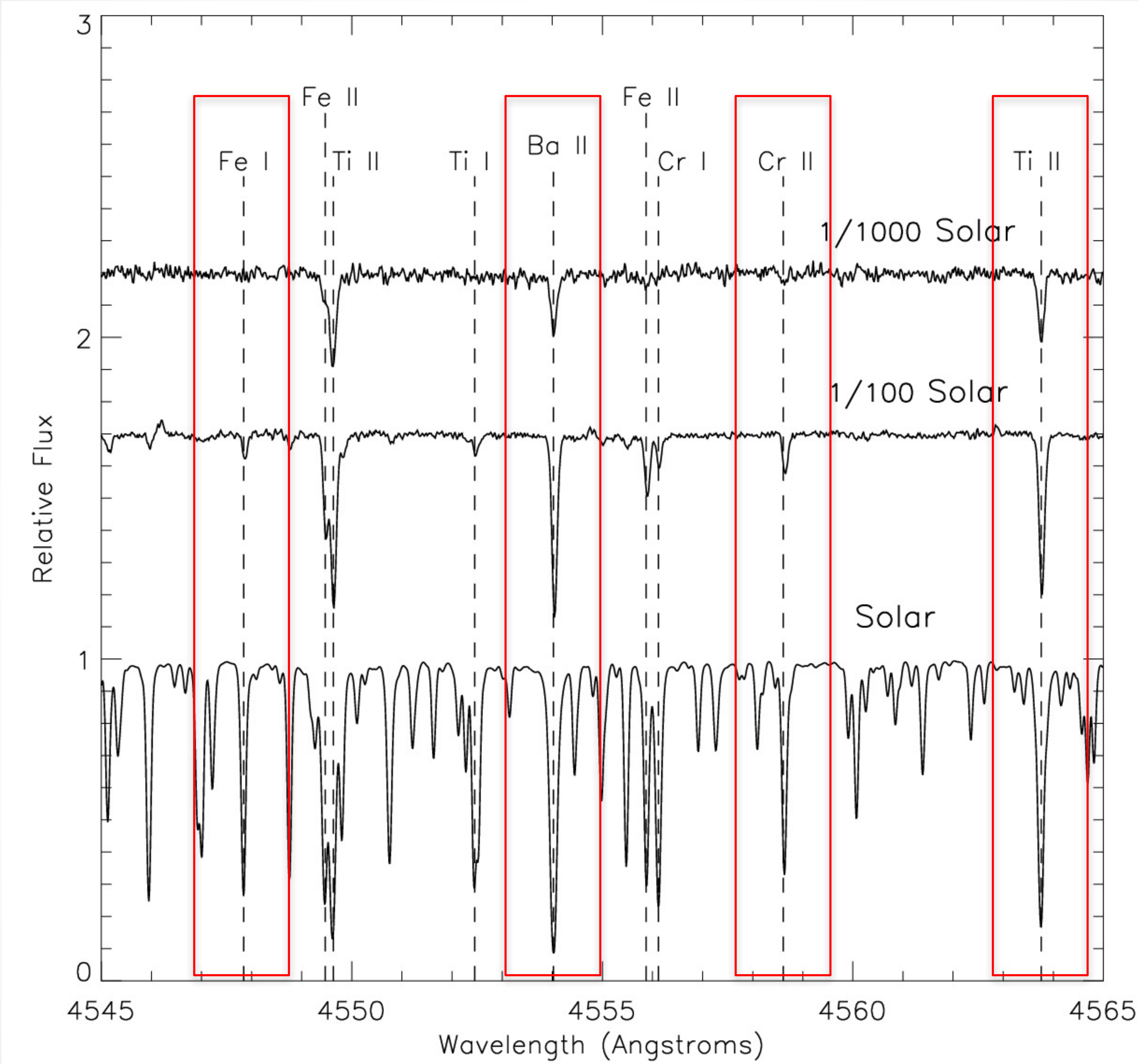


- BB+1 second: electrons, photons, neutrons, protons
- BB+2 minutes: some H^2 (p+n) produced
- BB+4 minutes: He production+tiny amount of Be, B and Li
- That's all! Universe has expanded to 10^9K and a density of only 10 g/cm^2

Fossils and time machines



- Stars like the Sun live for 10 billion years and some in the Galaxy are very old: these preserve the properties of the chemical composition of the Galaxy from the time of their formation
- Searching for "first stars" has been going on since 1970s



Fe II
Fe II
Fe I
Ti II
Ti I
Ba II
Cr I
Cr II
Ti II
1/1000 Solar
1/100 Solar
Solar
Relative Flux
Wavelength (Angstroms)
3
2
1
0
4545
4550
4555
4560
4565

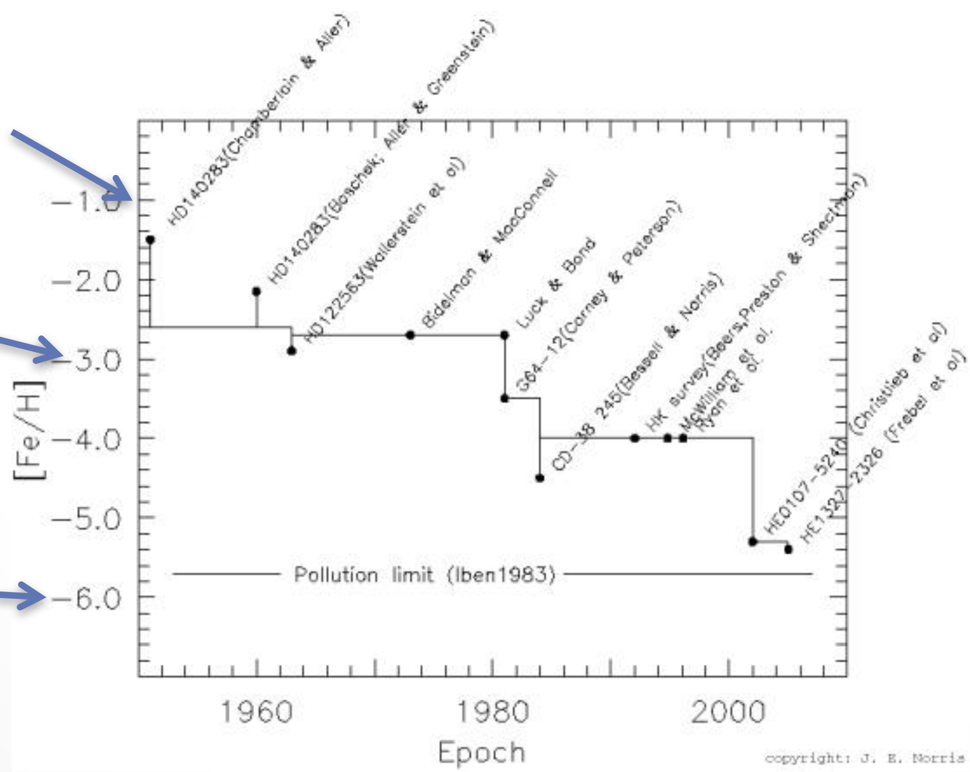
Big Bang Nucleosynthesis

- Is this story right? Makes some predictions:
  - The oldest stars in the Galaxy are deficient in the abundance of elements heavier than Helium

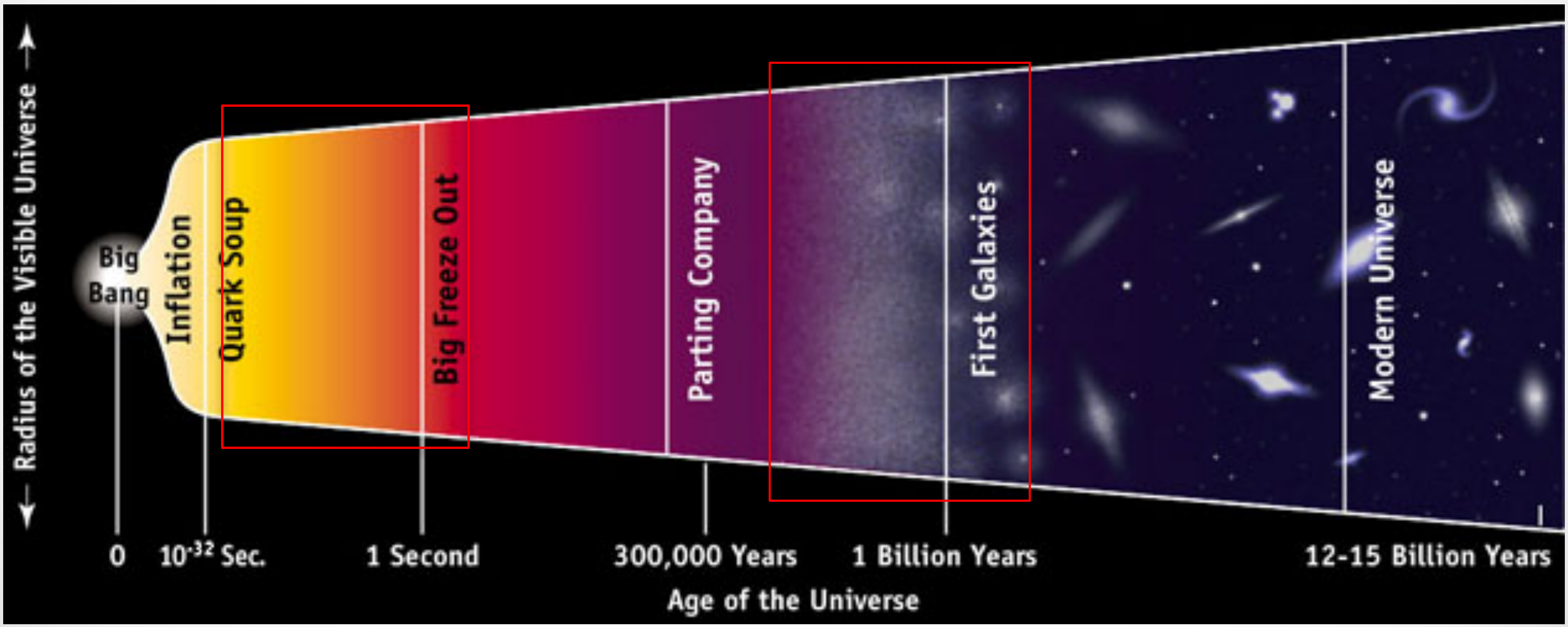
Factor 10 lower abundance of Fe compared to the Sun

Factor 10^3 lower abundance of Fe compared to the Sun

Factor 10^6 lower abundance of Fe compared to the Sun

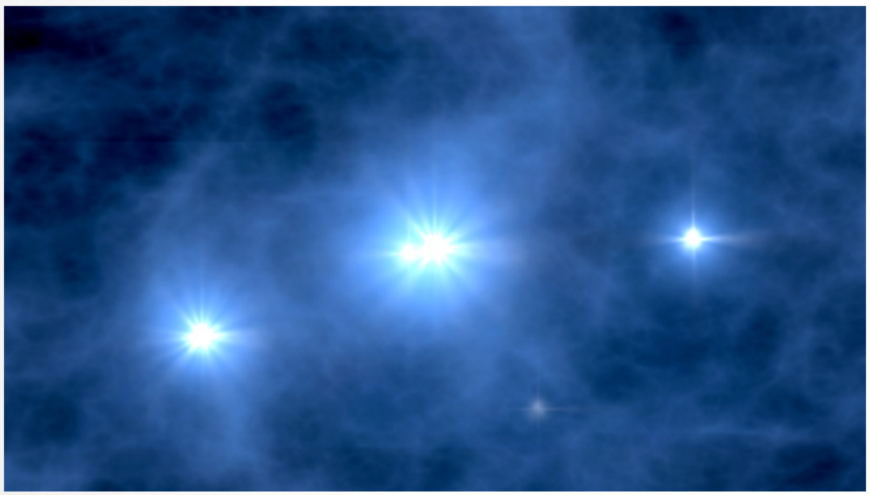


Universe Timeline



It was too hot for stars or galaxies to form for the first 200-400 million years, but at some point the First Stars formed from "primordial" material

The First Stars



- Have looked hard for stars composed of only HBB material
- Have gotten very close, but there may not be any true first-generation stars in the Galaxy
  - Galaxy formed from enriched gas
  - Stars made of only H and He all have masses $>0.9M_{Sun}$ and lifespans <12 Gyr

Pristine Gas in the Early Universe

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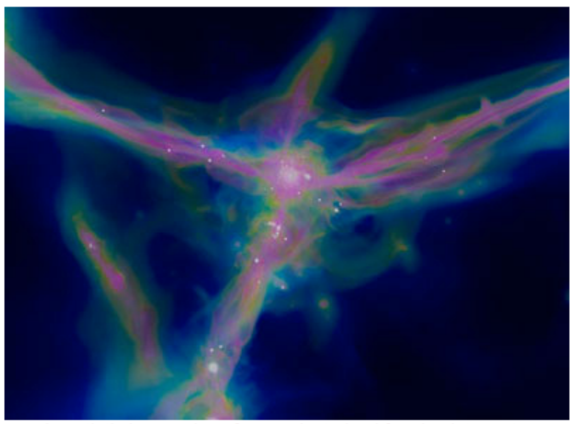
Pristine relics of the Big Bang spotted

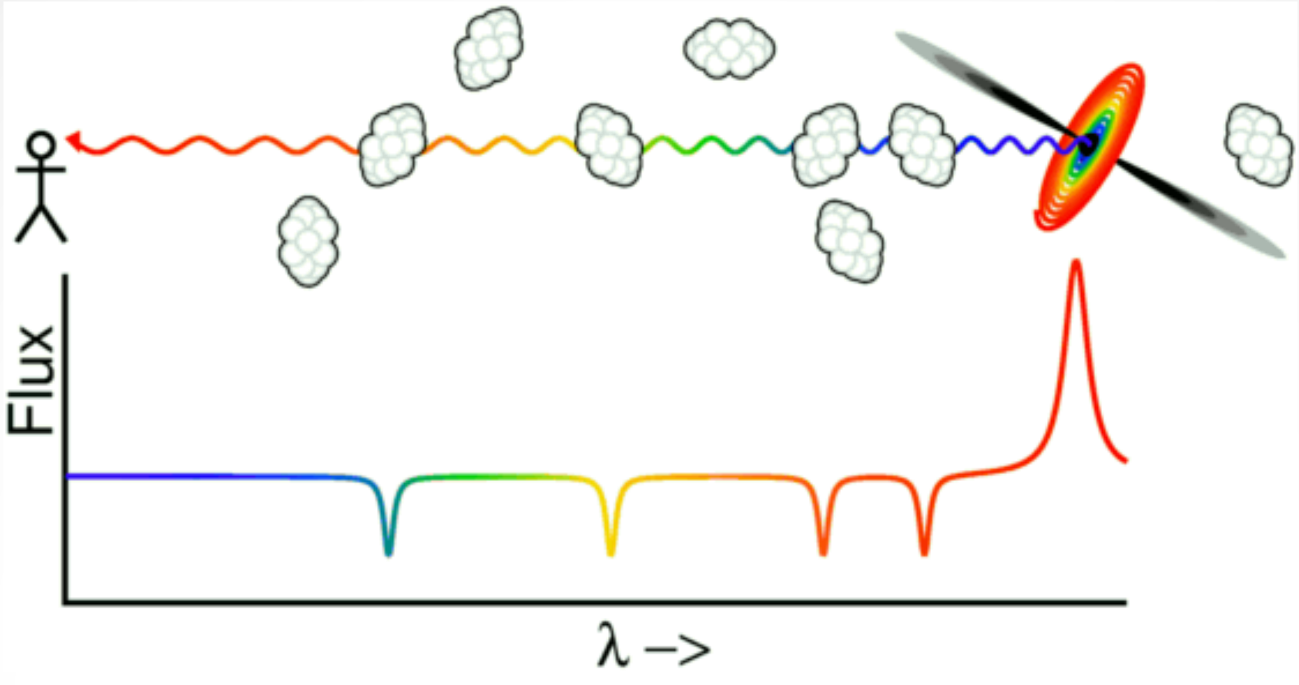
Nov 10, 2011 27 comments

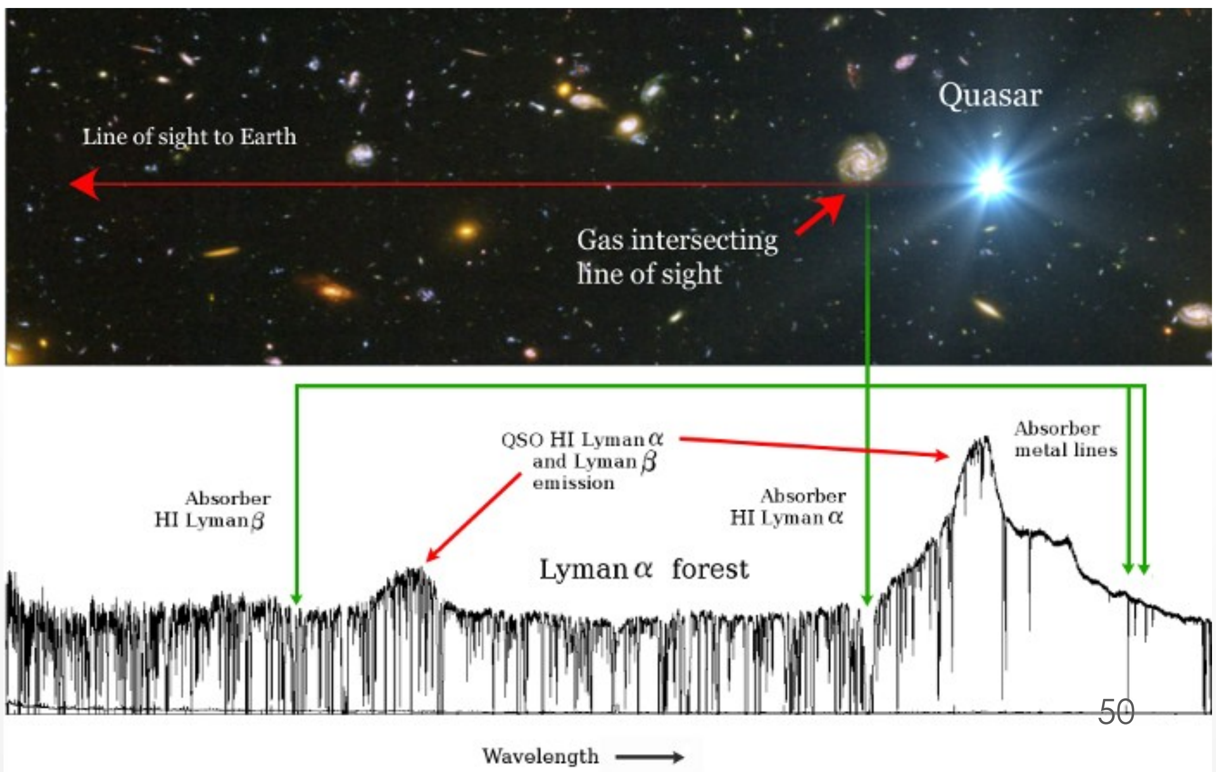


An artist's impression of an ancient cloud forming into a star

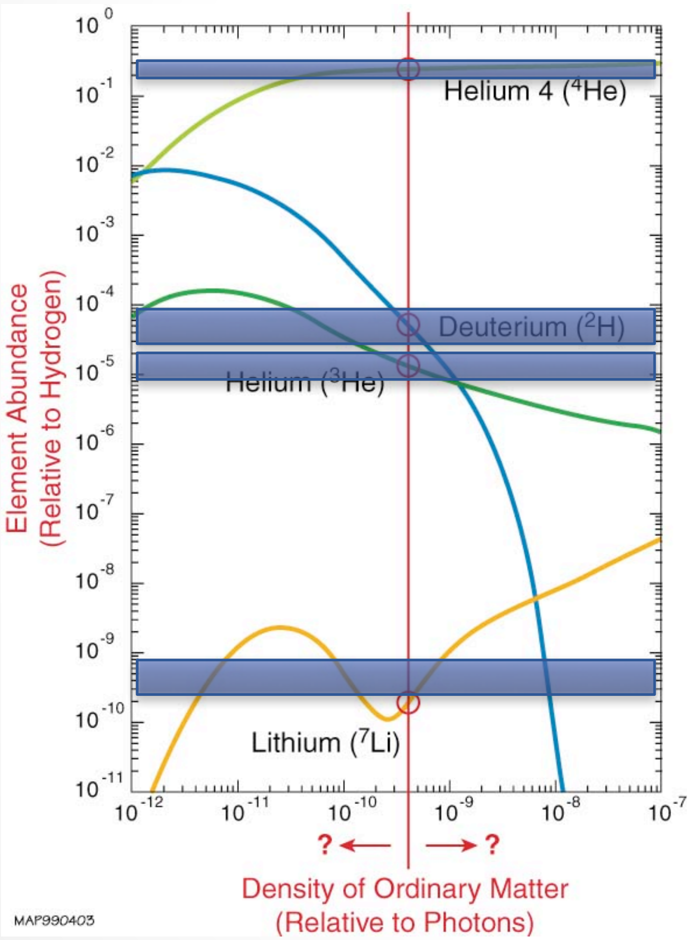
For the first time, astronomers have discovered two distant clouds of gas that seem to be pure relics from the Big Bang. Neither cloud contains any detectable elements forged by stars; instead, each consists only of the light elements that arose in the Big Bang some 14 billion years ago. Furthermore, the relatively high abundance of deuterium seen in one of the clouds agrees with predictions of Big Bang theory.

- In addition to looking for old stars made of primordial material, can look directly back in time at large look-back times
- 2011, UCSC astronomers made the first discovery of pristine clouds of primordial material


Flux
λ ->


Quasar
Line of sight to Earth
Gas intersecting
line of sight
QSO HI Lyman α
and Lyman β
emission
Absorber
metal lines
Absorber
HI Lyman β
Absorber
HI Lyman α
Lyman α forest
Wavelength

HBB Nucleosynthesis



- Calculations also make specific predictions for the "primordial" abundance of H, H^2, He^3, He^4 and Li^7
- Depends on the ratio of "baryons" to photons at the time of nucleosynthesis (plus a few other interesting things like the number of neutrino families)

HBB nucleosynthesis looks pretty good!

Run through the Standard Model physics for the Universe evolving from ~10^{-30} to 10^{3} seconds and we predict quite accurately the abundances of primordial elements and match particle accelerator results

Origin of the Elements

HBB stars SNI r- and s-process (SNII and AGB stars